

EDF R&D



FLUID DYNAMICS, POWER GENERATION AND ENVIRONMENT DEPARTMENT  
SINGLE PHASE THERMAL-HYDRAULICS GROUP

6, QUAI WATIER  
F-78401 CHATOU CEDEX

TEL: 33 1 30 87 75 40  
FAX: 33 1 30 87 79 16

DECEMBER 2013

*Code\_Saturne* documentation

***Code\_Saturne* version 3.2.0 practical user's guide**

contact: [saturne-support@edf.fr](mailto:saturne-support@edf.fr)



EDF R&D	<b><i>Code_Saturne</i> version 3.2.0 practical user's guide</b>	<i>Code_Saturne</i> documentation Page 1/ <a href="#">202</a>
---------	---	---

## ABSTRACT

*Code\_Saturne* is a system designed to solve the Navier-Stokes equations in the cases of 2D, 2D axisymmetric or 3D flows. Its main module is designed for the simulation of flows which may be steady or unsteady, laminar or turbulent, incompressible or potentially dilatable, isothermal or not. Scalars and turbulent fluctuations of scalars can be taken into account. The code includes specific modules, referred to as “specific physics”, for the treatment of lagrangian particle tracking, semi-transparent radiative transfer, gas combustion, pulverised coal combustion, electricity effects (Joule effect and electric arcs) and compressible flows. *Code\_Saturne* relies on a finite volume discretisation and allows the use of various mesh types which may be hybrid (containing several kinds of elements) and may have structural non-conformities (hanging nodes).

The present document is a practical user's guide for *Code\_Saturne* version 3.2.0. It is the result of the joint effort of all the members in the development team. It presents all the necessary elements to run a calculation with *Code\_Saturne* version 3.2.0. It then lists all the variables of the code which may be useful for more advanced utilisation. The user subroutines of all the modules within the code are then documented. Eventually, for each key word and user-modifiable parameter in the code, their definition, allowed values, default values and conditions for use are given. These key words and parameters are grouped under headings based on their function. An alphabetical index list is also given at the end of the document for easier consultation.

*Code\_Saturne* is free software; you can redistribute it and/or modify it under the terms of the GNU General Public License as published by the Free Software Foundation; either version 2 of the License, or (at your option) any later version. *Code\_Saturne* is distributed in the hope that it will be useful, but WITHOUT ANY WARRANTY; without even the implied warranty of MERCHANTABILITY or FITNESS FOR A PARTICULAR PURPOSE. See the GNU General Public License for more details.



EDF R&D	<i>Code_Saturne</i> version 3.2.0 practical user's guide	<i>Code_Saturne</i> documentation Page 2/ <a href="#">202</a>
---------	---	---

## TABLE OF CONTENTS

<b>1</b>	<b>Introduction</b>	9
<b>2</b>	<b>Quick start</b>	10
2.1	RUNNING A CALCULATION	10
2.2	TROUBLESHOOTING	10
<b>3</b>	<b>Practical information about Code_Saturne</b>	12
3.1	SYSTEM ENVIRONMENT FOR Code_Saturne	12
3.1.1	<i>Preliminary settings</i>	12
3.1.2	<i>Configuration file</i>	12
3.1.3	<i>Standard directory hierarchy</i>	12
3.1.4	<i>Code_Saturne Kernel library files</i>	14
3.2	SETTING UP AND RUNNING A CALCULATION	14
3.2.1	<i>Step by step calculation</i>	14
3.2.2	<i>Temporary execution directory</i>	17
3.2.3	<i>Execution modes</i>	17
3.2.4	<i>Environment variables</i>	18
3.2.5	<i>Interactive modification of the target time step</i>	18
3.3	CASE PREPARER	19
3.4	SUPPORTED MESH AND POST-PROCESSING OUTPUT FORMATS	19
3.4.1	<i>Formats supported for input</i>	20
3.4.2	<i>Formats supported for input or output</i>	22
3.4.3	<i>Formats supported for output only</i>	26
3.4.4	<i>Meshing tools and associated formats</i>	26
3.4.5	<i>Meshing remarks</i>	26
3.5	PREPROCESSOR COMMAND LINE OPTIONS	27
3.6	KERNEL COMMAND LINE OPTIONS	27
3.7	LAUNCH SCRIPTS	28
3.8	GRAPHICAL USER INTERFACE	29
3.9	USER SUBROUTINES	30
3.9.1	<i>Preliminary comments</i>	30
3.9.2	<i>Example routines</i>	30
3.9.3	<i>Main variables</i>	31
3.9.4	<i>Using selection criteria in user subroutines</i>	50
3.10	FACE AND CELL MESH-DEFINED PROPERTIES AND SELECTION	51
<b>4</b>	<b>Importing and Preprocessing Meshes</b>	53
4.1	PREPROCESSOR OPTIONS	54

4.1.1	<i>Mesh selection</i>	54
4.1.2	<i>Post-processing output</i>	54
4.1.3	<i>Element orientation correction</i>	54
4.2	ENVIRONMENT VARIABLES	54
4.2.1	<i>System environment variables</i>	55
4.3	OPTIONAL FUNCTIONALITY	55
4.4	GENERAL REMARKS	55
4.5	FILES PASSED TO THE KERNEL	56
4.6	MESH PREPROCESSING	56
4.6.1	<i>Joining of non-conforming meshes</i>	56
4.6.2	<i>Periodicity</i>	57
4.6.3	<i>Parameters for conforming or non-conforming mesh joinings</i>	57
4.6.4	<i>Parameters for periodicity</i>	60
4.6.5	<i>Modification of the mesh geometry</i>	60
4.7	MESH SMOOTHING UTILITIES	61
4.7.1	<i>Fix by feature</i>	61
4.7.2	<i>Warped faces smoother</i>	61
5	<b>Partitioning for parallel runs</b>	61
5.1	PARTITIONING STAGES	62
5.2	PARTITIONER CHOICE	62
5.3	EFFECT OF PERIODICITY	62
6	<b>Basic modelling setup</b>	63
6.1	INITIALISATION OF THE MAIN PARAMETERS	63
6.2	SELECTION OF MESH INPUTS: <code>CS_USER_MESH_INPUT</code>	66
6.3	NON-DEFAULT VARIABLES INITIALISATION	68
6.4	MANAGE BOUNDARY CONDITIONS	71
6.4.1	<i>Coding of standard boundary conditions</i>	72
6.4.2	<i>Coding of non-standard boundary conditions</i>	74
6.4.3	<i>Checking of the boundary conditions</i>	76
6.4.4	<i>Sorting of the boundary faces</i>	77
6.4.5	<i>Boundary conditions with LES</i>	77
6.5	MANAGE THE VARIABLE PHYSICAL PROPERTIES	82
6.5.1	<i>Basic variable physical properties</i>	82
6.5.2	<i>Modification of the turbulent viscosity</i>	83
6.5.3	<i>Modification of the variable <math>C</math> of the dynamic LES model</i>	83
6.6	USER SOURCE TERMS	84
6.6.1	<i>In Navier-Stokes</i>	86

6.6.2	<i>For <math>k</math> and <math>\varepsilon</math></i>	86
6.6.3	<i>For <math>R_{ij}</math> and <math>\varepsilon</math></i>	86
6.6.4	<i>For <math>\varphi</math> and <math>\bar{f}</math></i>	86
6.6.5	<i>For <math>k</math> and <math>\omega</math></i>	87
6.6.6	<i>For <math>\tilde{\nu}_t</math></i>	87
6.6.7	<i>For user scalars</i>	87
6.7	PRESSURE DROPS (HEAD LOSSES) AND POROSITY	87
6.7.1	<i>Head losses</i>	87
6.7.2	<i>Porosity</i>	89
6.8	MANAGEMENT OF THE MASS SOURCES	89
6.9	USER LAW EDITOR OF THE GUI	90
<b>7</b>	<b>Results analysis</b>	92
7.1	DEFINITION OF POST-PROCESSING AND MESH ZONES	92
7.1.1	<i>Management of the post-processing intermediate outputs</i>	94
7.2	DEFINITION OF THE VARIABLES TO POST-PROCESS	95
7.3	MODIFICATION OF THE VARIABLES AT THE END OF A TIME STEP	96
7.4	NON-STANDARD MANAGEMENT OF THE CHRONOLOGICAL RECORD FILES	96
<b>8</b>	<b>Advanced modelling setup</b>	97
8.1	USE OF A SPECIFIC PHYSICS	97
8.2	PULVERISED COAL AND GAS COMBUSTION MODULE	102
8.2.1	<i>Boundary conditions</i>	104
8.2.2	<i>Initialisation of the options of the variables</i>	107
8.3	HEAVY FUEL OIL COMBUSTION MODULE	110
8.3.1	<i>Initialisation of transported variables</i>	110
8.3.2	<i>Boundary conditions</i>	110
8.4	RADIATIVE THERMAL TRANSFERS IN SEMI-TRANSPARENT GRAY MEDIA	110
8.4.1	<i>Initialisation of the radiation main parameters</i>	110
8.4.2	<i>Radiative transfers boundary conditions</i>	111
8.4.3	<i>Absorption coefficient of the medium, boundary conditions for the luminance and calculation of the net radiative flux</i>	114
8.4.4	<i>Encapsulation of the temperature-enthalpy conversion</i>	115
8.4.5	<i>Input of radiative transfer parameters</i>	115
8.5	CONJUGATE HEAT TRANSFER	116
8.5.1	<i>Thermal module in a 1D wall</i>	116
8.5.2	<i>Fluid-Thermal coupling with SYRTHES</i>	117
8.6	PARTICLE-TRACKING (LAGRANGIAN) MODULE	118
8.6.1	<i>General information</i>	118

8.6.2	<i>Activating the particle-tracking module</i>	118
8.6.3	<i>Basic guidelines for standard simulations</i>	118
8.6.4	<i>Prescribing the main modelling parameters (GUI and/or <b>uslag1</b>)</i>	119
8.6.5	<i>Prescribing particle boundary conditions (GUI and/or <b>uslag2</b>)</i>	120
8.6.6	<i>Advanced particle-tracking set-up</i>	123
8.7	COMPRESSIBLE MODULE	125
8.7.1	<i>Initialisation of the options of the variables</i>	126
8.7.2	<i>Management of the boundary conditions</i>	126
8.7.3	<i>Initialisation of the variables</i>	127
8.7.4	<i>Management of variable physical properties</i>	127
8.8	MANAGEMENT OF THE ELECTRIC ARCS MODULE	127
8.8.1	<i>Activating the electric arcs module</i>	127
8.8.2	<i>Initialisation of the variables</i>	127
8.8.3	<i>Variable physical properties</i>	128
8.8.4	<i>Boundary Conditions</i>	128
8.8.5	<i>Initialisation of the variable options</i>	129
8.8.6	<i>EnSight output</i>	130
8.9	Code_Saturne-Code_Saturne COUPLING	130
8.10	FLUID-STRUCTURE EXTERNAL COUPLING	130
8.11	ALE MODULE	131
8.11.1	<i>Initialisation of the options</i>	131
8.11.2	<i>Boundary conditions of mesh velocity</i>	132
8.11.3	<i>Modification of the viscosity</i>	133
8.11.4	<i>Fluid - Structure internal coupling</i>	133
8.12	MANAGEMENT OF THE STRUCTURE PROPERTY	134
8.13	MANAGEMENT OF THE ATMOSPHERIC MODULE	135
8.13.1	<i>Initialisation of the variables</i>	135
8.13.2	<i>Management of the boundary conditions</i>	135
9	<b>Key word list</b>	138
9.1	INPUT-OUTPUT	138
9.1.1	<i>"Calculation" files</i>	139
9.1.2	<i>Post-processing for EnSight or other tools</i>	140
9.1.3	<i>Chronological records of the variables on specific points</i>	141
9.1.4	<i>Time averages</i>	142
9.1.5	<i>Others</i>	144
9.2	NUMERICAL OPTIONS	145
9.2.1	<i>Calculation management</i>	145

9.2.2	<i>Scalar unknowns</i>	146
9.2.3	<i>Definition of the equations</i>	148
9.2.4	<i>Definition of the time advancement</i>	149
9.2.5	<i>Turbulence</i>	151
9.2.6	<i>Time scheme</i>	156
9.2.7	<i>Gradient reconstruction</i>	161
9.2.8	<i>Solution of the linear systems</i>	162
9.2.9	<i>Convective scheme</i>	164
9.2.10	<i>Pressure-continuity step</i>	164
9.2.11	<i>Error estimators for Navier-Stokes</i>	166
9.2.12	<i>Calculation of the distance to the wall</i>	167
9.2.13	<i>Others</i>	170
9.3	NUMERICAL, PHYSICAL AND MODELLING PARAMETERS	171
9.3.1	<i>Numeric Parameters</i>	171
9.3.2	<i>Physical parameters</i>	171
9.3.3	<i>Physical variables</i>	172
9.3.4	<i>modelling parameters</i>	176
9.4	ALE	180
9.5	THERMAL RADIATIVE TRANSFERS: GLOBAL SETTINGS	180
9.6	ELECTRIC MODULE (JOULE EFFECT AND ELECTRIC ARCS): SPECIFICITIES	182
9.7	COMPRESSIBLE MODULE: SPECIFICITIES	184
9.8	LAGRANGIAN MULTIPHASE FLOWS	185
9.8.1	<i>Global settings</i>	185
9.8.2	<i>Specific physics models associated with the particles</i>	187
9.8.3	<i>Options for two-way coupling</i>	188
9.8.4	<i>Numerical modelling</i>	188
9.8.5	<i>Volume statistics</i>	189
9.8.6	<i>Display of particles and trajectories</i>	190
9.8.7	<i>Display of the particle/boundary interactions and the statistics at the boundaries</i>	191
10	Bibliography	194
	Index of the main variables and keywords	196

EDF R&D	<i>Code_Saturne</i> version 3.2.0 practical user's guide	<i>Code_Saturne</i> documentation Page 8/ <a href="#">202</a>
---------	---	---

# 1 Introduction

*Code\_Saturne* is an application designed to solve the Navier-Stokes equations in the cases of 2D, 2D axi-symmetric or 3D flows. Its main module is designed for the simulation of flows which may be steady or unsteady, laminar or turbulent, incompressible or potentially dilatant, isothermal or not. Scalars and turbulent fluctuations of scalars can be taken into account. The code includes specific modules, referred to as “specific physics”, for the treatment of Lagrangian particle tracking, semi-transparent radiative transfer, gas combustion, pulverised coal combustion, electricity effects (Joule effect and electric arcs) and compressible flows.

*Code\_Saturne* is free software; you can redistribute it and/or modify it under the terms of the GNU General Public License as published by the Free Software Foundation; either version 2 of the License, or (at your option) any later version. *Code\_Saturne* is distributed in the hope that it will be useful, but WITHOUT ANY WARRANTY; without even the implied warranty of MERCHANTABILITY or FITNESS FOR A PARTICULAR PURPOSE. See the GNU General Public License for more details.<sup>1</sup>

*Code\_Saturne* relies on a finite volume discretisation and allows the use of various mesh types which may be hybrid (containing several kinds of elements) and may have structural non-conformities (hanging nodes).

*Code\_Saturne* is composed of two main elements and an optional GUI, as shown on Figure 1:

- the Kernel module is the numerical solver
- the Preprocessor module is in charge of mesh import

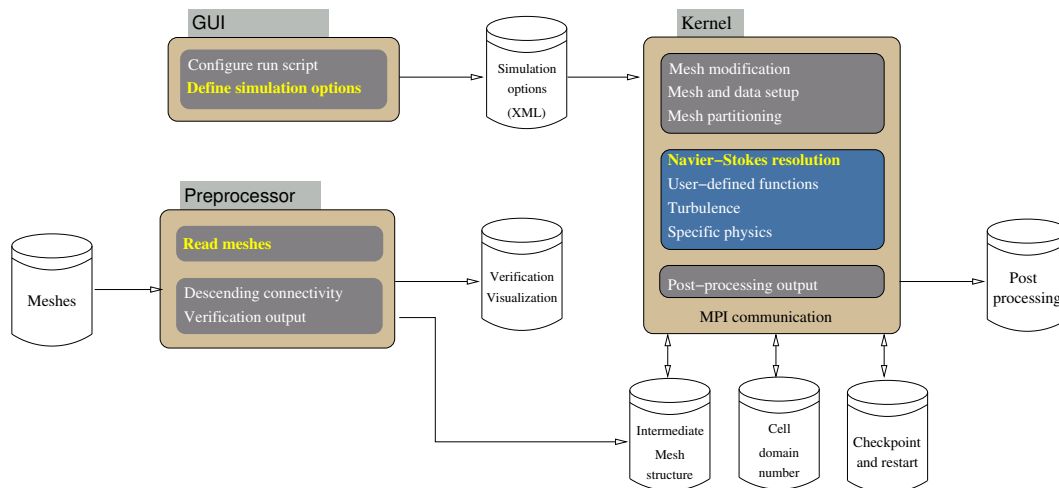


Figure 1: *Code\_Saturne* elements

*Code\_Saturne* also relies on the PLE (Parallel Location and Exchange) library (by the same team, under LGPL licence) for the management of code coupling; this library can also be used independently.

This document is a practical user guide for *Code\_Saturne* version 3.2.0. It is the result of the joint effort of all the members in the development team.

This document gives practical information for the usage of *Code\_Saturne*. For more details about the algorithms and their numerical implementation, please refer to the reports [1], [4] and [10], and to the theoretical documentation [11].

<sup>1</sup>You should have received a copy of the GNU General Public License along with *Code\_Saturne*; if not, write to the Free Software Foundation, Inc., 51 Franklin St, Fifth Floor, Boston, MA 02110-1301 USA



The latest updated version of this document is available on-line with the version of *Code\_Saturne* and accessible through the command `code_saturne info --guide theory`.

This document first presents all the necessary elements to run a calculation with *Code\_Saturne* version 3.2.0. It then lists all the variables of the code which may be useful for more advanced utilisation. The user subroutines of all the modules within the code are then documented. Eventually, for each keyword and user-modifiable parameter in the code, their definition, allowed values, default values and conditions for use are given. These keywords and parameters are grouped under headings based on their function. An alphabetical index is also given at the end of the document for easier consultation.

## 2 Quick start

### 2.1 Running a calculation

We assume in this section that the user has at his disposal the calculation data file (calculation set up) or already prepared it following for instance the step-by-step guidance provided in *Code\_Saturne* tutorial. The steps described below are intended to provide the user a way to run quickly on a workstation a calculation through the Graphical User Interface (GUI).

The first thing to do before running *Code\_Saturne* is to define an alias to the `code_saturne` script (see §3.1.1), for example:

```
alias cs='${prefix}/bin/code_saturne'.
```

When using the *bash* shell, a completion file may be sourced so as to allow for syntax auto-completion:

```
source ${prefix}/etc/bash_completion.d/code_saturne'.
```

The second thing is to prepare the computation directories. For instance, the study directory `T_JUNCTION`, containing a single calculation directory `CASE1`, will be created by typing the command (see §3.3):

```
code_saturne create -s T_JUNCTION
```

The mesh files should be copied in the directory `MESH` and the Fortran user files necessary for the calculation in the directory `CASE1/SRC`. Finally, the calculation data file `case_name.xml` read by the GUI should be copied to the directory `CASE1/DATA`. Once these steps completed, the user should go in the directory `CASE1/DATA` and type de command line `./SaturneGUI case_name.xml` to load the calculation file into the interface. A window similar to Figure2 will appear. Click on the heading “Calculation management”, select the heading “Prepare batch calculation”, see Figure 3. After having chosen the number of processors, press “start calculation” to run the calculation.

If no problem arises, the simulation results can be found in the directory `CASE1/RESU` and be read directly by *ParaView* or *EnSight* in `CASE1/RESU/<YYYYMMDD-hhmm>/postprocessing`. Calculation history can be found in the file `<YYYYMMDD-hhmm>/listing`.

### 2.2 Troubleshooting

If the calculation does not run properly, the user is advised to check the following points in `CASE1/RESU/<YYYYMMDD-hhmm>`:

- if the calculation stops in the pre-processor, the user should check for error messages in the file `preprocessor*.log`.
- if the problem is related to boundary conditions, the user should visualise the file `error.ensight` with *EnSight* or *ParaView*,



Figure 2: Identity and paths



Figure 3: Prepare execution

- if the calculation stops in the *Code\_Saturne* core, the user should look for messages at the end of the files **listing** and **error\***. In addition, the user can track the following keywords in the listing. They are specific error signals:
  - **SIGFPE**: a floating point exception occurred. It happens when there is a division by 0, when the calculation did not converge, or when a real number reached a value over  $10^{300}$ . Depending on the architecture *Code\_Saturne* is running on, this type of exception may be caught or ignored.
  - **SIGSEGV**: a memory error such as a segmentation violation occurred. An array may have exceeded its allocated memory size and a memory location in use was overwritten.

In order to easily find the problem, it is also advised to use a debug version of *Code\_Saturne* (see the installation documentation) in combination with the use of the valgrind tool (if it is installed). The use of valgrind can be specified in the GUI in the advanced options of the item “Prepare batch calculation” under the heading “Calculation management” or without the GUI, in the `cs_user_scripts.py` file (this file can be found in DATA/REFERENCE and should be copied in DATA, see §3.2.1).

## 3 Practical information about *Code\_Saturne*

### 3.1 System Environment for *Code\_Saturne*

#### 3.1.1 Preliminary settings

In order to use *Code\_Saturne*, every user should define the following alias (in their `.bashrc`, or equivalent, or `.alias` file, depending on the environment):

```
alias cs='${install_directory}/bin/code_saturne'
```

where `install_directory` is the base directory where *Code\_Saturne* and its components have been installed<sup>2</sup>.

This step may be skipped if `${install_directory}` is in a standard location (such as `/usr` or `/usr/local`).

#### 3.1.2 Configuration file

A configuration file for *Code\_Saturne* is available in `${install_directory}/etc`. This file can be useful as a post-install step for computing environments using a batch system, for separate front-end and compute systems (such as Blue Gene systems), or for coupling with SYRTHES 4 or *Code\_Aster* (see the installation documentation for more details).

To use this file, copy or rename the `${install_directory}/etc/code_saturne.cfg.template` to `${install_directory}/etc/code_saturne.cfg`, and uncomment and define the applicable sections. Two other options in the `code_saturne.cfg` file could be useful for the user:

- Set the temporary directory (see §3.2.2 for more details on the temporary execution directory).
- Set the mesh database directory: it is possible to indicate a path where meshes are stored. In this case, the GUI will propose this directory automatically for mesh selection. Without the GUI, it is then possible to fill in the `cs_user_scripts.py` file (see §3.2.1) with the name of the desired mesh of the database directory and the code will find it automatically (be careful if you have the same name for a mesh in the database directory and in the `MESH` directory, the mesh in `MESH` will be used).

#### 3.1.3 Standard directory hierarchy

The standard architecture for the simulation studies is:

An optional study directory containing:

- A directory `MESH` containing the mesh(es) necessary for the study
- A directory `POST` for the potential post-processing scripts (not used directly by the code)
- One or several calculation directories

Every calculation directory contains:

- A directory `SRC` for the potential user subroutines necessary for the calculation
- A directory `DATA` for the calculation data (data file from the interface, input profiles, thermo-chemical data, ...), the user script and the XML file.

---

<sup>2</sup>Without this step, using the absolute path is still possible

EDF R&D	<b><i>Code_Saturne</i> version 3.2.0 practical user's guide</b>	<i>Code_Saturne</i> documentation Page 13/202
---------	---	---

- A directory **SCRIPTS** for the launch script
- A directory **RESU** for the results  
To improve the calculation traceability, the files and directories sent to **RESU** after a calculation are placed in a subdirectory named after that run's "id", which is by default based on the run date and time, using the format: **YYYYMMDD-hhmm**. It is also possible to force a specific run id, using the **--id** option of **code\_saturne run**.

In the standard cases, **RESU/<run\_id>** contains a **postprocessing** directory with the post-processing (visualization) files, a **restart** directory for the calculation restart files, a **monitoring** directory for the files of chronological record of the results at specific locations (probes), **preprocessor.log** and **listing** files reporting the Preprocessor and the Kernel execution. All files from the **DATA** directory not in subdirectories are also copied. For a tracing of the modifications in prior calculations, the user-subroutines used in a calculation are stored in a **src\_saturne** subdirectory. The data files (such as the XML Interface data file and thermo-chemical data files) and launch script are also copied into the results directory. **compile.log** and **summary** are respectively reports of the compilation stage and general information on the calculation (type of machine, user, version of the code, ...).

Below are typical contents of a case directory **CASE1** in a study **STUDY**

<b>STUDY/CASE1/DATA:</b>	<b><i>Code_Saturne</i> data</b>
SaturneGUI	Graphical User Interface launch script
study.xml	Graphical User Interface parameter file
REFERENCE	Example of user scripts and meteorological or thermochemical data files (used with the specific physics modules)
<b>STUDY/CASE1/SRC:</b>	<b><i>Code_Saturne</i> user subroutines</b>
REFERENCE	Available user subroutines
EXAMPLES	Examples of user subroutines
cs_user_boundary_conditions.f90	user subroutines used for the present calculation
cs_user_parameters.f90	
<b>STUDY/CASE1/RESU/YYYYMMDD-hhmm:</b>	<b>Results</b> for the calculation <b>YYYYMMDD-hhmm</b>
postprocessing	Directory containing the <i>Code_Saturne</i> post-processing output in the <i>EnSight</i> format (both volume and boundary); copy of the <i>Code_Saturne</i> user subroutines used for the calculation
src_saturne	Directory containing the chronological records for <i>Code_Saturne</i>
monitoring	Directory containing the <i>Code_Saturne</i> restart files
checkpoint	Compilation log
compile.log	Graphical User Interface parameter file used for the calculation
study.xml	Copy of the launch script used for the calculation
runcase	Execution report for the <i>Code_Saturne</i> Preprocessor
preprocessor.log	Execution report for the Kernel module of <i>Code_Saturne</i>
listing	General information (machine, user, version, ...)
summary	
<b>STUDY/CASE1/SCRIPTS:</b>	<b>Launch script</b>
runcase	Launch script (which may contain batch system keywords)

When running, the code may use additional files or directories inside its execution directory, set by the execution script, which include a **mesh\_input** file or directory, as well as a **restart** directory (which is a link or copy of a previous run's **checkpoint** directory), as well as a **run\_solver.sh** script.

For coupled calculations, whether with *Code\_Saturne* itself or SYRTHES, each coupled calculation domain is defined by its own directory (bearing the same name as the domain), but results are placed in a **RESU\_COUPLING** directory, with a subdirectory for each run, itself containing one subdirectory per coupled domain. Coupled cases are not run through the standard **STUDY/CASE1/SCRIPTS/runcase** script or through the **code\_saturne run** command, but through a **STUDY/runcase\_coupling** script.

So in the coupled case, calculation results would not be placed in `STUDY/CASE1/RESU/YYYYMMDD-hhmm`, but in `STUDY/RESU_COUPLING/YYYYMMDD-hhmm/CASE1`, with the `summary` file being directly placed in `STUDY/RESU_COUPLING/YYYYMMDD-hhmm` (as it references all coupled domains).

Below are typical additional contents with a coupled SYRTHES case SOLID1 in a study STUDY	
<code>STUDY/runcase_coupling</code>	Coupled launch script
<code>STUDY/SOLID1/ DATA :</code>	<b>SYRTHES data</b>
<code>syrthes.data.syd</code>	SYRTHES data file
<code>syrthes.py</code>	SYRTHES script
<code>usr_examples</code>	SYRTHES user subroutine examples
<code>STUDY/RESU_COUPLING/YYYYMMDD-hhmm/SOLID1 :</code>	<b>results (file names defined in <code>syrthes.env</code>)</b>
<code>src</code>	SYRTHES user subroutines used in the calculation
<code>compile.log</code>	SYRTHES compilation report
<code>listsyr</code>	Execution log
<code>geoms</code>	SYRTHES solid geometry file
<code>histos1</code>	SYRTHES chronological records at specified monitoring points
<code>resus1</code>	SYRTHES calculation restart file (1 time step)
<code>resusc1</code>	SYRTHES chronological solid post-processing file (may be transformed into the <i>EnSight</i> or <i>MED</i> format with the <i>syrthes4ensight</i> or <i>syrthes4med30</i> utility)

### 3.1.4 Code\_Saturne Kernel library files

Information about the content of the *Code\_Saturne* base directories is given below. It is not of vital interest for the user, but given only as general information. Indeed, the case preparer command `code_saturne create` automatically extracts the necessary files and prepares the launch script without the user having to go directly into the *Code\_Saturne* base directories (see §3.3). The `code_saturne info` command gives direct access to the most needed information (especially the user and programmer's guides and the tutorial) without the user having to look for them in the *Code\_Saturne* directories.

The subdirectories `{install_directory}/lib` and `{install_directory}/bin` contain the libraries and compiled executables respectively.

The data files (for instance thermochemical data) are located in the directory `data`.

The user subroutines are available in the directory `users`, under subdirectories corresponding to each module: `base` (general routines), `cfbl` (compressible flows), `cogz` (gas combustion), `cplv` (pulverised coal combustion), `ctwr` (cooling towers modelling), `elec` (electric module), `fuel` (heavy fuel oil combustion module), `lagr` (Lagrangian module), `pprt` (general specific physics routines) and `rayt` (radiative heat transfer). The case preparer command `code_saturne create` copies all these files in the user directory `SRC/REFERENCE` during the case preparation.

The directory `bin` contains an example of the launch script, the compilation parameter files and various utility programs.

## 3.2 Setting up and running a calculation

### 3.2.1 Step by step calculation

This paragraph summarises the different steps which are necessary to prepare and run a standard case:

- Check the version of *Code\_Saturne* set for use in the environment variables (`code_saturne info --version`). If it does not correspond to the desired version, update the user profile or aliases to get the required version, logging out of the session and in again if necessary (cf. §3.1.1).

- Prepare the different directories using the `code_saturne create` command (see §3.3).
- It is recommended to place the mesh(es) in the directory `MESH`, but they may be selected from other directories, either with the Graphical User Interface (GUI) or the `cs_user_scripts.py` file (see below). Make sure they are in a format compliant with *Code\_Saturne* (see §3.4.5). There can be several meshes in case of mesh joining or coupling with SYRTHES<sup>3</sup>.
- Go to the directory `DATA` and launch the GUI using the command `./SaturneGUI`.
- If not using the GUI, copy the `DATA/REFERENCE/cs_user_scripts.py` file to `DATA` and edit it, so that the correct run options and paths may be set. For advanced uses, this file may also be used in conjunction with the GUI. Just as with user Fortran subroutines below, settings defined in this file have priority over those defined in the GUI.
- Place the necessary user subroutines in the directory `SRC` (see §3.9). When not using the Interface, some subroutines are compulsory.

#### For all physics:

*compulsory without Graphical User Interface:*

- `usipph` (in `cs_user_parameters.f90`) to specify the turbulence model (and whether  $C_p$  is uniform or not)
- `usipsu` (in `cs_user_parameters.f90`) to define most user parameters
- `cs_user_boundary_conditions` to manage the boundary conditions

*very useful without Graphical User Interface:*

- `usinsc` (in `cs_user_parameters.f90`) to define user scalars (species)
- `usipsc` (in `cs_user_parameters.f90`) to define parameters depending on the number of user scalars.
- `usipgl` (in `cs_user_parameters.f90`) to define some global parameters (error indicators, `idilat` (variable density), `ipucou` (reinforced velocity-pressure coupling), and `iphydr` (improved pressure interpolation in stratified flow))
- `usipes` (in `cs_user_parameters.f90`) to define monitoring points and additional parameters for results outputs

*very useful:*

- `usphyv` (in `cs_user_physical_properties.f90`) to manage variable physical properties (fluid density, viscosity ...)
- `cs_user_initialization` to manage the non-standard initialisations

#### For the “gas combustion” specific physics:

*compulsory without Graphical User Interface:*

- `usppmo` (in `cs_user_parameters.f90`) to select a specific physics module and combustion model

*very useful:*

- `usebu1`, `usd3p1` or `uslwc1` (in `cs_user_parameters.f90`), depending on the selected combustion model, to specify the calculation options for the variables corresponding to combustion model

#### For the “pulverized fuel combustion” specific physics:

*compulsory without Graphical User Interface:*

- `usppmo` (in `cs_user_parameters.f90`) to select the specific physics module

*very useful:*

- `uscpi1`, `user_coal_ini1` (in `cs_user_parameters.f90`) to specify the calculation options for the variables corresponding to pulverized fuel combustion

---

<sup>3</sup>SYRTHES 4 uses meshes composed of 4-node tetrahedra

or `user_fuel_ini1`

**For the “heavy fuel combustion” specific physics:**

(not accessible through the Graphical User Interface in version 3.2.0)

*compulsory:*

- `usppmo` (in `cs_user_parameters.f90`) to select the specific physics module
- `user_fuel_ini1` (in `cs_user_parameters.f90`) to specify the calculation options for the variables corresponding to heavy fuel combustion

**For the “atmospheric module” specific physics:**

*compulsory without Graphical User Interface:*

- `usppmo` (in `cs_user_parameters.f90`) to select the specific physics module

*very useful:*

- `usati1` (in `cs_user_parameters.f90`) to manage the reading of the meteo file
- `usadtvor` `usatsoil` (in `cs_user_atmospheric_model.f90`) to manage the options to the specific physics

**For the “electric module” specific physics (Joule effect and electric arcs):**

*compulsory without Graphical User Interface:*

- `usppmo` (in `cs_user_parameters.f90`) to select the specific physics module
- `cs_user_initialization` to initialise the enthalpy in case of Joule effect
- `uselph` (in `cs_user_physical_properties.f90`) to define the physical properties in case of Joule effect

*very useful:*

- `useli1` (in `cs_user_parameters.f90`) to manage the options related to the variables corresponding to the electric module

**For the “Lagrangian module” (dispersed phase):**

(the continuous phase is managed in the same way as for a case of standard physics)

*compulsory without Graphical User Interface:*

- `uslag1` to manage the calculation conditions
- `uslag2` to manage the boundary conditions for the dispersed phase

**For the “compressible module”:**

*compulsory without Graphical User Interface:*

- `usppmo` (in `cs_user_parameters.f90`) to select the specific physics module

*very useful:*

- `uscfx1` and `uscfx2` (in `cs_user_parameters.f90`) to manage the calculation parameters
- `uscfpv` (in `cs_user_physical_properties`) to manage the variable physical properties

The comprehensive list of the user subroutines and their instructions for use are given in §3.9.

- If necessary, place in the directory `DATA` the different external data (input profiles, thermochemical data files, ...)
- Prepare the launch script `runcase`, directly or through the Graphical Interface (see §3.7), or prepare the `DATA/cs_user_scripts.py` file.
- Run the calculation and analyse the results
- If necessary, purge the temporary files (in `RESU/<run_id>` or `<scratch>/<run_id>` directory) (see §3.2.2).



### 3.2.2 Temporary execution directory

During a calculation, *Code\_Saturne* may use a temporary directory for the compilation and the execution if such a “scratch” directory is defined in the global or user `code_saturne.cfg` file. In that case, the result files are only copied at the end in the directory `RESU`. This is recommended if the compute environment includes different file-systems, some better suited to data storage, others to intensive I/O. If this is not the case, there is no point in running in a scratch directory rather than the results directory, as this incurs additional file copies.

If the environment variable `CS_SCRATCHDIR` is defined, its value has priority over the value defined in the preference file so if necessary, it is possible to define a setting specific to a given run using this mechanism.

*WARNING: in case of an error, the temporary directories are not deleted after a calculation, so that they may be used for debugging. They may then accumulate and may hinder the correct operation of the machine.*

**It is therefore essential to remove them regularly.**

### 3.2.3 Execution modes

As explained before, *Code\_Saturne* is composed of two main modules, the Preprocessor and the Kernel (solver). The Preprocessor reads the meshes. The resulting data is transferred to the Kernel through specific files, named `mesh_input`, or placed in a directory of that name when multiple meshes are imported. In a standard calculation, the files are not copied from the temporary execution directory to the results directory, as they have no interest for data analysis, and are considered “internal” files, whose format or contents is not guaranteed not to change between *Code\_Saturne* versions.

Yet, the Preprocessor does not run in parallel and may require a large amount of memory. The launch scripts therefore allow specifically choosing which modules to run, either through the GUI or through the `cs.user_scripts.py` file:

If a `mesh_input` file or directory is defined (which may be either a `mesh_input` from a previous Preprocessor run or a `mesh_output` from a previous solver run), the script will copy or link it to the execution directory, and the Preprocessor will not be run again.

If `domain.exec_kernel = False`, the Kernel will not be run. This is useful when only the mesh import stage is required.

In a similar manner, the Kernel accepts several command-line options relative to execution mode, notably `domain.solver_args.kernel = '--preprocess'` or `'--quality'`, restricting the run to the preprocessing stages, or preprocessing stages augmented by mesh quality criteria computation. Whenever the preprocessing stages defined lead to an effective mesh modification, a `mesh_output` file is produced, which can be used directly as an input for a successive calculation.

The GUI presents the range of possibilities as 4 execution modes:

- **mesh import:** the Preprocessor is run to transform one or more meshes into an internal `mesh_input` file (or directory in case of multiple meshes).
- **mesh preprocessing:** the Kernel is run in preprocessing mode, so as to handle all mesh modification operations, such as joining, periodicity, smoothing, *etc.* If a `mesh_input` file or directory is provided, it is used directly. Otherwise, mesh import is run first.
- **mesh quality criteria:** similar to preprocessing, with the addition of mesh quality criteria computation, and post-processing output of those criteria. Some additional mesh consistency checks are also run.
- **standard:** this includes preprocessing, followed by a standard computation.



Note that to allow preprocessing in multiple passes, all defined preprocessing operations are run even on previously preprocessed meshes. In most cases, those will not produce additional changes (such as joining already joined meshes), but in the case of mesh smoothing, they might lead to small changes. So when using a previously preprocessed mesh it is recommended to skip the preprocessing stage.

It is encouraged to separate the preprocessing and calculation runs, as this not only speeds up calculations, but also ensures that the mesh is identical, regardless of the architecture or number of processors it is run on. Indeed, when running the same pre-processing stages such as mesh joining on a different machine or a different number of processors, very minor floating-point truncation errors may lead to very slightly different preprocessed meshes.

Note also that mesh partitioning is done directly by the Kernel. Depending on the partitioning algorithm used, a partition map (`partition_output/domain_number_*`) may be output, allowing the use of the same partitioning in an further calculations. By default, this file is output when using graph-based partitioner, which may use randomization and do not guarantee a reproducible output, and is not output when using a deterministic space-filling curve based partition.

If the code was built only with a serial partitioning library, graph-based partitioning may best be run in a serial pre-processing stage. In some cases, serial partitioning might also provide better partitioning quality than parallel partitioning, so if both are available, comparing the performance of the code may be worthwhile, at least for calculations expected to run for many iterations.

### 3.2.4 Environment variables

Setting a few environment variables specific to *Code\_Saturne* allows modifying its default behaviour. The environment variables used by *Code\_Saturne* are described here:

#### CS\_SCRATCHDIR

Allows defining the execution directory (see §3.2.2), overriding the default path or settings from the global or user `code_saturne.cfg`.

#### CS\_MPIEXEC\_OPTIONS

This variable allows defining extra arguments to be passed to the MPI execution command by the run scripts. If this option is defined, it will have priority over the value defined in the preference file (or by computed defaults), so if necessary, it is possible to define a setting specific to a given run using this mechanism. This may be useful when tuning the installation to a given machine, for example experimenting MPI mapping and “bind to core” features.

### 3.2.5 Interactive modification of the target time step

During a calculation, it is possible to change the limit time step number (`ntmabs`) specified through the GUI or in `cs_user_parameters.f90`. To do so, a file named `control_file` must be placed in the execution directory (see §3.2.2). This file must contain a line indicating the value of the new limit number of time steps.

If this new limit has already been reached, *Code\_Saturne* will stop properly at the end of the current time step (the results and restart files will be written correctly).

This procedure allows the user to stop a calculation in a clean and interactive way whenever they wish.

The `control_file` may also contain a few other commands:

max_time_step	<time_step_number>
max_time_value	<time_value>
checkpoint_time_step	<time_step_number>
checkpoint_time_value	<time_value>
checkpoint_wall_time	<wall_clock_time>
checkpoint_time_step_interval	<time_step_interval>
checkpoint_time_value_interval	<time_interval>
checkpoint_wall_time_interval	<wall_time_interval>
control_file_wtime_interval	<wall_time_interval>
postprocess_time_step	<time_step_number> [writer_id]
postprocess_time_value	<time_step_value> [writer_id]

### 3.3 Case preparer

The case preparer command `code_saturne create` automatically creates a study directory according to the typical architecture and copies and pre-fills an example of calculation launch script.

The syntax of `code_saturne create` is as follows:

```
code_saturne create --study STUDY CASE_NAME1 CASE_NAME2...
creates a study directory STUDY with case subdirectories CASE_NAME1 and CASE_NAME2... If no case
name is given, a default case directory called CASE1 is created.
```

```
code_saturne create --case Flow3 --case Flow4
executed in the directory STUDY adds the case directories Flow3 and Flow4. Whenever multiple
cases are created simultaneously, it is assumed they may be coupled, so a runcase_coupling file
and RESU_COUPLING directory are also created.
```

In the directory `DATA`, the `code_saturne create` command places a subdirectory `REFERENCE` containing examples of thermochemical data files used for pulverised coal combustion, gas combustion, electric arcs, or a meteo profile. The file to be used for the calculation must be copied directly in the `DATA` directory and its name may either must be unchanged, or be referenced using the GUI or using the `usppmo` subroutine in `cs_user_parameters.f90`. As a rule of thumb, all files in `DATA` except for `SaturneGUI` are copied, but subdirectories are not.

The `code_saturne create` command also places in the directory `DATA` the launch script for the Graphical User Interface: `SaturneGUI`.

In the directory `SRC`, the `code_saturne create` command creates a subdirectory `REFERENCE` containing all the available user subroutines, and the subdirectory `EXAMPLES` containing examples of user subroutines. Only the user subroutines placed directly under the directory `SRC` will be considered. The others will be ignored.

In the directory `SCRIPTS`, the `code_saturne create` command copies an example of the launch script: `runcase`. The XML file may be specified in the script (see §3.7), and using the GUI sets it automatically.

### 3.4 Supported mesh and post-processing output formats

*Code\_Saturne* supports multiple mesh formats, all of these having been requested at some time by users or projects based on their meshing or post-processing tools. All of these formats have advantages and disadvantages (in terms of simplicity, functionality, longevity, and popularity) when compared to each other. The following formats are currently supported by *Code\_Saturne*:

- SIMAIL (NOPO)
- I-deas universal
- MED

- CGNS
- EnSight 6
- EnSight Gold
- GAMBIT neutral
- Gmsh
- STAR-CCM+
- Catalyst (co-processing)

These formats are described in greater detail in the following sections. Unless a specific option is used, the Preprocessor determines the mesh format directly from the file suffix: “.case” for EnSight (6 or Gold), “.ccm” for STAR-CCM+, “.cgns” for CGNS, “.des” for SIMAIL, “.med” for MED, “.msh” for Gmsh, “.neu” for GAMBIT neutral, “.unv” for I-deas universal.

Note that the preprocessor can read gzipped mesh files directly (for Formats other than MED or CGNS, which use specific external libraries) on most machines.

### 3.4.1 Formats supported for input

#### 3.4.1.1 NOPO/SIMAIL (INRIA/Distene)

This format is output by SIMAIL, which was used heavily at EDF until a few years ago. We do not currently handle cylindrical or spherical coordinates, but it seems that SIMAIL always outputs meshes in Cartesian coordinates, even if points have been defined in another system. Most “classical” element types are usable, except for pyramids.

Note that depending on the architecture on which a file was produced by SIMAIL<sup>4</sup>, it may not be directly readable by SIMAIL on a different machine, while this is not a problem for the Preprocessor, which automatically detects the byte ordering and the 32/64 bit variant and adjusts accordingly.

Default extension:	.des
File type:	semi-portable “Fortran” binary (IEEE integer and floating-point numbers on 4 or 8 bytes, depending on 32 or 64 bit SIMAIL version, bytes also ordered based on the architecture)
Surface elements:	triangles, quadrangles (+ volume element face references)
Volume elements:	tetrahedra, prisms, hexahedra
Zone selection:	element face references and volume sub-domains (interpreted as numbered groups)
Compatibility:	all files of this type as long as the coordinate system used is Cartesian and not cylindrical or spherical
Documentation:	Simail user documentation and release notes or MODULEF documentation: <a href="http://www-rocq.inria.fr/modulef">http://www-rocq.inria.fr/modulef</a> Especially: <a href="http://www-rocq.inria.fr/modulef/Doc/FR/Guide2-14/node49.html">http://www-rocq.inria.fr/modulef/Doc/FR/Guide2-14/node49.html</a>

#### 3.4.1.2 I-deas universal file

This format was very popular in the 1990’s and early 2000’s, and though the I-deas tool has not focused on the CFD (or even meshing) market since many years, it is handled (at least in part) by many tools,

<sup>4</sup>“little endian” on Intel or AMD processors, or “big endian” on most others, and starting with SIMAIL 7, 32-bit or 64-bit integer and floating-point numbers depending on architecture

and may be considered as a major “legacy” format. It may contain many different datasets, relative to CAD, meshing, materials, calculation results, or part representation. Most of these datasets are ignored by *Code\_Saturne*, and only those relative to vertex, element, group, and coordinate system definitions are handled.

This format’s definition evolves with I-deas versions, albeit in a limited manner: some datasets are declared obsolete, and are replaced by others, but the definition of a given dataset type is never modified. Element and Vertex definitions have not changed for many years, but group definitions have gone through several dataset variants through the same period, usually adding minor additional group types not relevant to meshing. If one were to read a file generated with a more recent version of I-deas for which this definitions would have changed with no update in the Preprocessor, as the new dataset would be unknown, it would simply be ignored.

Note that this is a text format. Most element types are handled, except for pyramids.

Default extension:	.unv
File type:	text
Surface elements:	triangles, quadrangles
Volume elements:	tetrahedra, prisms, hexahedra
Zone selection:	colors (always) and named groups
Compatibility:	I-deas ( <i>Master Series</i> 5 to 9, <i>NX Series</i> 10 to 12) at least
Documentation:	Online I-deas NX Series documentation

### 3.4.1.3 GAMBIT neutral

This format may be produced by Ansys FLUENT’s GAMBIT meshing tool. As this tool does not export meshes to other formats directly handled by the Preprocessor (though FLUENT itself may export files to the CGNS or I-deas universal formats), it was deemed useful to enable the Preprocessor to directly read files in GAMBIT neutral format.

Note that this is a text format. “Classical” element types are usable.

Default extension:	.neu
File type:	text
Surface elements:	triangles, quadrangles
Volume elements:	tetrahedra, pyramids, prisms, hexahedra
Zone selection:	boundary conditions for faces, element groups for cells (interpreted as named groups)
Documentation:	GAMBIT on-line documentation

### 3.4.1.4 EnSight 6

This format is used for output by the Harpoon meshing tool, developed by Sharc Ltd (also the distributor of EnSight for the United Kingdom). This format may represent all “classical” element types.

Designed for post processing, it does not explicitly handle the definition of surface patches or volume zones, but allows the use of many *parts* (i.e. groups of elements) which use a common vertex list. A possible convention (used at least by Harpoon) is to add surface elements to the volume mesh, using one *part* per group. The volume mesh may also be separated into several *parts* so as to identify different zones. As *part* names may contain up to 80 characters, we do not transform them into groups (whose names could be unwieldy), so we simply convert their numbers to group names.

Also note that files produced by Harpoon may contain badly oriented prisms, so the Preprocessor orientation correction option (`--reorient`) may must be used. Meshes built by this tool also contain hanging nodes, with non-conforming elements sharing some vertices. Mesh joining must thus also be used, and is not activated automatically, as the user may prefer to specify which surfaces should be joined, and which ones should not (*i.e.* to conserve thin walls).

Default extension:	<b>.case</b>
File type:	text file (extension <i>.case</i> ), and text, binary, or Fortran binary file with ( <i>.geo</i> extension), describing the integers describing integers and floats in the IEEE format, using 32 bits
Surface elements:	triangles, quadrangles
Volume elements:	tetrahedra, pyramids, prisms, hexahedra
Zone selection:	part numbers interpreted as numbered groups
Compatibility:	All files of this type
Documentation:	on-line documentation, also available at: <a href="http://www3.ensight.com/EnSight10_Docs/UserManual.pdf">www3.ensight.com/EnSight10_Docs/UserManual.pdf</a>

### 3.4.1.5 Gmsh

This format is used by the free [Gmsh](http://gmsh.org/) tool. This tool has both meshing and post-processing functionality, but *Code\_Saturne* only imports meshes.

Note that some meshes produced by Gmsh may contain some badly oriented elements, so the Preprocessor's `-reorient` option may be necessary.

The Preprocessor handles versions 1 and 2 of this array. In version 1, two labels are associated with each element: the first defines the element's physical entity number, the second defines its elementary entity number. Using version 2, it is possible to associate an arbitrary number of labels with each element, but files produced by Gmsh use 2 labels, with the same meanings as with version 1.

We chose to convert physical entity numbers to groups. It is possible to build a mesh using Gmsh without defining any physical entities (in which case all elements will belong to the same group, but the Gmsh documentation clearly says that geometric entities are to be used so as to group elementary entities having similar "physical" meanings.

So as to obtain distinct groups with a mesh generated by Gmsh, it is thus necessary for the user to define physical entities. This requires an extra step, but allows for fine-grained control over the groups associated with the mesh, while using only elementary entities could lead to a high number of groups.

Default extension:	<b>.msh</b>
File type:	text or binary file
Surface elements:	triangles, quadrangles
Volume elements:	tetrahedra, pyramids, prisms, hexahedra
Zone selection:	physical entity numbers interpreted as numbered groups
Compatibility:	all files of this type
Documentation:	included documentation, also available at: <a href="http://www.geuz.org/gmsh">http://www.geuz.org/gmsh</a>

## 3.4.2 Formats supported for input or output

### 3.4.2.1 EnSight Gold

This format may represent all "classical" element types, as well as arbitrary polygons and convex polyhedra.

This format evolves slightly from one EnSight version to another, keeping backwards compatibility. For example, polygons could not be used in the same *part* as other element types prior to version 7.4, which removed this restriction and added support for polyhedra. Version 7.6 added support for material type definitions.

This format offers many possibilities not used by *Code\_Saturne*, such as defining values on part of a mesh only (using "undefined" marker values or partial values), assigning materials to elements, defining rigid motion, or defining per-processor mesh parts with ghost cells for parallel runs. Note

that some libraries allowing direct EnSight Gold support do not necessarily support the whole format specification. Especially, VTK does not support material types. Also, both EnSight Gold (8.2 and above) and VTK allow for automatic distribution, reducing the usefulness of pre-distributed meshes with per-processor files.

Note that when using ParaView, if multiple parts (i.e. meshes) are present in a give case, using the “Extract Blocks” filter is required to separate those parts and obtain a proper visualization. The VisIt software does not seem to handle multiple parts in an EnSight case, so different meshes must be assigned to different *writers* (see §7.1) when using this tool.

This format may be used as an input format, similar to EnSight 6. Compared to the latter, each *part* has its own coordinates and vertex connectivity, so as a convention, we consider that surface or volume zones may only be considered to be part of the same mesh if the file defines vertex IDs (which we consider to be unique vertex labels). In this case, *part* numbers are interpreted as group names. Without vertex IDs, only one part is read, and no groups are assigned.

Default extension:	directory { <i>case_name</i> }. <b>ensight</b> , containing a file with the <b>.case</b> extension
File type:	multiple binary or text files
Surface elements:	triangles, quadrangles, polygons
Volume elements:	tetrahedra, pyramids, prisms, hexahedra, convex polyhedra
Zone selection:	possibility of defining element materials (not used), or interpret part number as group name if vertex IDs are given
Compatibility:	files readable by EnSight 7.4 to 10.0, as well as tools based on the <b>VTK</b> library, especially ParaView ( <a href="http://www.paraview.org">http://www.paraview.org</a> )
Documentation:	online documentation, also available at: <a href="http://www3.ensight.com/EnSight10_Docs/UserManual.pdf">www3.ensight.com/EnSight10_Docs/UserManual.pdf</a>

### 3.4.2.2 MED

Initially defined by EDF R&D, this format (*Modèle d'échanges de Données*, or *Model for Exchange of Data*) has been defined and maintained through a MED working group comprising members of EDF R&D and CEA (the *Code\_Saturne* team being represented). This is the reference format for the **SALOME** environment. This format is quite complete, allowing the definition of all “classical” element types, in nodal or descending connectivity. It may handle polygonal faces and polyhedral cells, as well as the definition of structured meshes.

This format, which requires a library also depending on the free HDF5 library, allows both for reading and writing meshes with their attributes (“families” of color/attribute and group combinations), as well as handling calculation data, with the possibility (unused by *Code\_Saturne*) of defining variables only on a subset (“profile”) of a mesh.

The MED library is available under a **LGPL** licence, and is even packaged in some Linux distributions (at least Debian and Ubuntu). *Code\_Saturne* requires at least MED 3.0.2, which in turn requires HDF5 1.8. This format is upwards-compatible with MED 2.3, so files in that version of the format may be read, though not output.

Default extension:	.med
File type:	portable binary, based on the HDF5 library ( <a href="http://www.hdfgroup.org/HDF5/index.html">http://www.hdfgroup.org/HDF5/index.html</a> )
Surface elements:	triangles, quadrangles, simple polygons
Volume elements:	tetrahedra, pyramids, prisms, hexahedra, simple polyhedra
Zone selection:	element families ( <i>i.e.</i> colors and groups)
Input compatibility:	MED 2.3 or MED 3.0 (only unstructured nodal connectivity is supported)
Output compatibility:	MED 3.0 and above
Documentation:	on-line documentation. Download link at <a href="http://files.salome-platform.org/Salome/other/med-3.0.6.tar.gz">http://files.salome-platform.org/Salome/other/med-3.0.6.tar.gz</a>

### 3.4.2.3 CGNS

Promoted especially by the AIAA, NASA, Boeing Commercial, and ANSYS ICEM CFD (as well as ONERA in France), this format (*CFD General Notation System*) is quite well established in the world of CFD. The concept is similar to that of MED, with a bigger emphasis on normalization of variable names or calculation information, and even richer possibilities. Contrary to MED, the first version of this format was limited to multi-bloc structured meshes, unstructured meshes having been added in CGNS 2.

Slightly older than MED, this library was free from the start, with a good English documentation, and is thus much better known. It is more focused on CFD, where MED is more generic. A certain number of tools accompany the CGNS distribution, including a mesh visualizer (which does not handle polygonal faces although the format defines them), and an interpolation tool.

We should be able to read almost any mesh written in this format, though meshes with over-set interfaces may not be usable for a calculation (calculations with over-set interfaces may be possible in the context of coupling *Code\_Saturne* with itself but with two separate meshes). Other (abutting) interfaces are not handled automatically (as there are at least 3 or 4 ways of defining them, and some mesh tools do not export them<sup>5</sup>), so the user is simply informed of their existence in the Preprocessor's log file, with a suggestion to use an appropriate conformal joining option. Structured zones are converted to unstructured zones immediately after being read.

Boundary condition information is interpreted as groups with the same name. The format does not yet provide for selection of volume elements, as only boundary conditions are defined in the model (and can be assigned to faces in the case of unstructured meshes, or vertices in any case). Note that boundary conditions defined at vertices are not ignored by the Preprocessor, but are assigned to the faces of which all vertices bear the same condition.<sup>6</sup>

The Preprocessor also has the capability of building additional volume or surface groups, based on the mesh sections to which cells or faces belong. This may be activated using a sub-option of the mesh selection, and allows obtaining zone selection information from meshes that do not have explicit boundary condition information but that are subdivided in appropriate zones or sections (which depends on the tool used to build the mesh).

When outputting to CGNS, an unstructured connectivity is used for the calculation domain, with no face joining information or face boundary condition information.<sup>7</sup>

Though many tools support CGNS, that support is often quite disappointing, at least for unstructured meshes. Thus, some editors seem to use different means to mark zones to associate with boundary

<sup>5</sup>For example, ICEM CFD can join non-conforming meshes, but it exports joining surfaces as simple boundary faces with user-defined boundary conditions.

<sup>6</sup>If one of a face's vertices does not bear a boundary condition, that condition is not transferred to the face.

<sup>7</sup>Older versions of the documentation specified that a field must be defined on all elements of a zone, so that adding faces on which to base boundary conditions to a volume mesh would have required also defining volume fields on these faces. More recent versions of the documentation make it clear that a field must be defined on all elements of maximum dimension in a zone, not on all elements.



conditions than the ones recommended in the CGNS documentation, and some behaviours are worse. Also, many readers do not allow the user to choose between multiple CGNS bases (meshes in the *Code\_Saturne* sense), so when outputting to CGNS, it may be necessary to output each post-processing mesh using a separate writer. VisIt 2.4.2 may fail to read a volume mesh output by *Code\_Saturne*, but read a surface mesh correctly, while the same volume mesh may be read with no problems by EnSight 10. The support of polygons (*ngons* in the CGNS standard), is even worse, and even the verification tools published alongside the CGNS library were recently unable to handle them, and reported errors in valid files containing such elements. For mesh input, some ICEM CFD versions used a CGNS 3 beta library, which led to some issues. CGNS 3 output from ICEM CFD 13 is known to work well with *Code\_Saturne*, but that same version is unable to read files generated by *Code\_Saturne*, as it seems to “cheat” with CGNS version numbers and confuses CGNS 3 and 3.1 specs. ICEM 14 seems to have fixed that bug.

Default extension:	.cgns
File type:	portable binary (uses the ADF library specific to CGNS, or HDF5)
Surface elements:	triangles, quadrangles, simple polygons
Volume elements:	tetrahedra, pyramids, prisms, hexahedra (simple polyhedra allowed by CGNS 3 but not supported yet)
Zone selection:	Surface zone selection using boundary conditions, no volume zone selection, but the Preprocessor allows creation of groups associated to zones or sections in the mesh using mesh selection sub-options
Input compatibility:	CGNS 2.5 or CGNS 3.1
Output compatibility:	CGNS 3.1
Documentation:	See CGNS site: <a href="http://www.cgns.org">http://www.cgns.org</a>

### 3.4.2.4 STAR-CCM+

This polyhedral format is the current CD-Adapco format, and is based on CD-Adapco's libccmio, which is based on ADF (the low-level file format used by CGNS prior to the shift to HDF-5). libccmio comes with a version of ADF modified for performance, but also works with a standard version from CGNS.

Currently, geometric entity numbers are converted to numbered groups, with the corresponding names printed to the Preprocessor log. Depending on whether the names were generated automatically or set by the user, it would be preferable to use the original group names rather than base their names on their numbers.

This format may also be used for output, though its limitations make this a less general solution than other output formats: only 3D meshes are handled, though values can be output on boundary face regions (which may not overlap). As such, to ensure consistency, output using this format is limited as follows:

- output of the full volume mesh and cell or vertex data on that mesh is handled normally.
- output of the full surface mesh and per face data on that mesh handled normally, only if output of the full volume mesh to this format is also enabled. It is ignored otherwise.
- output of sub-meshes or meshes built during the preprocessing stage and all other data is ignored.

As such, this format may be useful for interoperability of data with a CCMIO-based tool-chain, but simultaneously using another output format to visualize possible error output is recommended.

The CCMIO library is distributed freely by CD-Adapco upon demand.



Default extension:	.ccm
File type:	binary file using modified ADF library.
Surface elements:	polygons
Volume elements:	polyhedra
Zone selection:	named face and cell sets (interpreted as numbered groups, with names appearing in log)
Compatibility:	all files of this type?
Documentation:	documentation and source code provided by CD-adapco

### 3.4.3 Formats supported for output only

#### 3.4.3.1 Catalyst

This is not a “true” output format in the sense that output is not written directly to file, but is exported to the Catalyst co-processor. In turn, this co-processor will execute operations based on a special ParaView Python script, and directly generate output such as images or movies.

Co-processing scripts may be generated using ParaView 4, with the CoProcessing plugin, using initial output in another format (such as EnSight Gold). A *Code\_Saturne* postprocessing writer will try to read a script named `<writer_name>.py`, which should be placed in a case's DATA directory.

Note that this output is still an experimental stage, and is heavily dependent on ParaView. Some operations may work very well, while other, similar operations may fail.

Default extension:	not applicable
File type:	co-processing
Surface elements:	triangles, quadrangles, polygons
Volume elements:	tetrahedra, pyramids, prisms, hexahedra, convex polyhedra
Compatibility:	Catalyst from <a href="#">ParaView</a> 4.0
Documentation:	online documentation and Wiki, at: <a href="http://paraview.org/Wiki/Main_Page">http://paraview.org/Wiki/Main_Page</a>

### 3.4.4 Meshing tools and associated formats

Most often, the choice of a mesh format is linked to the choice of a meshing tool. Still, some tools allow exporting a mesh under several formats handled by *Code\_Saturne*. This is the case of FLUENT and ICEM CFD, which can export meshes to both the I-deas universal and CGNS formats (FLUENT's GAMBIT is also able to export to I-deas universal format).

Traditionally, users exported files to the I-deas universal format, but it does not handle pyramid elements, which are often used by these tools to transition from hexahedral to tetrahedral cells in the case of hybrid meshes. The user is encouraged to export to CGNS, which does not have this limitation.

Tools related to the SALOME platform should preferably use SALOME's native MED format (export to I-deas universal is also possible, but has some limitations).

### 3.4.5 Meshing remarks

**WARNING:** Some turbulence models ( $k-\varepsilon$ ,  $R_{ij}-\varepsilon$  SSG, ...) used in *Code\_Saturne* are “High-Reynolds” models. Therefore the size of the cells neighbouring the wall must be greater than the thickness of the viscous sub-layer (at the wall,  $y^+ > 2.5$  is required, and  $30 < y^+ < 100$  is preferable). If the mesh does not match this constraint, the results may be false (particularly if thermal phenomena are involved). For more details on these constraints, see the keyword `iturb`.

## 3.5 Preprocessor command line options

The main options are:

- **--help**: gives a summary of the different command line options
- **<mesh>**: the last argument is used to specify the name of the mesh file. The launch script automatically calls the Preprocessor for every mesh in the `MESHES[]` list specified by the user.
- **--reorient**: try to re-orient badly-oriented cells if it is necessary to compensate for mesh-generation software whose output does not conform to the format specifications.

## 3.6 Kernel command line options

In the standard cases, the compilation of *Code\_Saturne* and its execution are entirely controlled by the launch script. The potential command line options are passed through user modifiable variables at the beginning of the `cs_user_scripts.py` file (this file may be copied from the `DATA/REFERENCE` to the `DATA` and edited). This way, the user only has to fill these variables and doesn't need to search deep in the script for the Kernel command line. For more advanced usage, the main options are described below:

- **--app-name**: specifies the application name. This is useful only in the case of code coupling, where the application name is used to distinguish between different code instances launched together.
- **--mpi**: specifies that the calculation is running with MPI communications. The number of processors used will be determined automatically by the Kernel. With most MPI implementations, the code will detect the presence of an MPI environment automatically, and this option is redundant. It is only kept for the rare case in which the MPI environment might not be detected.
- **--preprocess**: triggers the preprocessing-only mode. The code may run without any Interface parameter file nor any user subroutine. Only the initial operations such as mesh joining and modification are executed.
- **-q** or **--quality**: triggers the verification mode. The code may run without any Interface parameter file nor any user subroutine. This mode includes the preprocessing stages, and adds elementary tests:
  - the quality criteria of the mesh are calculated (non-orthogonality angles, internal faces offset, ...) and corresponding visualizable post-processing output is generated.
  - a few additional mesh consistency tests are run.
- **--benchmark**: triggers the benchmark mode, for a timing of elementary operations on the machine. A secondary option **--mpitrace** can be added. It is to be activated when the benchmark mode is used in association with a MPI trace utility. It restricts the elementary operations to those implying MPI communications and does only one of each elementary operation, to avoid overfilling the MPI trace report.  
This command is to be placed in the `textttdomain.solver_args` variable in the `cs_user_scripts.py` file to be added automatically to the Kernel command line.
- **--log n**: specifies the destination of the output for a single-processor calculation or for the processor of rank 0 in a parallel calculation.
  - n=0**: output directed towards the standard output

**n=1**: output redirected towards a file **listing** (default behaviour)  
This option can be specified in the **domain.logging\_args** field of the user script.

- **--logp n**: specifies the destination of the output for the processors of rank 1 to  $N - 1$  in a calculation in parallel on  $N$  processors (*i.e.* the redirection of all but the first processor).
  - n=-1**: no output for the processors of rank 1 to  $N - 1$  (default behaviour).
  - n=0**: no redirection. Every processor will write to the standard output. This might be useful in case a debugger is used, with separate terminals for each processor.
  - n=1**: one file for the output of each processor. The output of the processors of rank 1 to  $N - 1$  are directed to the files **listing\_n0002** to **listing\_nN**. This option can be specified in the **domain.logging\_args** field of the user script.
- **-p <filename>** or **--param <filename>**: specifies the name of the GUI parameter file to use for the calculation.  
The value of **<filename>** is to be defined by the **--param** option of **code\_saturne run**, either directly or in the standard **runcase** script (the file will be searched for in the **data** directory, though an absolute path name may also be defined).
- **-h** or **--help**: to display a summary of the different command line options.

### 3.7 Launch scripts

The case preparer command **code\_saturne create** places an example of launch script, **runcase**, in the **SCRIPTS** directory. This script is quite minimalist and is known to work on every architecture *Code\_Saturne* has been tested on. If a batch system is available, this script will contain options for batch submission. The script will then contain a line setting the proper **PYTHONPATH** variable for *Code\_Saturne* to run. Finally, it simply contains the **code\_saturne run** command, possible with a **--param** option when a parameters file defined by the GUI is used. Other options recognized by **code\_saturne run** may be added.

In the case of a coupled calculation, this script also exists, and may be used for preprocessing stages, but an additional **runcase\_coupling** is added in the directory above the coupled case directories, and may be used to define the list of coupled cases, as well as global options, such as MPI options of the temporary execution directory. An additional **runcase\_batch** file will contain batch submission options when a batch system is available (and is the file that should be submitted when using a batch system).

When not using the GUI, or if additional options must be accessed, the **cs\_user\_scripts.py** file may be copied from the **DATA/REFERENCE** to the **DATA** and edited. This file contains several Python functions:

- **define\_domain\_parameter\_file** allows defining the choice of a parameters file produced by the GUI. This is generally not useful, as the parameters file may be directly defined in **runcase** or **runcase\_coupling**, or passed as an option to **code\_saturne run**, but could be useful when running more complex parametric scripts, and is provided for the sake of completeness.
- **define\_domain\_parameters** allows defining most parameters relative to case execution for the current domain, including advanced options not accessible through the GUI. This function is the most important one in the user scripts file, and contains descriptions of the various options. Note that in most examples, setting of options is preceded by a **if domain.param == None:** line, ensuring the settings are only active if no GUI-defined parameters file is present. This is used to prevent accidental override of parameters defined by the GUI: parameters defined through the user script have priority over the GUI parameters file, so if both are used, these tests may be removed for parameters which should be defined through user scripts.
- **define\_case\_parameters** allows defining most parameters relative to the global calculation, such as the number of processors or the execution directory. To avoid potentially conflicting

definitions, this function is ignored for coupled calculations, where the corresponding parameters may be defined in the `runcase_coupling` script.

- `define_mpi_environment` allows defining advanced MPI parameters or redefining MPI options if the automatic settings are incorrect, and its use should only rarely be necessary. To avoid potentially conflicting definitions, this function is ignored for coupled calculations, where the corresponding parameters may be defined in the `runcase_coupling` script.

## 3.8 Graphical User Interface

A Graphical User Interface is available with *Code\_Saturne*. This Interface creates or reads an XML file according to a specific *Code\_Saturne* schema which is then interpreted by the code.

In version 3.2.0, the Graphical Interface manages calculation parameters, standard initialisation values and boundary conditions for standard physics, pulverised fuel combustion, gas combustion, atmospheric flows, Lagrangian module, electrical model, compressible model and radiative transfers (user subroutines can still be completed though).

The Interface is optional. Every data that can be specified through the Interface can also be specified in the user subroutines. In case of conflict, all calculation parameters, initialisation value or boundary condition set directly in the user subroutines will prevail over what is defined by the Interface. However, it is no longer necessary to redefine everything in the user subroutines. Only what was not set or could not be set using the Graphical Interface should be specified.

**WARNING:** There are some limitations to the changes that can be made between the Interface and the user routines. In particular, it is not possible to specify a certain number of solved variables in the Interface and change it in the user routines (for example, it is not possible to specify the use of a  $k - \varepsilon$  model in the Interface and change it to  $R_{ij} - \varepsilon$  in `cs_user_parameters.f90`, or to define additional scalars in `cs_user_parameters.f90` with respect to the Interface). Also, all boundaries should be referenced in the Interface, even if the associated conditions are intended to be modified in `cs_user_boundary_conditions`, and their nature (entry, outlet, wall<sup>8</sup>, symmetry) should not be changed.

For example, in order to set the boundary conditions of a calculation corresponding to a channel flow with a given inlet velocity profile, one should:

- set the boundary conditions corresponding to the wall and the output using the Graphical Interface
- set a dummy boundary condition for the inlet (uniform velocity for instance) - set the proper velocity profile at inlet in `cs_user_boundary_conditions`. The wall and output areas must not appear in `cs_user_boundary_conditions`. The dummy velocity entered in the Interface will not be taken into account.

The Graphical User Interface is launched with the `./SaturneGUI` command in the directory `DATA`. The first step is then to load an existing parameter file (in order to modify it) or to open a new one. The headings to be filled for a standard calculation are the following:

- Identity and paths: definition of the calculation directories (STUDY, CASE, DATA, SRC, SCRIPTS, MESH).
- Calculation environment: definition of the mesh file(s), stand-alone execution of the Preprocessor module (used by the Interface to get the groups of the boundary faces).
- Thermophysical models: physical model, ALE mobile mesh features, turbulence model, thermal model, coupling with SYRTHES.
- Additional scalars: definition, initialisation of the scalars, and physical characteristics.

---

<sup>8</sup>smooth and rough walls are considered of the same nature

- Physical properties: reference pressure, fluid characteristics, gravity. It is also possible to write user laws for the density, the viscosity, the specific heat and the thermal conductivity in the interface through the use of a formulae interpreter.
- Volume conditions: initialisation of the variables, and definition of the zones where to apply head losses or source terms.
- Boundary conditions: definition of the boundary conditions for each variable. The colors of the boundary faces may be read directly from a “preprocessor.log\*” files created by the Preprocessor or a “listing” file from a previous Kernel run.
- Numerical parameters: number and type of time step, advanced parameters for the numerical solution of the equations.
- Calculation control: parameters concerning the time averages, time step, location of the probes where some variables will be monitored over time, definition of the frequency of the outputs in the calculation listing and in the chronological records and of the EnSight outputs. The item *Profiles* allows to save, with a given frequency, 1D profiles on an axis defined from two points provided by the user.
- Calculation management: management of the calculation restarts, updating of the launch script (temporary execution directory, parallel computing, user data or result files, ...) and interactive launch of the calculation.

The *Code\_Saturne* tutorial [14] offers a step-by-step guidance to the setting up of some simple calculations with the *Code\_Saturne* Interface.

To launch *Code\_Saturne* using an XML parameter file, the name of the file must be given using the `--param` option of `code_saturne run` in the launch script (see §3.7). When the launch script is edited from the Interface (Calculation management → Prepare batch analysis), this option is set automatically.

## 3.9 User subroutines

### 3.9.1 Preliminary comments

The user can run the calculations with or without an interface, with or without the user subroutines. Without interface, some user subroutines are needed (see §3.2.1). With interface, all the user subroutines are optional.

The parameters can be read in the interface and then in the user subroutines. In the case that a parameter is specified in the interface and in a user subroutine, it is the value in the user subroutine that is taken into account. For this reason, all the examples of user subroutines are placed in the **EXAMPLES** directory by the case setup `code_saturne create` (and available subroutines in the directory **REFERENCE**).

### 3.9.2 Example routines

Some user subroutines may be used for many different user definitions. As including enough examples in those subroutines would make them very difficult to read, these routines provided as templates only, with separate examples in a case's **EXAMPLES** subdirectory of its **SRC** directory.

Example file names are defined by inserting the name of the matching example in the file name. For example, a basic example for `cs_user_boundary_conditions.f90` is provided in `cs_user_boundary_conditions-base.f90`, while an example dedicated to atmospheric flows is provided in `cs_user_boundary_conditions-atmospheric.f90`.

EDF R&D	<i>Code_Saturne</i> version 3.2.0 practical user's guide	<i>Code_Saturne</i> documentation Page 31/202
---------	--	---

The user is encouraged to check what examples are available, and to study those that are to a given setup.

Template user subroutines contain three sections the user may define, marked by the following strings:

- INSERT\_VARIABLE\_DEFINITIONS\_HERE
- INSERT\_ADDITIONAL\_INITIALIZATION\_CODE\_HERE
- INSERT\_MAIN\_CODE\_HERE

Comparing template and example files with a graphical file comparison tool should help the user highlights the matching sections from the examples, so it is recommended as good practice for those not already very familiar with those user subroutines.

### 3.9.3 Main variables

This section presents a non-exhaustive list of the main variables which may be encountered by the user. Most of them should not be modified by the user. They are calculated automatically from the data. However it may be useful to know what they represent. Developers can also refer to [11].

These variables are listed in the alphabetical index at the end of this document (see § 9).

The type of each variable is given: integer [i], real number [r], integer array [ia], real array [ra].

#### 3.9.3.1 Array sizes

<b>ndim:</b>	Space dimension (ndim=3).
<b>ncel:</b>	Number of real cells in the mesh.
<b>ncelet:</b>	Number of cells in the mesh, including the ghost cells of the “halos” (see note 1).
<b>nfac:</b>	Number of internal faces (see note 2).
<b>nfabor:</b>	Number of boundary faces (see note 2).
<b>ncelbr:</b>	Number of cells with at least one boundary face (see note 2).
<b>lndfac:</b>	Size of the array <b>nodfac</b> of internal faces - nodes connectivity (see note 3).
<b>lndfbr:</b>	Size of the array <b>nodfbr</b> of boundary faces - nodes connectivity (see note 3).
<b>nnod:</b>	Number of vertices in the mesh.
<b>nfml:</b>	Number of referenced families of entities (boundary faces, elements, ...).
<b>nprfml:</b>	Number of properties per referenced entity family.
<b>nvar:</b>	Number of solved variables (must be lower than <b>nvrmax</b> ).
<b>nscamx:</b>	Maximum number of scalars solutions of an advection equation, apart from the variables of the turbulence model ( $k, \varepsilon, R_{ij}, \omega, \varphi, \bar{f}, \alpha, \nu_t$ ), that is to say the temperature and other scalars (passive or not, user-defined or not).
<b>nscal:</b>	Effective number of scalars solutions of an advection equation, apart from the variables of the turbulence model ( $k, \varepsilon, R_{ij}, \omega, \varphi, \bar{f}, \alpha, \nu_t$ ), that is to say the temperature and other scalars (passive or not, user-defined or not). These scalars can be divided into two distinct groups: <b>nscaus</b> user-defined scalars and <b>nscapp</b> scalars related to a “specific physics”. <b>nscal=nscaus+nscapp</b> , and <b>nscal</b> must be less than or equal to <b>nscamx</b> .

- nscapp:** Effective number of scalars related to a “specific physics”. These scalars are solutions of an advection equation and distinct from the scalars of the turbulence model ( $k$ ,  $\varepsilon$ ,  $R_{ij}$ ,  $\omega$ ,  $\varphi$ ,  $\bar{f}$ ,  $\alpha$ ,  $\nu_t$ ). They are automatically defined by the choice of the selected specific physics model (gas combustion with Eddy Break-Up model, pulverised coal combustion, ...). For example: mass fractions, enthalpy, ....
- nscaus:** Effective number of user-defined scalars. These scalars are solutions of an advection equation and distinct from the scalars of the turbulence model ( $k$ ,  $\varepsilon$ ,  $R_{ij}$ ,  $\omega$ ,  $\varphi$ ,  $\bar{f}$ ,  $\alpha$ ,  $\nu_t$ ) and from the **nscapp** scalars related to the “specific physics”. For example: passive tracers, temperature (when no specific physics model is selected), ....
- nestmx:** Maximum number of error estimators for Navier-Stokes.
- npromx:** Maximum number of physical properties. They will be stored in the arrays **propce** or **propfb**.
- nproce:** Number of properties defined at the cells. They will be stored in the array **propce**.
- nprofb:** Number of properties defined at the boundary faces. They will be stored in the array **propfb**.
- nvisls:** Number of scalars with variable diffusivity.
- nushmx:** Maximum number of user chronological files (in the case where **ushist** is used).
- nbmomt:** Effective number of calculated time-averages. **nbmomt** must be less than or equal to **nbmomx**.
- nbmomx:** Maximum number of calculated time-averages (default value: 50).
- ndgmox:** Maximum degree of the time-averages (default value: 5).
- nclacp:** Number of coal classes for the pulverised coal combustion module. It is the total number of classes, *i.e.* the sum of the number of classes for every represented coal. **nclacp** must be less than or equal to **nclcpm**.
- nclcpm:** Maximum number of coal classes for the pulverised coal combustion module.

#### NOTE 1: GHOST CELLS - “HALOS”

A cell (real cell) is an elementary mesh element of the spatial discretisation of the calculation domain. The mesh is made of **ncel** cells.

When using periodicity and parallelism, extra “ghost” cells (called “halo” cells) are defined for temporary storage of some information (on a given processor). The total number of real and ghost cells is **ncelet**.

Indeed, when periodicity is enabled, the cells with periodic faces do not have any real neighbouring cell across these particular faces. Their neighbouring cell is elsewhere in the calculation domain (its position is determined by the periodicity). In order to temporarily store the information coming from this “distant” neighbouring cell, a ghost cell (“halo”) is created.

The same kind of problem exists in the case of a calculation on parallel machines: due to the decomposition of the calculation domain, some cells no longer have access to all their neighbouring cells, some of them being treated by another processor. The creation of ghost cells allows to temporarily store the information coming from real neighbouring cells treated by other processors.

The variables are generally arrays of size **ncelet** (number of real and fictitious cells). The calculations (loops) are made on **ncel** cells (only the real cells, the fictitious cells are only used to store information).

#### NOTE 2: INTERNAL FACES

An internal face is an interface shared by two cells (real or ghost ones) of the mesh. A boundary face is a face which has only one real neighbouring cell. In the case of periodic calculations, a periodic face is an internal face. In the case of parallel running calculations, the faces situated at the boundary of a partition may be internal faces or boundary faces (of the whole mesh);



EDF R&D	<b>Code_Saturne version 3.2.0 practical user's guide</b>	Code_Saturne documentation Page 33/202
---------	--	--

#### NOTE 3: FACES-NODES CONNECTIVITY

The faces - nodes connectivity is stored by means of four integer arrays: **ipnfac** and **nodfac** for the internal faces, **ipnfbr** and **nodfbr** for the boundary faces. **nodfac** (size **lndfac**) contains the list of all the nodes of all the internal faces; first the nodes of the first face, then the nodes of the second face, and so on. **ipnfac** (size: **nfac+1**) gives the position **ipnfac(ifac)** in **nodfac** of the first node of each internal face **ifac**. Therefore, the reference numbers of all the nodes of the internal face **ifac** are: **nodfac(ipnfac(ifac))**, **nodfac(ipnfac(ifac)+1)**, ..., **nodfac(ipnfac(ifac)+1)-1**. In order for this last formula to be valid even for **ifac=nfac**, **ipnfac** is of size **nfac+1** and **ipnfac(nfac+1)** is equal to **lndfac+1**.

The composition of the arrays **nodfbr** and **ipnfbr** is similar.

#### NOTE 4: MODULES

**The user will not modify the existing modules.** This would require the recompilation of the complete version, operation which is not allowed in standard use.

### 3.9.3.2 Geometric variables

The main geometric variables are available in most of the subroutines and directly accessible through the following arrays, defined in the **mesh** module (i.e. **use mesh**).

**cdgfac(ndim,nfac)** [ra]: Coordinates of the centres of the internal faces.

**cdgfb(nfac,nfabor)** [ra]: Coordinates of the centres of the boundary face.

**ifacel(2,nfac)** [ia]: Index-numbers of the two (only) neighbouring cells for each internal face.

**ifabor(nfabor)** [ia]: Index-number of the (unique) neighbouring cell for each boundary face.

**ipnfac(nfac+1)** [ia]: Position of the first node of the each internal face in the array **nodfac** (see note 3 in paragraph 3.9.3.1).

**ipnfbr(nfabor+1)** [ia]: Position of the first node of the each boundary face in the array **nodfbr** (see note 3 in paragraph 3.9.3.1).

**nodfac(lndfac)** [ia]: Index-numbers of the nodes of each internal face (see note 3 in paragraph 3.9.3.1).

**nodfbr(lndfbr)** [ia]: Index-numbers of the nodes of each boundary face (see note 3 in paragraph 3.9.3.1).

**surf(nfac)** [ra]: Surface vector of the internal faces. Its norm is the surface of the face and it is oriented from **ifacel(1,.)** to **ifacel(2,.)**.

**surfbo(nfabor)** [ra]: Surface vector of the boundary faces. Its norm is the surface of the face and it is oriented outwards.

**volume(ncelet)** [ra]: Volume of each cell.

**xyzcen(ndim,ncelet)** [ra]: Coordinates of the cell centres.

**xyznod(ndim,nnod)** [ra]: Coordinates of the mesh vertices.

In addition, other geometric variables are useful for gradients reconstruction. The main variables of this type are the following:

**dijpf(ndim,nfac)** [ra]: For every internal face, the three components of the vector  $\underline{I'J'}$ , where **I'** and **J'** are respectively the orthogonal projections of the neighbouring cell centres **I** and **J** on a straight line orthogonal to the face and passing through its center.

**diipb(ndim,nfabor)** [ra]: For every boundary face, the three components of the vector  $\underline{II'}$ . **I'** is the orthogonal projection of **I**, center of the neighbouring cell, on the straight line perpendicular to the face and passing through its center.



EDF R&D	<b>Code_Saturne version 3.2.0 practical user's guide</b>	Code_Saturne documentation Page 34/202
---------	--	--

**dist(nfac)** [ra]: For every internal face, dot product of the vectors  $\underline{IJ}$  and  $\underline{n}$ . I and J are respectively the centres of the first and the second neighbouring cell. The vector  $\underline{n}$  is the unit vector normal to the face and oriented from the first to the second cell.

**distb(nfabor)** [ra]: For every boundary face, dot product between the vectors  $\underline{IF}$  and  $\underline{n}$ . I is the center of the neighbouring cell. F is the face center. The vector  $\underline{n}$  is the unit vector normal to the face and oriented to the exterior of the domain.

**dofij(ndim,nfac)** [ra]: For every internal face, the three components of the vector  $\underline{OF}$ . O is the intersection point between the face and the straight line joining the centres of the two neighbouring cells. F is the face center.

**icelbr(ncelbr)** [ia]: List of cells having at least one boundary face.

**pond(nfac)** [ra]: For every internal face,  $\frac{FJ.n}{IJ.n}$ . Regarding the quality of mesh, its ideal value is 0.5.

**surfan(nfac)** [ra]: Norm of the surface vector of the internal faces.

**surfbn(nfabor)** [ra]: Norm of the surface of the boundary faces.

### 3.9.3.3 Physical variables

The main physical variables are available in the majority of the subroutines and brought together according to their type in the multidimensional arrays listed below. In some particular subroutines, some variables may be given a more explicit name, in order to ease the comprehension.

**propce(ncelet,nproce)** [ra]: Properties defined at the cell centres. For instance: density, viscosity, ....

**propfb(nfabor,nprofb)** [ra]: Properties defined at the boundary faces. For instance: mass flow across boundary faces, density at boundary faces, ....

**rtp(ncelet,nvar)** [ra]: Array storing the values of the solved variables at the current time step.

**rtpa(ncelet,nvar)** [ra]: Array storing the values of the solved variables at the previous time step.

Progressively, this system is being replaced by the **m [o]: d.ule**, so some variables and properties are now only accessible as fields (either through their name or their id). This is the case for example for mass flux across interior and boundary faces.

#### About **rtp** and **rtpa**

The indexes allowing to mark out the different variables (from 1 to **nvar**) are integers available in a “module” called **numvar**.

For example, **ipr** refers to the variable “pressure”: the pressure in the cell **iel** at the current time step is therefore **rtp(iel,ipr)**.

The list of integers referring to solved variables is given below. These variable index-numbers are not only used for the **rtp** and **rtpa** arrays, but also for some arrays of variable associated options (for instance, **blencv(ik)** is the percentage of second-order convective scheme for the turbulent energy when a corresponding turbulent model is used).

- **ipr**: pressure <sup>9</sup>.
- **iu**: velocity along the X axis.
- **iv**: velocity along the Y axis.

---

<sup>9</sup>**ipr** corresponds to a reduced pressure, from which the standard hydrostatic pressure has been deduced. The total solved pressure is stored in the **PROPCE** array

EDF R&D	<b>Code_Saturne version 3.2.0 practical user's guide</b>	Code_Saturne documentation Page 35/202
---------	--	--

- **iw**: velocity along the Z axis.
- **ik**: turbulent energy, in  $k - \varepsilon$ ,  $k - \omega$  modelling or v2f ( $\varphi$ -model and BL-v2/k model) modelling.
- **ir11**: Reynolds stress R11, in  $R_{ij} - \varepsilon$  or SSG modelling.
- **ir22**: Reynolds stress R22, in  $R_{ij} - \varepsilon$  or SSG modelling.
- **ir33**: Reynolds stress R33, in  $R_{ij} - \varepsilon$  modelling.
- **ir12**: Reynolds stress R12, in  $R_{ij} - \varepsilon$  modelling.
- **ir13**: Reynolds stress R13, in  $R_{ij} - \varepsilon$  modelling.
- **ir23**: Reynolds stress R23, in  $R_{ij} - \varepsilon$  modelling.
- **iep**: turbulent dissipation in  $k - \varepsilon$ ,  $R_{ij} - \varepsilon$  or v2f ( $\varphi$ -model and BL-v2/k model) modelling.
- **iomg**: Specific dissipation rate  $\omega$ , in  $k - \omega$  SST modelling.
- **iphi**: variable  $\varphi = \overline{v^2}/k$  in v2f ( $\varphi$ -model and BL-v2/k model).
- **ifb**: variable  $\overline{f}$  in v2f ( $\varphi$ -model).
- **ial**: variable  $\alpha$  in elliptic blending models (BL-v2/k and EBRSM).
- **inusa**: variable  $tildenu_t$  in Spalart-Allmaras model.
- **isca(j)**: scalar j ( $1 \leq j \leq \text{nscal}$ ).

Concerning the solved scalar variables (apart from the variables pressure,  $k$ ,  $\varepsilon$ ,  $R_{ij}$ ,  $\omega$ ,  $\varphi$ ,  $\overline{f}$ ,  $\alpha$ ,  $\nu_t$ ), the following are highly important:

- The designation “scalar” refers to scalar variables which are solution of an advection equation, apart from the variables of the turbulence model ( $k$ ,  $\varepsilon$ ,  $R_{ij}$ ,  $\omega$ ,  $\varphi$ ,  $\overline{f}$ ,  $\alpha$ ,  $\nu_t$ ): for instance the temperature, scalars which may be passive or not, “user” or not. The mean value of the square of the fluctuations of a “scalar” is a “scalar”, too. The scalars may be divided into two groups: **nscaus** “user” scalars and **nscapp** “specific physics” scalars, with **nscal=nscaus+nscapp**. **nscal** must be less than or equal to **nscamx**.
- The  $j^{\text{th}}$  user scalar is, in the whole list of the **nscal** scalars, the scalar number **j**. In the list of the **nvar** solved variables, it corresponds to the variable number **isca(j)**, its value in the cell **iel** at the current time step is given by **rtp(iel,isca(j))**.
- The  $j^{\text{th}}$  scalar related to a specific physics is, in the whole list of the **nscal** scalars, the scalar number **iscapp(j)**. In the list of the **nvar** solved variables, it corresponds to the variable number **isca(iscapp(j))**, its value in the cell **iel** at the current time step is given by **rtp(iel,isca(iscapp(j)))**.
- Apart from specific physics, the temperature (or the enthalpy) is the scalar number **iscalt** in the list of the **nscal** scalars. It corresponds to the variable number **isca(iscalt)** and its value in the cell **iel** is **rtp(iel,isca(iscalt))**. if there is no thermal scalar, **iscalt** is equal to -1.
- A “user” scalar number **j** may represent the average of the square of the fluctuations of a scalar **k** (i.e. the average  $\overline{\varphi'\varphi'}$  for a fluctuating scalar  $\varphi$ ). This can be made either *via* the interface or by indicating **iscavr(j)=k** in **cs\_user\_parameters.f90** (if the scalar in question is not a “user” scalar, the selection is made automatically). For instance, if **j** and **k** are “user” scalars, the variable  $\varphi$  corresponding to **k** is the variable number **isca(k)=isca(iscavr(j))**, and its value in the cell **iel** is **rtp(iel,isca(k))=rtp(iel,isca(iscavr(j)))**.

EDF R&D	<b>Code_Saturne version 3.2.0 practical user's guide</b>	Code_Saturne documentation Page 36/202
---------	--	--

The variable corresponding to the mean value of the square of the fluctuations<sup>10</sup> is the variable number `isca(j)` and its value in the cell `iel` is `rtp(iel,isca(j))`.

#### About `propce` and `propfb`

In *Code\_Saturne*, the physical properties<sup>11</sup> are stored in the `propce` and `propfb` arrays. Some properties, like the density, are only stored for cells and boundary faces. To avoid having different index numbers for a physical property, depending on the array it is used in, the following structure is used in *Code\_Saturne*:

- All the properties (used or not) have a unique and distinct index-number, given automatically by the code and stored in an integer or an integer array (its size may be the maximum number of scalars or the maximum number of variables).
- The indexes referring to the different properties stored in the `propxx` arrays are given respectively by the following integer arrays:

`ipproc(npromx)` [ia]: Rank `i` in `propce(.,i)` of the properties defined at the cell centres.

`ipprob(npromx)` [ia]: Rank `i` in `propfb(.,i)` of the properties defined at the boundary faces.

For instance, the index number corresponding to the density is `irom`.

In the list of the properties defined at the cell center, the density is therefore the `ipproc(irom)`<sup>th</sup> property: its value at the center of the cell `iel` is given by `propce(iel,ipproc(irom))`.

In the same way, in the list of the properties defined at the boundary faces, the density is the `ipprob(irom)`<sup>th</sup> property: its value at the boundary face is given by `propfb(iel,ipprob(irom))`.

The list of properties accessible in the `propxx` arrays is given below (this does not include the properties linked to the specific physics modules):

`irom` [ia]: Property number corresponding to the density (*i.e.*  $\rho$  in  $kg.m^{-3}$ ).  
Stored at the cells and the boundary faces.

`iroma` [ia]: Property number corresponding to the density (*i.e.*  $\rho$  in  $kg.m^{-3}$ ) at the previous time step, in the case of a second-order extrapolation in time.  
Stored at the cells and the boundary faces.

`ivisc1` [ia]: Property number corresponding to the fluid molecular dynamic viscosity (*i.e.*  $\mu$  in  $kg.m^{-1}.s^{-1}$ ).  
Stored at the cells.

`ivisla` [ia]: Property number corresponding to the fluid molecular dynamic viscosity (*i.e.*  $\mu$  in  $kg.m^{-1}.s^{-1}$ ) at the previous time step, in the case of a second-order extrapolation in time.  
Stored at the cells.

`ivisct` [ia]: Property number corresponding to the fluid turbulent dynamic viscosity (*i.e.*  $\mu_t$  in  $kg.m^{-1}.s^{-1}$ ).  
Stored at the cells.

`ivista` [ia]: Property number corresponding to the fluid turbulent dynamic viscosity (*i.e.*  $\mu_t$  in  $kg.m^{-1}.s^{-1}$ ) at the previous time step, in the case of a second-order extrapolation in time.  
Stored at the cells.

`icp` [ia]: Property number corresponding to the specific heat, when it is variable (*i.e.*  $C_p$  in  $m^2.s^{-2}.K^{-1}$ ).  
See note below.  
Stored at the cells.

---

<sup>10</sup>it is really  $\overline{\varphi'\varphi'}$ , and not  $\sqrt{\varphi'\varphi'}$

<sup>11</sup>other variables are stored in the arrays `propce` and `propfb`. They are not “physical properties” strictly speaking, but it is convenient to have them in the same array as the proper physical properties

EDF R&D	<b>Code_Saturne version 3.2.0 practical user's guide</b>	Code_Saturne documentation Page 37/202
---------	--	--

- icpa** [ia]: Property number corresponding to the specific heat, when it is variable (*i.e.*  $C_p$  in  $m^2.s^{-2}.K^{-1}$ ), at the previous time step, in the case of a second-order extrapolation in time. See note below.  
Stored at the cells.
- itsnsa** [ia]: In the case of a calculation run with a second-order discretisation in time with extrapolation of the source terms, property number corresponding to the source term of Navier-Stokes at the previous time step ( $kg.m^{-1}.s^{-2}$ ).  
Stored at the cells.
- itstua** [ia]: In the case of a calculation run with a second-order discretisation in time with extrapolation of the source terms, property number corresponding to the source terms of the turbulence at the previous time step.  
Stored at the cells.
- itssca** [ia]: In the case of a calculation run with a second-order discretisation in time with extrapolation of the source terms, property number corresponding to the source terms of the equations solved for the scalars at the previous time step ( $kg.m^{-1}.s^{-2}$ ).  
Stored at the cells.
- iestim(nestmx)** [ia]: Property number for the **nestmx** error estimators for Navier-Stokes. The estimators currently available are **iestim(iespre)**, **iestim(iesder)**, **iestim(iescor)**, **iestim(iestot)** stored at the cells.
- ifluaa(nvarmx)** [ia]: Property number corresponding to the mass flow associated with each variable at the previous time step, in the case of a second-order extrapolation in time.  
Stored at the internal faces and boundary faces.
- ivisls(nscamx)** [ia]: Property number corresponding to the diffusivity of scalars for which it is variable (*i.e.* thermal conductivity  $\lambda$  for the temperature, in  $W.m^{-1}.K^{-1}$ ). It must be noticed that the diffusivity is associated with the scalars rather than with the variables. See note below.  
Stored at the cells.
- ivissa(nscamx)** [ia]: Property number corresponding to the diffusivity of scalars for which it is variable (*i.e.* thermal conductivity  $\lambda$  for the temperature, in  $W.m^{-1}.K^{-1}$ ) at the previous time step, in the case of a second-order extrapolation in time.  
Stored at the cells.
- ismago** [i]: Property number corresponding to the variable  $C$  of the dynamic model, *i.e* so that  $\mu_t = \rho C \bar{\Delta}^2 \sqrt{2S_{ij}S_{ij}}$  (with the notations of [3]).  $C$  corresponds to  $C_s^2$  in the classical model of Smagorinsky.  
Stored at the cells.
- icour** [i]: CFL number in each cell at the present time step.  
Stored at the cells.
- ifour** [i]: Fourier number in each cell at the present time step.  
Stored at the cells.
- ipttot** [i]: Total pressure in each cell.  
Stored at the cells.
- ivisma(1 or 3)** [ia]: When the ALE method for deformable meshes is activated, **ivisma** corresponds to the “mesh viscosity”, allowing to limit the deformation in certain areas. This mesh viscosity can be isotropic or be taken as a diagonal tensor (depending on the value of the parameter **iortvm**).  
Stored at the cells.
- icmome(nbmomx)** [ia]: Property number corresponding to the time averages defined by the user. More precisely, it is not the time average that is stored, but a summation over time (the division

by the cumulated duration is done just before the results are written).  
Stored at the cells.

`icdtmo(nbmomx)` [ia]: Property number corresponding to the cumulated duration associated with each time average defined by the user, when this duration is not spatially uniform (see note below).  
Stored at the cells.

#### NOTE: VARIABLE PHYSICAL PROPERTIES

Some physical properties such as specific heat or diffusivity are often constant (choice made by the user). In that case, in order to limit the necessary memory, these properties are stored as a simple real number rather than in a domain-sized array of reals.

- It is the case for the specific heat  $C_p$ .
  - If  $C_p$  is constant, it can be specified in the interface or by indicating `icp=0` in `cs.user_parameters.f90`, and the property will be stored in the real number `cp0`.
  - If  $C_p$  is variable, it can be specified in the interface or by indicating `icp=1` in `cs.user_parameters.f90`. The code will then modify this value to make `icp` refer to the effective property number corresponding to the specific heat, in a way which is transparent for the user. For each cell `iel`, the value of  $C_p$  is then given in `usphyv` and stored in the array `propce(iel,iproc(icp))`.
- It is the same for the diffusivity  $K$  of each scalar `iscal`.
  - If  $k$  is constant, it can be specified in the interface or by indicating `ivisls(iscal)=0` in `cs.user_parameters.f90`, and the property will be stored in the real number `visls0(iscal)`.
  - If  $k$  is variable, it can be specified in the interface or by indicating `ivisls(iscal)=1` in `cs.user_parameters.f90`. The code will then modify this value to make `ivisls(iscal)` refer to the effective property number corresponding to the diffusivity of the scalar `iscal`, in a way which is transparent for the user. For each cell `iel`, the value of  $k$  is then given in `usphyv` and stored in the `propce(iel,iproc(ivisls(iscal)))` array.

#### NOTE: CUMULATED DURATION ASSOCIATED WITH THE AVERAGES DEFINED BY THE USER

The cumulated duration associated with the calculation of a time averages defined by the user is often a spatially uniform value. In this case, it is stored in a simple real number: for the mean value `imom`, it is the real number `dtcmom(-idtmom(imom))` (`idtmom(imom)` is negative in this case).

When this cumulated duration is not spatially uniform (for instance in the case of a spatially variable time step), it is stored in `propce`. It must be noted that the cumulated duration associated with the calculation of the average `imom` is variable in space if `idtmom(imom)` is strictly positive. The number of the associated property in `propce` is then `icdtmo(idtmom(imom))`. For instance, for the average `imom`, the cumulated duration in the cell `iel` will be `propce(iel,iproc(icdtmo(idtmom(imom))))`. The user may have a look to the examples given for `cs.user_extra_operations` to know how to calculate a time averages in a particular cases (printing of extreme values, writing of results, ...).

Two other variables, `hbord` and `tbord`, should be noted here, although they are relatively local (they appear only in the treatment of the boundary conditions) and are used only by developers.

`hbord(nfabor)` [ra]: Array of the exchange coefficient for temperature (or enthalpy) at the boundary faces. The table is allocated only if `isvhb` is set to 1 in the subroutine `tridim` (which is note a user subroutine), which is done automatically, but only if the coupling with SYRTHES or the 1D thermal wall module are activated..

EDF R&D	<b><i>Code_Saturne</i> version 3.2.0 practical user's guide</b>	<i>Code_Saturne</i> documentation Page 39/202
---------	---	---

**tbord**(nfabor) [ra]: Temperature (or enthalpy) at the boundary faces<sup>12</sup>. The table is allocated only if **isvtb** is set to 1 in the subroutine **tridim** (which is not a user subroutine), which is done automatically but only if the coupling with SYRTHES or the 1D thermal wall module are activated..

Tables **hbord** and **tbord** are of size **nfabor**, although they concern only the wall boundary faces.

### 3.9.3.4 Variables related to the numerical methods

The main numerical variables and “pointers” are displayed below.

#### BOUNDARY CONDITIONS

**coefa**(nfabor,\*) [ra]: Boundary conditions: see note 2.

**coefau**(3,nfabor) [ra]: Boundary conditions for the velocity vector with the coupled velocity components algorithm (**ivelco**=1): see note 2.

**cofafu**(3,nfabor) [ra]: Boundary conditions for the velocity diffusion flux with the coupled velocity components algorithm (**ivelco**=1): see note 2.

**coefb**(nfabor,\*) [ra]: Boundary conditions: see note 2.

**coefbu**(3,nfabor) [ra]: Boundary conditions for the velocity vector with the coupled velocity components algorithm (**ivelco**=1): see note 2.

**cofbfu**(3,nfabor) [ra]: Boundary conditions for the velocity diffusion flux with the coupled velocity components algorithm (**ivelco**=1): see note 2.

**iclrrtp**(nvarmx,3) [ia]: For each variable **ivar** ( $1 \leq \text{ivar} \leq \text{nvar} \leq \text{nvarmx}$ ), rank in **coefa** and **coefb** of the boundary conditions. See note 2.

**icoef** [i]: Rank in **iclrrtp** of the rank in **coefa** and **coefb** of the gradient boundary conditions. See note 2.

**icoeff** [i]: Rank in **iclrrtp** of the rank in **coefa** and **coefb** of the diffusive flux type boundary conditions. See note 2.

**icoefr** [i]: Rank in **iclrrtp** of the rank in **coefa** and **coefb** of the divergence type boundary conditions (especially for Reynolds stresses in the momentum equation). See note 2.

**ifmfbr**(nfabor) [ia]: Family number of the boundary faces. See note 1.

**iprfml**(nfml,nprfml) [ia]: Properties of the families of referenced entities. See note 1.

**isympha**(nfabor) [i]: Integer to mark out the “symmetry” (**itypfb**=**isymet**) boundary faces where the mass flow has to be canceled when the ALE module is switched off (these faces are impermeable). For instance, if the face **ifac** is symmetry face, **isympha**(**ifac**)=0, otherwise **isympha**(**ifac**)=1.

**itrifb**(nfabor) [ia]: Indirection array allowing to sort the boundary faces according to their boundary condition type **itypfb**.

**itypfb**(nfabor) [ia]: Boundary condition type at the boundary face **ifac** (see user subroutine **cs\_user\_boundary\_conditions**).

**uetbor**(nfabor) [ra]: Friction velocity at the wall, in the case of a LES calculation with van Driest-wall damping.

---

<sup>12</sup>It is the physical temperature at the boundary faces, not the boundary condition for temperature. See [11] for more details on boundary conditions

EDF R&D	<b>Code_Saturne version 3.2.0 practical user's guide</b>	Code_Saturne documentation Page 40/202
---------	--	--

### DISTANCE TO THE WALL

**ifapat**(ncelet) [ra]: Number of the wall face (type **itypfb**=**iparoi** or **iparug**) which is closest to the center of a given volume when necessary ( $R_{ij} - \varepsilon$  with wall echo, LES with van Driest-wall damping, or  $k - \omega$  (SST) turbulence model) and when **icdpar**=2. The number of the wall face which is the closest to the center of the cell **iel** is **ifapat**(**iel1**). This calculation method is not compatible with parallelism and periodicity.

**dispar**(ncelet) [ra]: Distance between the center of a given volume and the closest wall, when it is necessary ( $R_{ij} - \varepsilon$  with wall echo, LES with van Driest-wall damping, or  $k - \omega$  (SST) turbulence model) and when **icdpar**=1. The distance between the center of the cell **iel** and the closest wall is **dispar**(**iel**).

**yplpar** [ra]: Non-dimensional distance  $y^+$  between a given volume and the closest wall, when it is necessary (LES with van Driest-wall damping) and when **icdpar**=1. The adimensional distance  $y^+$  between the center of the cell **iel** and the closest wall is therefore **yplpar**(**iel1**).

### PRESSURE DROPS AND POROSITY

**icepdc**(ncepdc) [ia]: Number of the **ncepdc** cells in which a pressure drop is imposed. See **iicepd** and the user subroutine **uskpdc**.

**ckupdc**(ncepdc,6) [ra]: Value of the coefficients of the pressure drop tensor of the **ncepdc** cells in which a pressure drop is imposed. Note the 6 values are sorted as follows: (k11, k22, k33, k12, k23, k33). See **ickpdc** and the user subroutine **uskpdc**.

**ncepdc** [ia]: Number of cells in which a pressure drop is imposed. See the user subroutine **uskpdc**.

**porosi**(ncelet) [ra]: Value of the porosity.

### MASS SOURCES

**icetsm**(ncetsm) [ia]: Number of the **ncetsm** cells in which a mass source term is imposed. See **iicesm** and the user subroutine **ustsma**.

**itypsm**(ncetsm,nvar) [ia]: Type of mass source term for each variable (0 for an injection at ambient value, 1 for an injection at imposed value). See the user subroutine **ustsma**.

**ncetsm** [i]: Number of cells with mass sources. See the user subroutine **ustsma**.

**smacel**(ncetsm,nvar) [ra]: Value of the mass source term for pressure. For the other variables, eventual imposed injection value. See the user subroutine **ustsma**.

### WALL 1D THERMAL MODULE

**nfpt1d** [i]: Number of boundary faces which are coupled with a wall 1D thermal module. See the user subroutine **uspt1d**.

**ifpt1d** [ia]: Array allowing to mark out the numbers of the **nfpt1d** boundary faces which are coupled with a wall 1D thermal module. The numbers of these boundary faces are given by **ifpt1d**(**ii**), with  $1 \leq ii \leq nfpt1d$ . See the user subroutine **uspt1d**.

**nppt1d** [ia]: Number of discretisation cells in the 1D wall for the **nfpt1d** boundary faces which are coupled with a wall 1D thermal module. The number of cells for these boundary faces is given by **nppt1d**(**ii**), with  $1 \leq ii \leq nfpt1d$ . See the user subroutine **uspt1d**.

**eppt1d** [ia]: Thickness of the 1D wall for the **nfpt1d** boundary faces which are coupled with a



EDF R&D	<b>Code_Saturne version 3.2.0 practical user's guide</b>	Code_Saturne documentation Page 41/202
---------	--	--

wall 1D thermal module. The wall thickness for these boundary faces is therefore given by `eppt1d(ii)`, with  $1 \leq ii \leq \text{nfpt1d}$ . See the user subroutine `uspt1d`.

#### OTHERS

`dt(ncelet)` [ra]: Value of the time step.

`ifmcel(ncelet)` [ia]: Family number of the elements. See note 1.

`s2kw(ncelet)` [ra]: Square of the norm of the deviatoric part of the deformation rate tensor ( $S^2 = 2S_{ij}^D S_{ij}^D$ ). This array is defined only with the  $k - \omega$  (SST) turbulence model.

`divukw` [ia]: Divergence of the velocity. More precisely it is the trace of the velocity gradient (and not a finite volume divergence term). In the cell `iel`,  $\text{div}(\underline{u})$  is given by `divukw(iel1)`. This array is defined only with the  $k - \omega$  SST turbulence model (because in this case it may be calculated at the same time as  $S^2$ ).

`ra(ifinra)` [ra]: Real work array containing `dt`, `rtp`, `rtpa`, `propce`, `tpucou`.

#### NOTE: BOUNDARY CONDITIONS

The **gradient** boundary conditions in *Code\_Saturne* boil down to determine a value for the current variable  $Y$  at the boundary faces  $f_b$ , that is to say  $Y_{f_b}$ , value expressed as a function of  $Y_{I'}$ , value of  $Y$  in  $I'$ , projection of the center of the adjacent cell on the straight line perpendicular to the boundary face and crossing its center:

$$Y_{f_b} = A_{f_b}^g + B_{f_b}^g Y_{I'}. \quad (1)$$

For a face `ifac`, the pair of coefficients  $A_{f_b}^g$ ,  $B_{f_b}^g$  is stored in `coefa(ifac,iclvar)` and `coefb(ifac,iclvar)`, where the integer `iclvar=iclrtp(ivar,icoef)` determines the rank in `coefa` and `coefb` of the set of boundary gradient conditions of the variable `ivar`.

The **flux** boundary conditions in *Code\_Saturne* boil down to determine the value of the diffusive flux of the current variable  $Y$  at the boundary faces  $f_b$ , that is to say  $D_{ib}(K_{f_b}, Y)$ , value expressed as a function of  $Y_{I'}$ , value of  $Y$  in  $I'$ , projection of the center of the adjacent cell on the straight line perpendicular to the boundary face and crossing its center:

$$D_{ib}(K_{f_b}, Y) = A_{f_b}^f + B_{f_b}^f Y_{I'}. \quad (2)$$

For a face `ifac`, the pair of coefficients  $A_{f_b}^f$ ,  $B_{f_b}^f$  is stored in `coefa(ifac,iclvarf)` and `coefb(ifac,iclvarf)`, where the integer `iclvarf=iclrtp(ivar,icoeff)` determines the rank in `coefa` and `coefb` of the set of boundary diffusive flux conditions of the variable `ivar`.

The **divergence** boundary conditions in *Code\_Saturne* boil down to determine a value for the current variable  $Y$  (mainly the Reynolds stress components, the divergence  $\underline{\text{div}}(\underline{R})$  used in the calculation of the momentum equation) at the boundary faces  $f_b$ , that is to say  $Y_{f_b}$ , value expressed as a function of  $Y_{I'}$ , value of  $Y$  in  $I'$ , projection of the center of the adjacent cell on the straight line perpendicular to the boundary face and crossing its center:

$$Y_{f_b} = A_{f_b}^d + B_{f_b}^d Y_{I'}. \quad (3)$$

For a face `ifac`, the pair of coefficients  $A_{f_b}^d$ ,  $B_{f_b}^d$  is stored in `coefa(ifac,iclvrr)` and `coefb(ifac,iclvrr)`, where the integer `iclvrr=iclrtp(ivar,icoefr)` determines the rank in `coefa` and `coefb` of the set of boundary divergence conditions of the variable `ivar`.



### 3.9.3.5 User arrays

Modules containing user arrays accessible from all user subroutines may be defined in the `user_modules.f90` file. This file is compiled before any other Fortran user file, to ensure modules may be accessed in other user subroutines using the `use <module>` construct. It may contain any routines or variables the user needs, and contains no predefined routines or variables (i.e. the only specificity of this file is that a file with this name is compiled before all others).

### 3.9.3.6 Parallelism and periodicity

Parallelism is based on domain partitioning: each processor is assigned a part of the domain, and data for cells on parallel boundaries is duplicated on neighbouring processors in corresponding “ghost”, or “halo” cells (both terms are used interchangeably). Values in these cells may be accessed just the same as values in regular cells. Communication is only required when cell values are modified using values from neighbouring cells, as the values in the “halo” can not be computed correctly (since the halo does not have access to all its neighbours), so halo values must be updated by copying values from the corresponding cells on the neighbouring processor.

Compared to other tools using a similar system, a specificity of *Code\_Saturne* is the separation of the halo in two parts: a standard part, containing cells shared through faces on parallel boundaries, and an extended part, containing cells shared through vertices, which is used mainly for least squares gradient reconstruction using an extended neighbourhood. Most updates need only to operate on the standard halo, requiring less data communication than those on the extended halos.



Figure 4: Parallel domain partitioning: halos

Periodicity is handled using the same halo structures as parallelism, with an additional treatment for vector and coordinate values: updating coordinates requires applying the periodic transformation to the copied values, and in the case of rotation, updating vector and tensor values also requires applying the rotation transformation. Ghost cells may be parallel, periodic, or both. The example of a pump combining parallelism and periodicity is given Figure 5. In this example, all periodic boundaries match with boundaries on the same domain, so halos are either parallel or periodic.

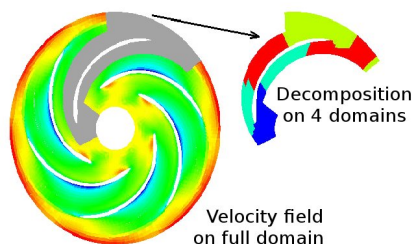


Figure 5: Combined parallelism and periodicity

### Activation

Parallelism is activated by means of the GUI or of the launch scripts in the standard cases:

- On clusters with batch systems, the launching of a parallel run requires to complete the batch cards located in the beginning of `runcase` or `runcase_batch` script, and set the number of MPI processes, or the numbers of physical nodes and processors per node (`ppn`) wanted. This can be done through the Graphical Interface or by editing the `runcase` or `runcase_batch` file directly. The number of processors defined here will override the number defined through the GUI in a non-batch environment (so that studies defined on one environment may be migrated to larger compute resources easily), but it may be overridden by the `define_case_parameters` function from the `cs_user_scripts.py` file, or by setting the `n_procs_weight`, `n_procs_min`, and `n_procs_max` parameters for the different domains defined in `runcase_coupling`.
- On clusters with unsupported batch systems, `runcase` file may have to be modified manually. Please do not hesitate to contact the *Code\_Saturne* support ([saturne-support@edf.fr](mailto:saturne-support@edf.fr)) so that these modifications can be added to the standard launch script to make it more general.
- A parallel calculation may be stopped in the same manner as a sequential one using the file `control_file` (see paragraph 3.2.5).
- The standard elements of information displayed in the listing (marked out with 'v' for the min/max values of the variables), 'c' for the data concerning the convergence and 'a' for the values before clipping) are global values for the whole domain and not related to each processor.

## User subroutines

The user can check in a subroutine

- that the presence of periodicity is tested with the variable `iperio` (=1 if periodicity is activated);
- that the presence of rotation periodicities is tested with the variable `iperot` (number of rotation periodicities);
- that running of a calculation in parallel is tested for with the variable `irangp` (`irangp` is worth -1 in the case of a non-parallel calculation and  $p - 1$  in the case of a parallel calculation,  $p$  being the number of the current processor)

Attention must be paid to the coding of the user subroutines. If conventional subroutines like `cs_user_parameters.f90` or `cs_user_boundary_conditions` usually do not cause any problem, some kind of developments are more complicated. The most usual cases are dealt with below.

Examples are given for the subroutine `cs_user_extra_operations`.

- **Access to information related to neighbouring cells in parallel and periodic cases.**  
When periodicity or parallelism are brought into use, some cells of the mesh become physically distant from their neighbours. Concerning parallelism, the calculation domain is split and distributed between the processors: a cell located at the “boundary” of a given processor may have neighbours on different processors.  
In the same way, in case of periodicity, the neighbouring cells of cells adjacent to a periodic face are generally distant.  
When data concerning neighbouring cells are required for the calculation, they must first be searched on the other processors or on the other edge of periodic frontiers. In order to ease the manipulation of these data, they are stored temporarily in virtual cells called “halo” cells, as can be seen in Figure 4. It is in particular the case when the following operations are made on a variable  $A$ :
  - calculation of the gradient of  $A$  (use of the subroutine `grdcel`);
  - calculation of an internal face value from the values of  $A$  in the neighbouring cells (use of `ifacel`).

The variable *A* must be exchanged before these operations can be made: to allow it, the subroutine `synsca` may be called.

- **Global operations in parallel mode.**

In parallel mode, the user must pay attention during the realisation of global operations. The following list is not exhaustive:

- calculation of extreme values on the domain (for instance, minimum and maximum of some calculation values);
- test of the existence of a certain value (for instance, do faces of a certain color exist?);
- verification of a condition on the domain (for instance, is a given flow value reached somewhere?);
- counting out of entities (for instance, how many cells have pressure drops?);
- global sum (for instance, calculation of a mass flow or the total mass of a pollutant).

The user may refer to the different examples present in the directory **EXAMPLES** in the `cs_user_extra_operations-parallel_operations.f90` file. Care should be taken with the fact that the boundaries between subdomains consist of **internal** faces shared between two processors (these are indeed internal faces, even if they are located at a “processor boundary”). They should not be counted twice (once per processor) during global operations using internal faces (for instance, counting the internal faces per processor and summing all the obtained numbers drives into over-evaluating the number of internal faces of the initial mesh).

- **Writing operations that should be made on one processor only in parallel mode.**

In parallel mode, the user must pay attention during the writing of pieces of information. Writing to the “listing” can be done simply by using the `nfecra` logical unit (each processor will write to its own “listing” file): use `write(nfecra,...)`.

If the user wants an operation to be done by only one processor (for example, open or write a file), the associated instructions must be included inside a test on the value of `irangp` (generally it is the processor 0 which realises these actions, and we want the subroutine to work in non-parallel mode, too: `if (irangp.le.0) then ...`).

### Some notes about periodicity

Note that periodic faces are not part of the domain boundary: periodicity is interpreted as a “geometric” condition rather than a classical boundary condition.

Some particular points should be reminded:

- Periodicity can also work when the periodic boundaries are meshed differently (periodicity of non-conforming faces), *except* for the case of a 180 degree rotation periodicity with faces coupled on the rotation axis.
- rotation periodicity is incompatible with
  - semi-transparent radiation,
  - reinforced velocity-pressure coupling (`ipucou=1`).
- although it has not been the case so far, potential problems might be met in the case of rotation periodicity with the  $R_{ij} - \varepsilon$  (LRR) model. They would come from the way of taking into account the orthotropic viscosity (however, this term usually has a low influence).

### 3.9.3.7 Geometry and particle arrays related to Lagrangian modelling

In this section is given a non-exhaustive list of the main variables which may be seen by the user in the Lagrangian module. Most of them should not be modified by the user. They are calculated automatically from the data. However it may be useful to know their meaning.

EDF R&D	<b><i>Code_Saturne</i> version 3.2.0 practical user's guide</b>	<i>Code_Saturne</i> documentation Page 45/202
---------	---	---

These variables are listed in the alphabetical index in the end of this document.

The type of each variable is given: integer [i], real number [r], integer array [ia], real array [ra].

#### SIZE OF THE LAGRANGIAN ARRAYS

**lndnod** [i]: Size of the array **icocel** concerning the cells  $\rightarrow$  faces connectivity (the faces  $\rightarrow$  nodes connectivity must be given to allow the construction of this connectivity. See note 3 of section 3.9.3.1).

**nbpmax** [i]: Maximum number of particles simultaneously allowed in the calculation domain.

**nvp** [i]: Number of variables describing the particles for which a stochastic differential equation (SDE) is solved.

**nvls** [i]: Number of variables describing the supplementary user particles for which a SDE is solved.

**nvep** [i]: Number of real state variables describing the particles.

**nivep** [i]: Number of integer state variables describing the particles.

**ntersl** [i]: Number of source terms representing the backward coupling of the dispersed phase on the continuous phase.

**nvlst** [i]: Number of volumetric statistical variables .

**nvlst** [i]: Number of supplementary user volumetric statistical variables.

**nvisbr** [i]: Number of boundary statistical variables.

**nusbor** [i]: Number of supplementary user boundary statistical variables.

**nvgaus** [i]: Number of Gaussian random variables.

#### LAGRANGIAN ARRAYS

**icocel(lndnod)** [ia]: Cells *rightarrow* internal/boundary faces connectivity. The numbers of the boundary faces are marked out in **icocel** with a negative sign.

**itycel(ncelet+1)** [ia]: Array containing the position of the first face surrounding every cell in the array **icocel** (see subroutine **lagdeb** for more details).

**ettp(nbpmax,nvp)** [ra]: Variables forming the state vector related to the particles: either at the current stage if the Lagrangian scheme is a second-order, or at the current time step if the scheme is a first-order. These variables are marked out by “pointers” whose value can vary between 1 and **nvp**:

$\rightarrow$  **jmp**: particle mass

$\rightarrow$  **jdp**: particle diameter

$\rightarrow$  **jxp, jyp, jzp**: particle coordinates

$\rightarrow$  **jup, jvp, jwp**: particle velocity components

$\rightarrow$  **juf, jvf, jwf**: locally undisturbed fluid flow velocity components

$\rightarrow$  **jtp, jtf**: particle and locally undisturbed fluid flow temperature ( $^{\circ}\text{C}$ )

$\rightarrow$  **jcp**: particle specific heat

$\rightarrow$  **jhp**: coal particle temperature ( $^{\circ}\text{C}$ )

- **jmch**: mass of reactive coal of the coal particle
- **jmck**: mass of coke of the coal particle
- **jvls(ii)**: **ii<sup>th</sup>** supplementary user variable

**ettpa(nbpmax,nvp)** [ra]: Variables forming the state vector related to the particles: either at the previous stage if the Lagrangian scheme is a second-order, or at the previous time step if the Lagrangian scheme is a first-order.

**itepa(nbpmax,nivep)** [ia]: Integer variables related to the particles. They are marked out by the following “pointers”:

- **jisor**: Number of the current cell containing the particle; this number is re-actualised during the trajectography step
- **jinch**: Number of the coal particle

**tepa(nbpmax,nvep)** [ra]: Real variables related to the particles. They are marked out by the following “pointers”:

- **jrtsp**: particle residence time
- **jrpoi**: particle statistic weight
- **jrdck**: coal particle shrinking core diameter
- **jrd0p**: coal particle initial diameter
- **jrr0p**: coal particle initial density

**indep(nbpmax)** [ia]: Storage of the cell number of every particle at the beginning of a Lagrangian iteration; this data is not modified during the iteration.

**vitpar(nbpmax,3)** [ra]: At the beginning of the trajectography, **vitpar** contains the particle velocity vector components; the modifications of the particle velocity following every particle/boundary interaction are saved in this array; after the trajectography and backward coupling steps, **ettp** is updated with **vitpar**.

**vitflu(nbpmax,3)** [ra]: At the beginning of the trajectography, **vitflu** contains the locally undisturbed fluid flow velocity vector components; the modifications of the locally undisturbed fluid flow velocity following every particle/boundary interaction are saved in this array; after the trajectography and backward coupling steps, **ettp** is updated with **vitflu**.

**gradpr(ncelet,3)** [ra]: Pressure gradient of the continuous phase.

**gradvf(ncelet,9)** [ra]: Gradient of the continuous phase fluid velocity (useful if the complete model is activated: see **modcpl**).

**cpgd1(nbpmax)** [ra]: First de-volatilisation term (light volatile matters) of the coal particles (useful in the case of backward coupling on the continuous phase).

**cpgd2(nbpmax)** [ra]: Second de-volatilisation term (heavy volatile matters) of the coal particles (useful in the case of backward coupling on the continuous phase).

**cpght(nbpmax)** [ra]: Heterogeneous combustion term of the coal particles (useful in the case of backward coupling on the continuous phase).

**statis(ncelet,nvlsta)** [ra]: Averages of the volumetric variables related to the dispersed phase  
They can be accessed by the following “pointers”:

- **ilvx,ilvy,ilvz**: mean dispersed phase velocity
- **ilfv**: dispersed phase volumetric concentration
- **ilpd**: sum of the statistical weights
- **iltp**: dispersed phase temperature (°C)
- **ildp**: dispersed phase mean diameter
- **ilmp**: dispersed phase mean mass
- **ilhp**: temperature of the coal particle cloud (°C)
- **ilmch**: mass of reactive coal of the coal particle cloud
- **ilmck**: mass of coke of the coal particle cloud
- **ilmck**: shrinking core diameter of the coal particle cloud
- **ilvu(ii)**: *iith* supplementary user volumetric statistics

**stativ(ncelet,nvlsta)** [ra]: Variances of the volumetric variables related to the dispersed phase.  
they can be accessed by using the same “pointers” as the ones used for the **stativ** array.

**parbor(nfavor,nvisbr)** [ra]: Boundary statistics related the dispersed phase; after every particle/boundary interaction it is possible to save some data and to calculate averages; the boundary statistics are marked out by the following “pointers”:

- **inbr**: number of particle/boundary interactions
- **iflm**: particle mass flow at the boundary faces
- **iang**: mean interaction angle with the boundary faces
- **ivit**: mean interaction velocity with the boundary faces
- **ienc**: mass of coal deposit at the walls
- **iusb(ii)**: *iith* supplementary user boundary statistics

**tslagr(ncelet,ntersl)** [ra]: Source terms corresponding to the backward coupling of the dispersed phase on the continuous phase. These source terms are marked out by the following “pointers”:

- **itsvx, itsvy, itsvz**: explicit source terms for the continuous phase velocity
- **itsli**: implicit source term for the continuous phase velocity and for the turbulent energy if the  $k - \varepsilon$  model is used
- **itske**: explicit source term for the turbulent dissipation and the turbulent energy if the  $k - \varepsilon$  turbulence model is used for the continuous phase
- **itsr11,... itsr33**: source terms for the Reynolds stress and the turbulent dissipation if the  $R_{ij} - \varepsilon$  turbulence model is used for the continuous phase

- `itsmas`: mass source term
- `itste`, `itsti`: explicit and implicit thermal source terms for the thermal scalar of the continuous phase
- `itsmv1(icha)`, `itsmv2(icha)`: source terms respectively for the light and heavy volatile matters
- `itsco`: source term for the carbon released during heterogeneous combustion

`croule(ncelet)` [ra]: Importance function for the technique of variance reduction (cloning/fusion of particles).

`vagaus(nbpmax,nvgaus)` [ra]: Vectors of Gaussian random variables.

`auxl(nbpmax,3)` [ra]: Auxiliary work array.

### 3.9.3.8 Variables saved to allow calculation restarts

The directory `checkpoint` contains:

- `main`: main restart file,
- `auxiliary`: auxiliary restart file (see `ileaux`, `iecaux`),
- `radiative_transfer`: restart file for the radiation module,
- `lagrangian`: main restart file for the Lagrangian module,
- `lagrangian_stats`: auxiliary restart file for the Lagrangian module (mainly for the statistics),
- `1dwall_module`: restart file for the 1D wall thermal module,
- `vortex`: restart file for the vortex method (see `ivrtex`).

The main restart file contains the values in every cell of the mesh for pressure, velocity, turbulence variables and all the scalars (user scalars et specific physics scalars. Its content is sufficient for a calculation restart, but the complete continuity of the solution at restart is not ensured<sup>13</sup>.

The auxiliary restart file completes the main restart file to ensure solution continuity in the case of a calculation restart. If the code cannot find one or several pieces of data required for the calculation restart in the auxiliary restart file, default values are then used. This allows in particular to run calculation restarts even if the number of faces has been modified (for instance in case of modification of the mesh merging or of periodicity conditions<sup>14</sup>). More precisely, the auxiliary restart file contains the following data:

- type and value of the time step, turbulence model,
- density value at the cells and boundary faces, if it is variable,
- values at the cells of the other variable physical properties, when they are extrapolated in time (molecular dynamic viscosity, turbulent or sub-grid scale viscosity, specific heat, scalar diffusivity); for the Joule effect, the specific heat is stored automatically (in case the user should need it at restart to calculate the temperature from the enthalpy before the new specific heat has been estimated),

<sup>13</sup>in other words, a restart calculation of  $n$  time steps following a calculation of  $m$  time steps will not yield strictly the same results as a direct calculation on  $m+n$  time steps, whereas it is the case when the auxiliary file is used

<sup>14</sup>imposing a periodicity changes boundary faces into internal faces



- time step value at the cells, if it is variable,
- mass flow value at the internal and boundary faces (at the last time step, and also at the previous time step if required by the time scheme),
- boundary conditions,
- values at the cells of the source terms when they are extrapolated in time,
- number of time-averages, and values at the cells of the associated cumulated values,
- for each cell, distance to the wall when it is required (and index-number of the nearest boundary face, depending on `icdpar`),
- values at the cells of the external forces in balance with a part of the pressure (hydrostatic, in general),
- for the D3P gas combustion model: massic enthalpies and temperatures at entry, type of boundary zones and entry indicators,
- for the EBU gas combustion model: temperature of the fresh gas, constant mixing rate (for the models without mixing rate transport), types of boundary zones, entry indicators, temperatures and mixing rates at entry,
- for the LWC gas combustion model: the boundaries of the probability density functions for enthalpy and mixing rate, types of boundary zones, entry indicators, temperatures and mixing rates at entry,
- for the pulverised coal combustion: coal density, types of boundary zones, variables `ientat`, `ientcp`, `inmoxy`, `timpat`, `x20` (in case of coupling with the Lagrangian module, `iencp` and `x20` are not saved),
- for the pulverised fuel combustion: types of boundary zones, variables `ientat`, `ientfl`, `inmoxy`, `timpat`, `qimpat`, `qimpfl`,
- for the electric module: the tuned potential difference `dpot` and, for the electric arcs module, the tuning coefficient `coejou` (when the boundary conditions are tuned), the Joule source term for the enthalpy (with the Joule effect is activated) and the Laplace forces (with the electric arc module).

It should be noted that, if the auxiliary restart file is read, it is possible to run calculation restarts with relaxation of the density<sup>15</sup>(when it is variable), because this variable is stored in the restart file. On the other hand, it is generally not possible to do the same with the other physical properties (they are stored in the restart file only when they are extrapolated in time, or with the Joule effect for the specific heat).

Apart from `vortex` which has a different structure and is always in text format, all the restart files are binary files. Nonetheless, they may be dumped or compared using the `cs_io_dump` tool.

In the case of parallel calculations, it should be noted that all the processors will write their restart data in the same files. Hence, for instance, there will always be one and only one `main` file, whatever the number of processors used. The data in the file are written according to the initial full domain ids for the cells, faces and nodes. This allows in particular to restart using  $p$  processors a calculation begun with  $n$  processors, or to make the restart files independent of any mesh renumbering that may be carried out in each domain.

**WARNING:** *if the mesh is composed of several files, the order in which they appear in the launch script or in the Graphical Interface must not be modified in case of a calculation restart*<sup>16</sup>.

<sup>15</sup>such a relaxation only makes sense for a stationary calculation

<sup>16</sup>when uncertain, the user can check the saved copy of the launch script in the `RESU` directory, or the head of the `preprocessor*.log` files, which repeat the command lines passed to the Preprocessor module



EDF R&D	<b>Code_Saturne version 3.2.0 practical user's guide</b>	Code_Saturne documentation Page 50/202
---------	--	--

*NOTE: when joining of faces or periodicity is used, two nodes closer than a certain (small) tolerance will be merged. Hence, due to numerical round-up errors, two different machines may yield different results. This might change the number of faces in the global domain<sup>17</sup> and make restart files incompatible. Should that problem arise when making a calculation restart on a different architecture, the solution is to ignore the **auxiliary** file and use only the **main** file, by setting **ileaux** = 0 in **cs\_user\_parameters.f90***

### 3.9.4 Using selection criteria in user subroutines

In order to use selection criteria (cf. §3.10) in Fortran user subroutines, a collection of utility subroutines is provided. The aim is to define a subset of the mesh, for example:

- boundary regions (cf. **cs\_user\_boundary\_conditions**, **usalcl**, **usray2**, **uslag2**,...),
- volume initialization (cf. **cs\_user\_initialization**,...),
- head-loss region (cf. **uskpdc**),
- source terms region (cf. **cs\_user\_source\_terms**),
- advanced post-processing (cf. **cs\_user\_postprocess.c**, **cs\_user\_extra\_operations**, ...),

This section explains how to define surface or volume sections, in the form of lists **lstelt** of **nlelt** elements (internal faces, boundary faces or cells). For each type of element, the user calls the appropriate Fortran subroutine: **getfbr** for boundary faces, **getfac** for internal faces and **getcel** for cells. All of these take the three following arguments:

- the character string which contains the selection criterion (see some examples below),
- the returned number of elements **nlelt**,
- the returned list of elements **lstelt**.

Several examples of possible selections are given here:

- call **getfbr('Face\_1, Face\_2', nlelt, lstelt)** to select boundary faces in groups Face\_1 or Face\_2,
- call **getfac('4', nlelt, lstelt)** to select internal faces of color 4,
- call **getfac('not(4)', nlelt, lstelt)** to select internal faces which have a different color than 4,
- call **getfac('4 to 8', nlelt, lstelt)** to internal faces with color between 4 and 8 internal faces,
- call **getcel('1 or 2', nlelt, lstelt)** to select cells with colors 1 or 2,
- call **getfbr('1 and y > 0', nlelt, lstelt)** to select boundary faces of color 1 which have the coordinate  $Y > 0$ ,
- call **getfac('normal[1, 0, 0, 0.0001]', nlelt, lstelt)** to select internal faces which have a normal direction to the vector (1,0,0),
- call **getcel('all[]', nlelt, lstelt)** to select all cells.

<sup>17</sup>the number of cells will not be modified, it is always the sum of the number of cells of the different meshes

The user may then use a loop on the selected elements.

For instance, in the subroutine `cs_user_boundary_conditions` used to impose boundary conditions, let us consider the boundary faces of color number 2 and which have the coordinate  $X \leq 0.01$  (so that call `getfbr('2 and x <= 0.01', nlelt, lstelt)`); we can do a loop (`do ilelt = 1, nlelt`) and obtain `ifac = lstelt(ilelt)`.

#### NOTE: LEGACY METHOD USING EXPLICIT FAMILIES AND PROPERTIES

The selection method for user subroutines by prior versions of *Code\_Saturne* is still available, though it may be removed in future versions. This method was better adapted to working with colors than with groups, and is explained here:

From *Code\_Saturne*'s point of view, all the references to mesh entities (boundary faces and volume elements) correspond to a number (color number or negative of group number) associated with the entity. An entity may have several references (for instance, one entity may have one color and belong to several groups). In *Code\_Saturne*, these references may be designated as "properties".

The mesh entities are gathered in equivalence classes on the base of their properties. These equivalence classes are called "families". All the entities of one family have the same properties. In order to know the properties (in particular the color) of an entity (a boundary face for example), the user must first determine the family to which it belongs.

For instance, let's consider a mesh whose boundary faces have all been given one color (for example using SIMAIL). The family of the boundary face `ifac` is `ifml=ifmfbr(ifac)`. The first (and only) property of this family is the color `icoul`, obtained for the face `ifac` with `icoul=iprfml(ifml,1)`. In order to know the property number corresponding to a group, the utility function `numgrp(nomgrp, lngnom)` (with a name `nomgrp` of the type `character*` and its lenght `lngnom` of the type `integer`) may be used.

## 3.10 Face and cell mesh-defined properties and selection

The mesh entities may be referenced by the user during the mesh creation. These references may then be used to mark out some mesh entities according to the need (specification of boundary conditions, pressure drop zones, ...). The references are generally of one of the two following types:

- color. A color is an integer possibly associated with boundary faces and volume elements by the mesh generator. Depending on the tool, this concept may have different names, which *Code\_Saturne* interprets as colors. Most tools allow only one color per face or element.
  - I-deas uses a color number with a default of 7 (green) for elements, be they volume elements or boundary "surface coating" elements. Color 11 (red) is used for vertices, but vertex properties are ignored by *Code\_Saturne*.
  - SIMAIL uses the equivalent notions of "reference" for element faces, and "subdomain" for volume elements. By default, element faces are assigned no reference (0), and volume elements domain 1.
  - Gmsh uses "physical property" numbers.
  - EnSight has no similar notion, but if several parts are present in an EnSight 6 file, or several parts are present *and* vertex ids are given in an EnSight Gold file, the part number is interpreted as a color number by the Preprocessor.
  - The MED 2.3 model allowed integer "attributes", though many tools working with this format ignored those and only handle groups.
- groups. Named "groups" of mesh entities may also be used with many mesh generators or formats. In some cases, a given cell or face may belong to multiple groups (as some tools allow new groups to be defined by boolean operations on existing groups). In *Code\_Saturne*, every group is assigned a group number (base on alphabetical ordering of groups).

- Ideas assigns a group number with each group, but by default, this number is just a counter. Only the group name is considered by *Code\_Saturne* (so that elements belonging to two groups with identical names and different numbers are considered as belonging to the same group).
- CGNS allows both for named boundary conditions and mesh sections. If present, boundary condition names are interpreted as group names, and groups may also be defined based on element section or zone names using additional Preprocessor options (`-grp-cel` or `-grp-fac` followed by `section` or `zone`).
- Using the MED format, it is preferable to use “groups” than colors, as many tools ignore the latter.

Selection criteria may be defined in a similar fashion whether using the GUI or in user subroutines. Typically, a selection criteria is simply a string containing the required color numbers or group names, possibly combined using boolean expressions. Simple geometric criteria are also possible.

A few examples are given below:

```
ENTRY
1 or 7
all[]
3.1 >= z >= -2 or not (15 or entry)
range[04, 13, attribute]
sphere[0, 0, 0, 2] and (not no_group[])
```

Strings such as group names containing white-space or having names similar to reserved operators may be protected using “escape characters”.<sup>18</sup> More complex examples of strings with protected strings are given here:

```
"First entry" or Wall\ or\ sym
entry or \plane or "noone's output"
```

The following operators and syntaxes are allowed (fully capitalized versions of keywords are also allowed, but mixed upper-case/lower-case versions are not):

#### escape characters

protect next character only:	\
protect string:	'string'    "string"

#### basic operators

priority:	(    )
not:	not    !    !=
and:	and    &    &&
or:	or               ,    ;
xor:	xor    ^

#### general functions

select all:	all[]
entities having no group or color:	no_group[]
select a range of groups or colors:	range[first, last]
	range[first, last, group]
	range[first, last, attribute]

For the range operator, *first* and *last* values are inclusive. For attribute (color) numbers, natural integer value ordering is used, while for group names, alphabetical ordering is used. Note also that in the bizarre (not recommended) case in which a mesh would contain for example both a color number

<sup>18</sup>Note that for defining a string in Fortran, double quotes are easier to use, as they do not conflict with Fortran's single quotes delimiting a string. In C, the converse is true. Also, in C, to define a string such as `\plane`, the string `\\plane` must be used, as the first `\` character is used by the compiler itself. Using the GUI, either notation is easy.

EDF R&D	<b>Code_Saturne version 3.2.0 practical user's guide</b>	Code_Saturne documentation Page 53/202
---------	--	--

15 and a group named “15”, using `range[15, 15, group]` or `range[15, 15, attribute]` could be used to distinguish the two.

Geometric functions are also available. The coordinates considered are those of the cell or face centres. Normals are of course usable only for face selections, not cell selections.

#### geometric functions

face normals:	<code>normal[x, y, z, epsilon]</code>
	<code>normal[x, y, z, epsilon = epsilon]</code>
plane, $ax + by + cz + d = 0$ form:	<code>plane[a, b, c, d, epsilon]</code>
	<code>plane[a, b, c, d, epsilon = epsilon]</code>
	<code>plane[a, b, c, d, inside]</code>
	<code>plane[a, b, c, d, outside]</code>
plane, normal + point in plane form:	<code>plane[n<sub>x</sub>, n<sub>y</sub>, n<sub>z</sub>, x, y, z, epsilon]</code>
	<code>plane[n<sub>x</sub>, n<sub>y</sub>, n<sub>z</sub>, x, y, z, epsilon = epsilon]</code>
	<code>plane[n<sub>x</sub>, n<sub>y</sub>, n<sub>z</sub>, x, y, z, inside]</code>
	<code>plane[n<sub>x</sub>, n<sub>y</sub>, n<sub>z</sub>, x, y, z, outside]</code>
box, extents form:	<code>box[x<sub>min</sub>, y<sub>min</sub>, z<sub>min</sub>, x<sub>max</sub>, y<sub>max</sub>, z<sub>max</sub>]</code>
box, origin + axes form:	<code>box[x<sub>0</sub>, y<sub>0</sub>, z<sub>0</sub>,</code> <code>dx<sub>1</sub>, dy<sub>1</sub>, dz<sub>1</sub>, dx<sub>2</sub>, dy<sub>2</sub>, dz<sub>2</sub>, dx<sub>3</sub>, dy<sub>3</sub>, dz<sub>3</sub>]</code>
cylinder:	<code>cylinder[x<sub>0</sub>, y<sub>0</sub>, z<sub>0</sub>, x<sub>1</sub>, y<sub>1</sub>, z<sub>1</sub>, radius]</code>
sphere:	<code>sphere[x<sub>c</sub>, y<sub>c</sub>, z<sub>c</sub>, radius]</code>
inequalities:	<code>&gt;, &lt;, &gt;=, &lt;=</code> associated with <code>x, y, z</code> or <code>X, Y, Z</code> keywords and coordinate value; <code>x<sub>min</sub> &lt;= x &lt; x<sub>max</sub></code> type syntax is allowed.

In the current version of *Code\_Saturne*, all selection criteria used are maintained in a list, so that re-interpreting a criterion already encountered (such as at the previous time step) is avoided. Lists of entities corresponding to a criteria containing no geometric functions are also saved in a compact manner, so re-using a previously used selection should be very fast. For criteria containing geometric functions, the full list of corresponding entities is not maintained, so each entity must be compared to the criterion at each time step. Heavy use of many selection criteria containing geometric functions may thus lead to reduced performance.

## 4 Importing and Preprocessing Meshes

Importing meshes is done by the Preprocessor module, while and preprocessing is done mainly by the code Kernel (except for element orientation checking, which is done by the Preprocessor).

The Preprocessor module of *Code\_Saturne* reads the mesh file(s) (under any supported format) and translates the necessary information into a Kernel input file.

When multiple meshes are used, the Preprocessor is called once per mesh, and each resulting output is added in a `mesh.input` directory (instead of a single `mesh.input` file).

The executable of the Preprocessor module is `cs_preprocess`, and is normally called through the run script, so it is not in standard paths (it is at `<prefix>/libexec/code_saturne/cs_preprocess`). Its most useful options and sub-options are described briefly here. To obtain a complete and up-to-date list of options and environment variables, use the following command: `cs_preprocess -h` or `cs_preprocess --help`. Many options, such as this one, accept a short and a long version.

Nonetheless, it may be useful to call the Preprocessor manually in certain situations, especially for frequent verification when building a mesh, so its use is described here. Verification may also be done using the GUI or the mesh quality check mode of the general run script.

The Preprocessor is controlled using command-line arguments. A few environment variables allow advanced users to modify some behaviours or to obtain a trace of memory management.

## 4.1 Preprocessor options

Main choices are done using command-line options. For example:

```
cs_preprocess --num 2 fluid.med
```

means that we read the second mesh defined in the `fluid.med` file, while:

```
cs_preprocess --no-write --post-volume med fluid.msh
```

means that we read file `fluid.msh`, and do not produce a `mesh_input` file, but do output a `fluid.med` file (effectively converting a Gmsh file to a MED file).

### 4.1.1 Mesh selection

Any use of the preprocessor requires one mesh file (except for `cs_preprocess` and `cs_preprocess -h` which respectively print the version number and list of options). This file is selected as the last argument to `cs_preprocess`, and its format is usually automatically determined based on its extension (c.f. 3.4.1 page 20) but a `--format` option allows forcing the format choice of the selected file.

For formats allowing multiple meshes in a single file, the `--num` option followed by a strictly positive integer allows selection of a specific mesh; by default, the first mesh is selected.

For meshes in CGNS format, we may in addition use the `--grp-cel` or `--grp-fac` options, followed by the `section` or `zone` keywords, to define additional groups of cell or faces based on the organization of the mesh in sections or zones. The sub-options have no effect on meshes of other formats.

### 4.1.2 Post-processing output

By default, the Preprocessor does not generate any post-processor output. By adding `--post-volume [format]`, with the optional `format` argument being one of `ensight`, `med`, or `cgns` to the command-line arguments, the output of the volume mesh to the default or indicated format is provoked.

In case of errors, output of error visualization output is always produced, and by adding `--post-error [format]`, the format of that output may be selected (from one of `ensight`, `med`, or `cgns`, assuming MED and CGNS are available),

### 4.1.3 Element orientation correction

We may activate the possible element orientation correction using the `--reorient` option.

Note that we cannot guarantee correction (or even detection) of a bad orientation in all cases. Not all local numbering possibilities of elements are tested, as we focus on “usual” numbering permutations. Moreover, the algorithms used may produce false positives or fail to find a correct renumbering in the case of highly non convex elements. In this case, nothing may be done short of modifying the mesh, as without a convexity hypothesis, it is not always possible to choose between two possible definitions starting from a point set.

With a post-processing option such as `--post-error` or, `--post-volume`, visualizable meshes of corrected elements as well as remaining badly oriented elements are generated.

## 4.2 Environment variables

Setting a few environment variables specific to the Preprocessor allows modifying its default behaviour. In general, if a given behaviour is modifiable through an environment variable rather than by a command-line option, it has little interest for a non-developer, or its modification is potentially hazardous. The environment variables used by the Preprocessor are described here:

#### CS\_PREPROCESS\_MEM\_LOG

Allows defining a file name in which memory allocation, reallocation, and freeing is logged.

#### CS\_PREPROCESS\_MIN\_EDGE\_LEN

Under the indicated length ( $10^{-15}$  by default), an edge is considered to be degenerate and its vertices will be merged after the transformation to descending connectivity. Degenerate edges and faces will thus be removed. Hence, the post-processed element does not change, but the Kernel may handle a prism where the preprocessor input contained a hexahedron with two identical vertex couples (and thus a face of zero surface). If the Preprocessor does not print any information relative to this type of correction, it means that it has not been necessary. To completely deactivate this automatic correction, a negative value may be assigned to this environment variable.

#### CS\_PREPROCESS\_IGNORE\_IDEAS\_COO\_SYS

If this variable is defined and is a strictly positive integer, coordinate systems in I-deas universal format files will be ignored. The behaviour of the Preprocessor will thus be the same as that of versions 1.0 and 1.1. Note that in any case, non Cartesian coordinate systems are not handled yet.

### 4.2.1 System environment variables

Some system environment variables may also modify the behaviour of the Preprocessor. For example, if the Preprocessor was compiled with MED support on an architecture allowing shared (dynamic) libraries, the LD\_PRELOAD environment variable may be used to define a “priority” path to load MED or HDF5 libraries, and thus experiment with another version of these libraries without recompiling the Preprocessor. To determine which shared libraries are used by an executable file, use the following command: `ldd {executable_path}`.

## 4.3 Optional functionality

Some functions of the Preprocessor are based on external libraries, which may not always be available. It is thus possible to configure and compile the Preprocessor so as not to use these libraries. When running the Preprocessor, the supported options are printed. The following optional libraries may be used:

- CGNS library. In its absence, [CGNS](#) format support is deactivated.
- MED-file library. In its absence, [MED](#) format is simply deactivated.
- libCCMIO library. In its absence, [CCM](#) format is simply deactivated.
- Read compressed files using Zlib. With this option, it is possible to directly read mesh files compressed with a *gzip* type algorithm and bearing a *.gz* extension. This is limited to formats not already based on an external library (i.e. it is not usable with CGNS, MED, or CCM files), and has memory and CPU time overhead, but may be practical. Without this library, files must be uncompressed before use.

## 4.4 General remarks

Note that the Preprocessor is in general capable of reading all “classical” element types present in mesh files (triangles, quadrangles, tetrahedra, pyramids, prisms, and hexahedra). Quadratic or cubic elements are converted upon reading into their linear counterparts. Vertices referenced by no element (isolated vertices or centres of higher-degree elements) are discarded. Meshes are read in the order defined by the user and are appended, vertex and element indices being incremented appropriately.<sup>19</sup>

<sup>19</sup>Possible entity labels are not maintained, as they would probably not be unique when appending multiple meshes.

At this stage, volume elements are sorted by type, and the fluid domain post-processing output is generated if required.

In general, groups assigned to vertices are ignored. selections are thus based on faces or cells. with tools such as SIMAIL, faces of volume elements may be referenced directly, while with I-deas or SALOME, a layer of surface elements bearing the required colors and groups must be added. Internally, the Preprocessor always considers that a layer of surface elements is added (i.e. when reading a SIMAIL mesh, additional faces are generated to bear cell face colors. When building the *faces* → *cells* connectivity, all faces with the same topology are merged: the initial presence of two layers of identical surface elements belonging to different groups would thus lead to a calculation mesh with faces belonging to two groups).

## 4.5 Files passed to the Kernel

Data passed to the Kernel by the Preprocessor is transmitted using a binary file, using “big endian” data representation, named `mesh_input` (or contained in a directory of that name).

When using the Preprocessor for mesh verification, data for the Kernel is not always needed. In this case, the `--no-write` option may avoid creating a Preprocessor output file.

## 4.6 Mesh preprocessing

### 4.6.1 Joining of non-conforming meshes

Conforming joining of possibly non-conforming meshes may be done by the solver, and defined either using the Graphical User Interface (GUI) or the `cs_user_join` user function. In the GUI, the user must add entries in the “Face joining” section of the “Meshes” tab in the item “Calculation environment → Meshes selection”. The user may specify faces to be joined, and can also modify basic joining parameters, see Figure 6. For a simple mesh, it is rarely useful to specify strict face selection criteria, as joining is sufficiently automated to detect which faces may actually be joined. For a more complex mesh, or a mesh with thin walls which we want to avoid transforming into interior faces, it is recommended to filter boundary faces that may be joined by using face selection criteria. This has the additional advantage of reducing the number of faces to test for in the intersection/overlap search, and thus reduced to the time required by the joining algorithm.

One may also modify tolerance criteria using 2 options:

<code>fraction <i>r</i></code>	assigns value <i>r</i> (where $0 < r < 0,49$ ) to the maximum intersection distance multiplier (0,1 by default). The maximum intersection distance for a given vertex is based on the length of the shortest incident edge, multiplied by <i>r</i> . The maximum intersection at a given point along an edge is interpolated from that at its vertices, as shown on the left of Figure 7;
<code>plane <i>c</i></code>	assigns the maximum angle between normals for two faces to be considered coplanar (25° by default); this parameter is used in the second stage of the algorithm, to reconstruct conforming faces, as shown on the right of figure 7.

In practice, we are sometimes led to increase the maximum intersection distance multiplier to 0.2 or even 0.3 when joining curved surfaces, so that all intersection are detected. As this influences merging of vertices and thus simplification of reconstructed faces, but also deformation of “lateral” faces, it is recommended only to modify it if necessary. As for the `plane` parameter, its use has only been necessary on a few meshes up to now, and always in the sense of reducing the tolerance so that face reconstruction does not try to generate faces from initial faces on different surfaces.





Figure 6: Conformal or non-conformal joining

## 4.6.2 Periodicity

Handling of periodicity is based on an extension of conforming joining, as shown on Figure 8. It is thus not necessary for the periodic faces to be conforming (though it usually leads to better mesh quality). All options relative to conforming joining of non-conforming faces also apply to periodicity. Note also that once pre-processed, 2 periodic faces have the same orientation (possibly adjusted by periodicity of rotation).

This operation can also be performed by the solver and specified either using the GUI or the `cs_user_periodicity` function.

As with joining, it is recommended to filter boundary faces to process using a selection criterion. As many periodicities may be built as desired, as long as boundary faces are present. Once a periodicity is handled, faces having periodic matches do not appear as boundary faces, but as interior faces, and are thus not available anymore for other periodicities.

## 4.6.3 Parameters for conforming or non-conforming mesh joinings

The setting of these parameters is done in the user subroutine `cs_user_join` (called once). The user can specify the parameters used for the joining of different meshes. Below is given the list of the standard parameters which can be modified:

- **fract**: the initial tolerance radius associated to each vertex is equal to the length of the shortest incident edge, multiplied by this fraction,
- **plane**: when subdividing faces, 2 faces are considered coplanar and may be joined if the angle



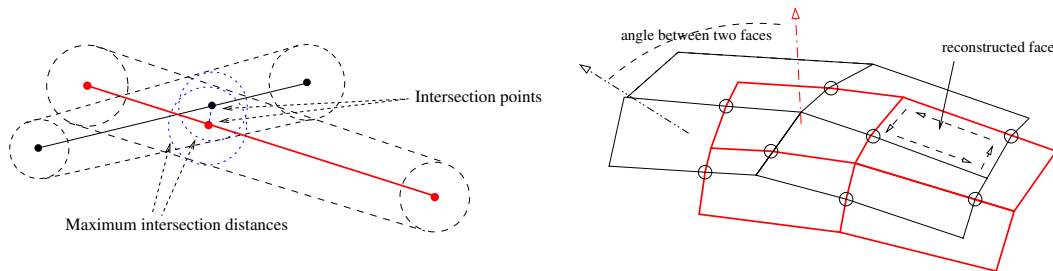


Figure 7: Maximum intersection tolerance and faces normal angle

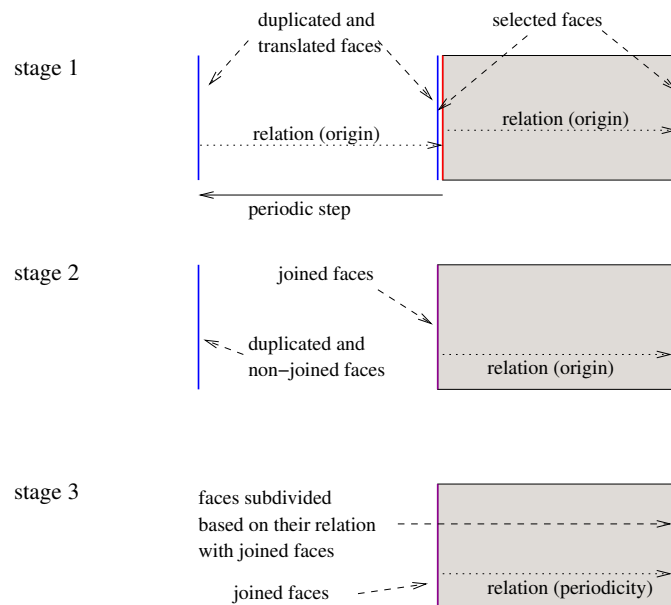


Figure 8: Matching of periodic faces

between their unit normals (in degrees) does not exceed this parameter,

- **iwarnj**: the associated verbosity level (debug level if over 3).

In the call of the function **cs\_join\_add**, a selection criteria for mesh faces to be joined is specified.

The call to the function '**cs\_join\_set\_advanced\_param**' allows defining parameters not available through the GUI.

The list of advanced modifiable parameters is given below:

- **mtf**: a merge tolerance factor, used to locally modify the tolerance associated to each vertex before the merge step. Depending on its value, four scenarii are possible:
  - if  $mtf = 0$ , no vertex merge
  - if  $mtf < 1$ , the vertex merge is more strict. It may increase the number of tolerance reduction and therefore define smaller subset of vertices to merge together but it can drive to loose intersections.
  - if  $mtf = 1$ , no change occurs
  - if  $mtf > 1$ , the vertex merge is less strict. The subset of vertices able to merge is greater.
- **pmf**: a pre-merge factor. This parameter is used to define a limit under which two vertices are merged before the merge step,
- **tcm**: a tolerance computation mode. If its value is:
  - 1 (default), the tolerance is the minimal edge length related to a vertex, multiplied by a fraction.
  - 2, the tolerance is computed like for 1 with, in addition, the multiplication by a coefficient equal to the maximum between  $\sin(e1)$  and  $\sin(e2)$ ; where  $e1$  and  $e2$  are two edges sharing the same vertex  $V$  for which we want to compute the tolerance.
  - 11, it is the same as 1 but taking into account only the selected faces.
  - 12, it is the same as 2 but taking into account only the selected faces.
- **icm**: the intersection computation mode. If its value is:
  - 1 (default), the original algorithm is used. Care should be taken to clip the intersection on an extremity.
  - 2, a new intersection algorithm is used. Caution should be used to avoid to clip the intersection on an extremity.
- **maxbrk**: defines the maximum number of equivalence breaks which is enabled during the merge step,
- **maxsf**: defines the maximum number of sub-faces used when splitting a selected face

The following are advanced parameters used in the search algorithm for face intersections between selected faces (octree structure). They are useful in case of memory limitation:

- **tml**: the tree maximum level is the deepest level reachable during the tree building,
- **tmb**: the tree maximum boxes is the maximum number of bounding boxes (BB) which can be linked to a leaf of the tree (not necessary true for the deepest level),
- **tmr**: the tree maximum ratio. The building of the tree structure stops when the number of bounding boxes is superior to the product of **tmr** with the number of faces to locate. This is an efficient parameter to reduce memory consumption.

#### 4.6.4 Parameters for periodicity

Periodicities can be set directly in the Graphical User Interface (GUI) or using the user function `cs_user_periodicity` (called once during the calculation initialisation). In both cases, the user can choose between 3 types of periodicities: translation, rotation, or mixed (see Figure9). Then specific parameters must be set.

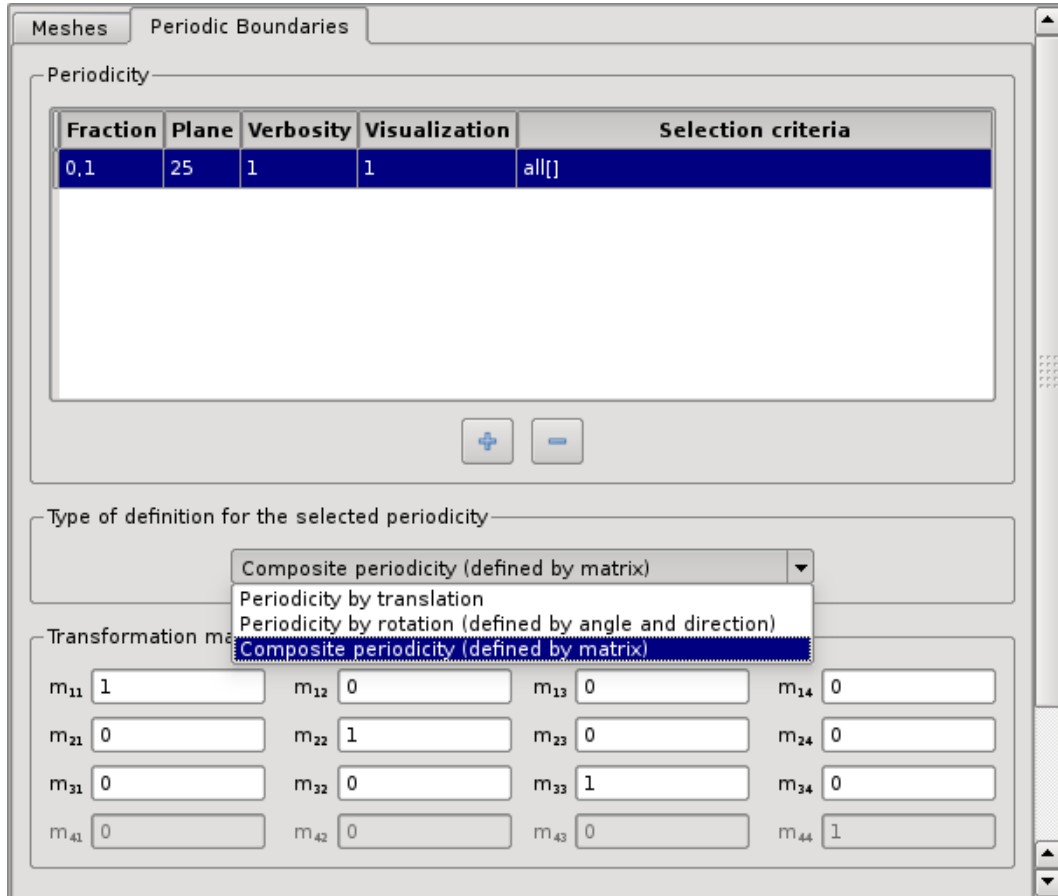


Figure 9: Periodicity

As periodicity is an extension of mesh joining, all parameters (whether basic or advanced) available for mesh joining are also applicable for periodicity, in addition to the parameters defining the periodicity transformation.

#### 4.6.5 Modification of the mesh geometry

*Functions called only once during the calculation initialisation.*

The user function `cs_user_mesh_input` allows a detailed selection of imported meshes read, reading files multiple times, applying geometric transformations, and renaming groups.

The user function `cs_user_mesh_modify` may be used for advanced modification of the main `cs_mesh_t` structure.

**WARNING:** Caution must be exercised when using this function along with periodicity. Indeed, the periodicity parameters are not updated accordingly, meaning that the periodicity may be unadapted after one changes the mesh vertex coordinates. It is particularly true when one rescales the mesh. Rescaling

*should thus be done in a separate run, before defining periodicity.*

The user function `cs_user_mesh_thinwall` allows insertion of thin walls in the calculation mesh. Currently, this function simply transforms the selected internal faces into boundary faces, on which boundary conditions can (and must) be applied.

Faces on each side of a thin wall will share the same vertices, so post-processing of the main volume mesh may not show the inserted walls, though they will appear in the main boundary output mesh.

## 4.7 Mesh smoothing utilities

*Function called once only during the calculation initialisation.*

The smoothing utilities may be useful when the calculation mesh has local defects. The principle of smoothers is to mitigate the local defects by averaging the mesh quality. This procedure can help for the calculus robustness or/and results quality.

The user function `cs_user_mesh_smoothe` allows to use different smoothing functions detailed below.

*WARNING 1: Caution must be exercised when using this function along with periodicity. Indeed, the periodicity parameters are not currently updated accordingly, meaning that the periodicity may be unadapted after one changes the mesh vertex coordinates. It is particularly true when one rescales the mesh. Rescaling should thus be done in a separate run, before defining periodicity.*

*WARNING 2: Caution must be exercised when using smoothing utilities because the geometry may be modified. In order to preserve geometry, the function `cs_mesh_smoother_fix_by_feature` allows to fix by a feature angle criterion the mobility of boundary vertices.*

### 4.7.1 Fix by feature

The vertex normals are defined by the average of the normals of the faces sharing the vertex. The feature angle between a vertex and one of its adjacent faces is defined by the angle between the vertex normal and the face normal.

This function sets a vertex if one of its feature angles is less than  $\cos(\theta)$  where  $\theta$  is the maximum feature angle (in degrees) defined by the user. In fact, if  $\theta = 0^\circ$  all boundary vertices will be fixed, and if  $\theta = 90^\circ$  all boundary vertices will be free.

Fixing all boundary vertices ensures the geometry is preserved, but reduces the smoothing algorithm's effectiveness.

### 4.7.2 Warped faces smoother

The function `cs_mesh_smoother_unwarp` allows to reduce face warping in the calculation mesh.

Be aware that, in some cases, this algorithm may degrade other mesh quality criteria.

## 5 Partitioning for parallel runs

Graph partitioning (using one of the optional METIS or SCOTCH libraries) is done by the Kernel. Unless explicitly deactivated, this stage produces one or several “cell  $\rightarrow$  domain” distribution files, named `domain_number_p` for a partitioning on  $p$  sub-domains, which may be read when starting a subsequent computation so as to avoid re-running that stage. These files are placed in a directory named `partition_output`.

The Kernel redistributes data read in `mesh_input` based on the associated (re-read or computed) partitioning, so there is no need to run any prior script when running on a different number of processors, although a previous partitioning may optionally be re-used.

Without a graph-based partitioning library, or based on the user's choice, the Kernel will use a built-in partitioning using a space-filling curve (Z or Hilber curve) technique. This usually leads to partitionings of lower quality than with graph partitioning, but parallel performance remains reasonable.

Partitioning options may be defined using the GUI or by calling the appropriate functions in the `cs_user_partition_options` user function.

## 5.1 Partitioning stages

Partitioning is always done just after reading the mesh, unless a partitioning input file is available, in which case the partitioning replaces this stage.

When a mesh modification implying a change of cell connectivity graph is expected, the mesh may be re-partitioned after the pre-processing stage, prior to calculation. By default, re-partitioning is only done if the partitioning algorithm chosen for that stage is expected to produce different results due to the connectivity change. This is the case for graph-based algorithms such as those of METIS or SCOTCH, when mesh joining is defined, or additional periodic matching is defined (and the algorithm is not configured to ignore periodicity information).

There are thus two possible partitioning stages:

- `CS_PARTITION_FOR_PREPROCESS`, which is optional, and occurs just after reading the mesh.
- `CS_PARTITION_MAIN`, which occurs just after reading the mesh if it is the only stage, or after mesh preprocessing (and before computation), if the partitioning for preprocessing stage is activated.

The number of partitioning stages is determined automatically based on information provided through `cs_partition_set_preprocess_hints()`, but re-partitioning may also be forced or inhibited using the `cs_partition_set_preprocess()` user function.

## 5.2 Partitioner choice

If the Kernel has been configured with both PT-SCOTCH or SCOTCH and PARMETIS or METIS libraries, PT-SCOTCH will be used by default<sup>20</sup>, but the user may force the selection of another partitioning type using either the GUI or user routines.

In addition to graph-based partitionings, a space-filling curve based algorithm is available, using either a Morton (Z) or Peano-Hilbert curve, in the computation domain's bounding box or bounding curve.

When partitioning for preprocessing, a space-filling curve is used, unless forced by calling `cs_partition_set_algorithm()` with the appropriate algorithm choice for the `CS_PARTITION_FOR_PREPROCESS` stage.

## 5.3 Effect of periodicity

By default, face periodicity relations are taken into account when building the "cell → cell" connectivity graph used for partitioning. This allows better partitioning optimization, but increases the probability of having groups of cells at opposite sides of the domain in a same sub-domain. This is not an issue for standard calculations, but may degrade performance of search algorithms based on bounding boxes. It is thus possible to ignore periodicity when partitioning a mesh.

Also, partitioning using a space-filling curve ignores periodicity.

Note that nothing guarantees that a graph partitioner will not place disjoint cells in the same sub-domain independently of this option, but this behaviour is rare.

<sup>20</sup>Though PARMETIS will be chosen before serial SCOTCH in a parallel run

## 6 Basic modelling setup

### 6.1 Initialisation of the main parameters

This operation is done in the Graphical User Interface (GUI) or by using the user subroutines in `cs_user_parameters.f90`. In the GUI, the initialisation is performed by filling the parameters displayed in Figure 10 to 27. If the option 'Mobile mesh' is activated in Figure 11, please see Section 8.11.4 for more details. The headings filled for the initialisation of the main parameters are the followings:

- Thermophysical model options: Steady or unsteady algorithm, specific physics, ALE mobile mesh, turbulence model, thermal model and species transport (definition of the scalars and their variances), see Figure 10 to Figure 14. If a thermal scalar, temperature or enthalpy, is selected, two other headings on conjugate heat transfer and radiative transfers can be filled in (see Figure 13).
- Physical properties: reference pressure, velocity and length, fluid properties (density, viscosity, thermal conductivity, specific heat and scalar diffusivity), gravity, see Figure 15 to Figure 17. If non-constant values are used for the fluid properties, and if the GUI is not used, the `cs_user_physical_properties` file must be used, see § 6.5.1.
- Volume conditions: definition of volume regions (for initialisation, head losses and source terms, see § 6.6 and § 6.7), initialisation of the variables (including scalars), Coriolis source term, see Figure 18 and Figure 19.
- Boundary conditions: definition and parametrisation of boundary conditions for all variables (including scalars). If the GUI is not used, the `cs_user_boundary_conditions` file must be used, see § 6.4.
- Numerical parameters: number and type of time steps, and advanced parameters for the numerical solution of the equations, see Figure 20 to Figure 22.
- Calculation control: parameters related to the time averages, the locations of the probes where some variables will be monitored over time (if the GUI is not used, this information is specified in § 6.3), the definition of the frequency of the outputs in the calculation listing, the post-processing output writer frequency and format options, and the post-processing output meshes and variables selection, see Figure 23, Figure 24, Figure 25, and Figure 26. The item "Profiles" allows to save, with a frequency defined by the user, 1D profiles on a parametric curve defined by its equation, see Figure 27.

With the GUI, the subroutine `cs_user_parameters.f90` is only used to modify high-level parameters which can not be managed by the interface. Without the GUI, this subroutine is compulsory and some of the headings must be completed (see § 3.2.1). `cs_user_parameters.f90` is used to indicate the value of different calculation basic parameters: constant and uniform physical values, parameters of numerical schemes, input-output management...

It is called only during the calculation initialisation.

For more details about the different parameters, please refer to the key word list (§9).

`cs_user_parameters.f90` is in fact constituted of 6 separate subroutines: `usipph`, `usinsc`, `usipsc`, `usipgl`, `usipsu` and `usipes`. Each one controls various specific parameters. The keywords which are not featured in the supplied example can be provided by the user in `SRC/REFERENCE/base`; in this case, understanding of the comments is required to add the keywords in the appropriate subroutine, it will ensure that the value has been well defined. The modifiable parameters in each of the subroutines of `cs_user_parameters.f90` are:

- `usipph`: `iturb`, `icp` and `irccor` (don't modify these parameters anywhere else)
- `usinsc`: `nscaus` (don't modify this parameter anywhere else)

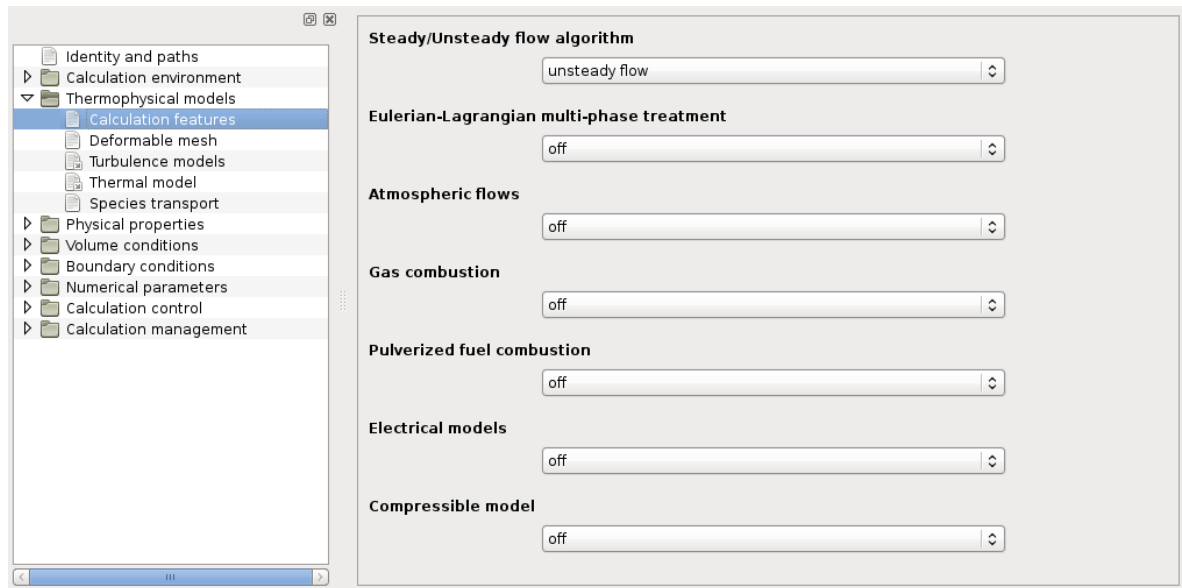


Figure 10: Calculation features options

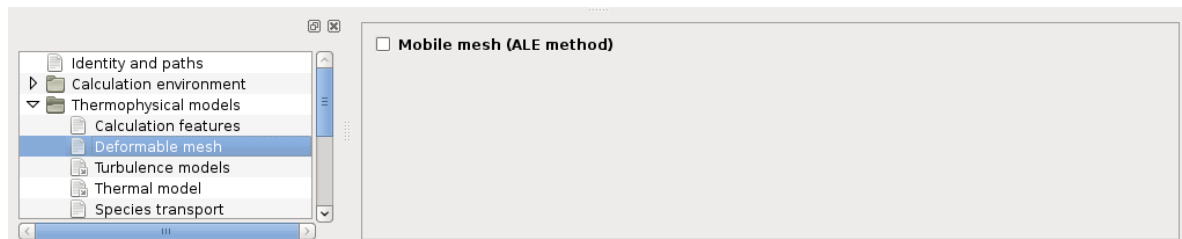


Figure 11: Mobile mesh option

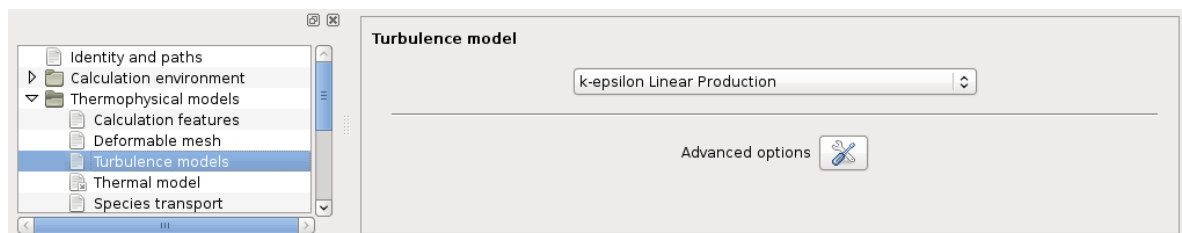


Figure 12: Turbulence model selection

- **usipsc**: `iscavr` and `ivisl`s (don't modify these parameters anywhere else)
- **usipgl**: `idtvar`, `ipucou`, `iphydr`, `idilat` and the parameters related to the error estimators (don't modify these parameters anywhere else).
- **usipsu**: physical parameters of the calculation (thermal scalar, physical properties,...), numerical parameters (time steps, number of iterations, ...), definition of the time averages.
- **usipes**: post-processing output parameters (periodicity, variable names, probe positions,...)

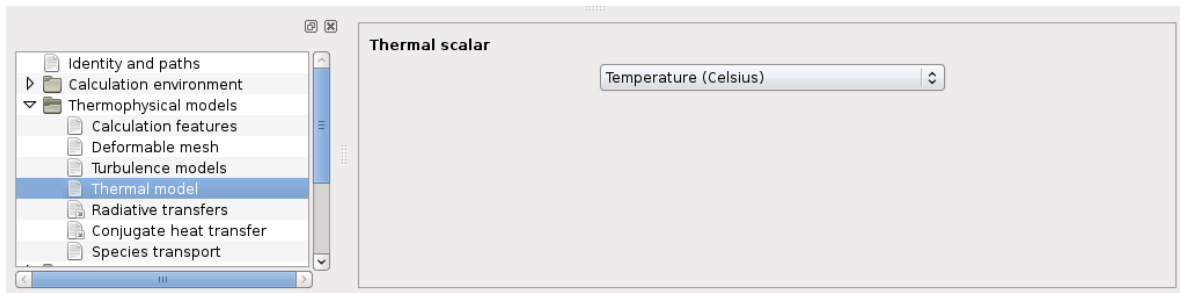


Figure 13: Thermal scalar selection

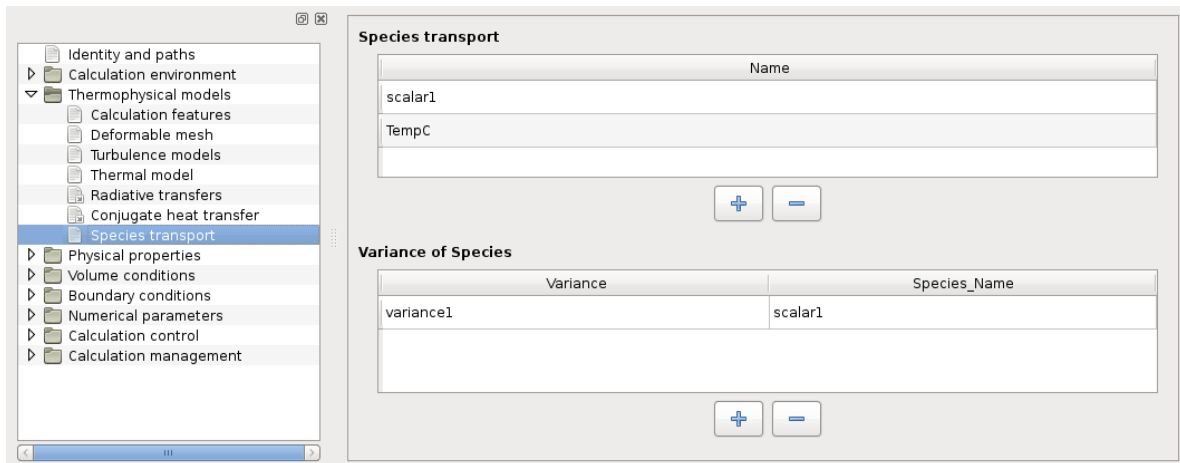


Figure 14: Definition of the transported species/scalars

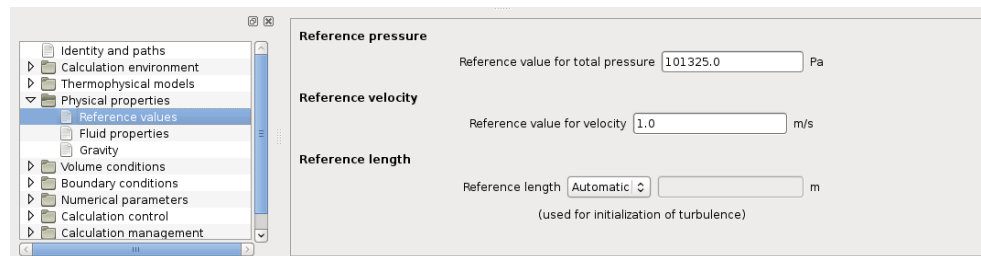


Figure 15: Setting of the reference values for pressure, velocity and length

For more details on the different parameters, see the list of keywords (§ 9). The names of the keywords can also be seen in the help sections of the interface.

## NOTES

- The table `iscavr` is filled with the user scalars which represent the mean square fluctuations of another scalar amongst the list of the `nscaus` scalars. For the other scalars, `iscavr` does not need to be completed (by default, `iscavr(ii) ≤ 0`). For instance, if the scalar `jj` represents the average of the square of the fluctuations of the scalar `kk`, the user must indicate `iscavr(jj)=kk` ( $1 ≤ kk ≤ nscaus$ ).
- When using the interface, only the additional parameters (which can not be defined in the interface) should appear in `cs_user_parameters.f90`. The user needs then only to activate examples which are useful for his case (replacing `if (.false.)` with `if (.true.)`, or removing such tests).



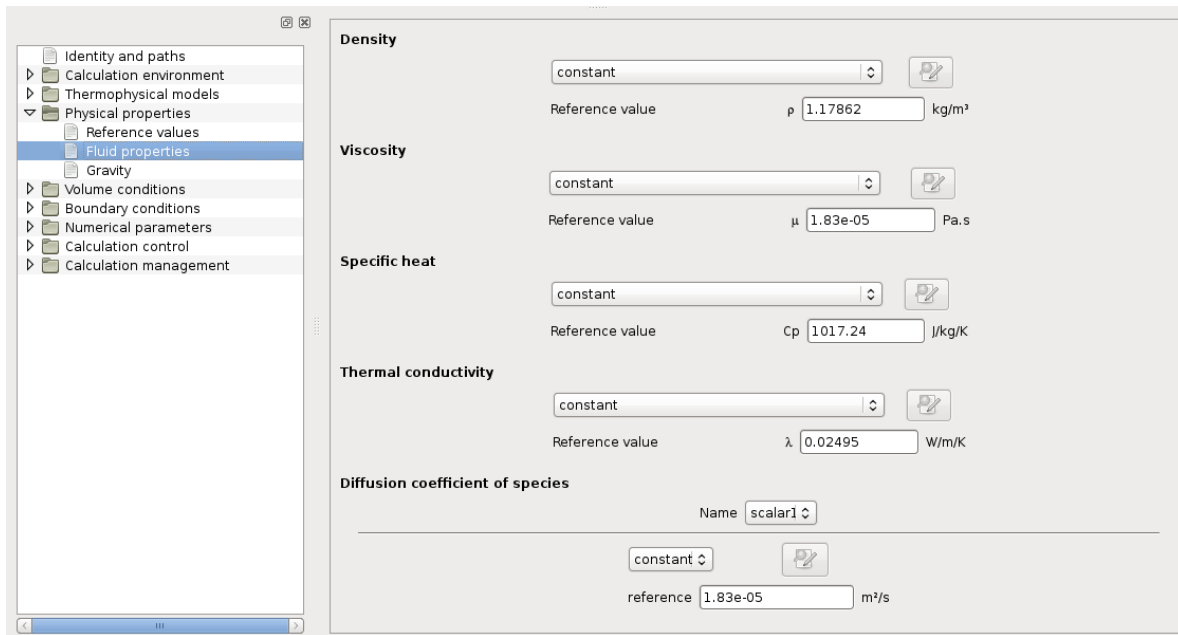


Figure 16: Fluid properties

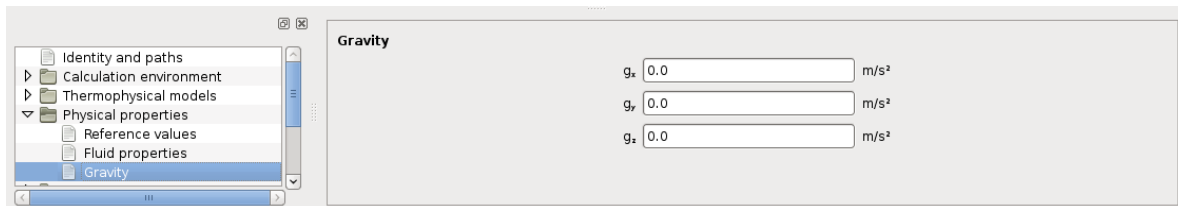


Figure 17: Setting of the gravity

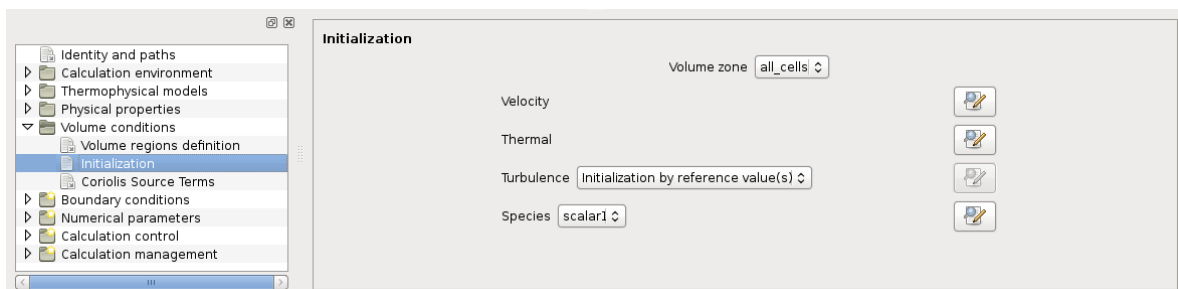


Figure 18: Initialisation of variables

## 6.2 Selection of mesh inputs: `cs_user_mesh_input`

*Subroutine called only during the calculation initialisation.*

This C function may be used to select which mesh input files are read, and apply optional coordinate transformations or group renumberings to them. By default, the input read is a file or directory named `mesh_input`, but if this function is used, any file may be selected, and the same file may be read multiple times (applying a different coordinate transformation each time). All inputs read through

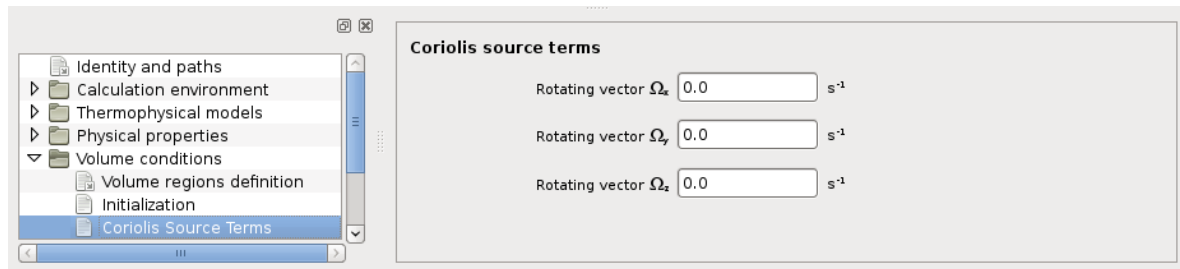


Figure 19: Setting of the Coriolis source term

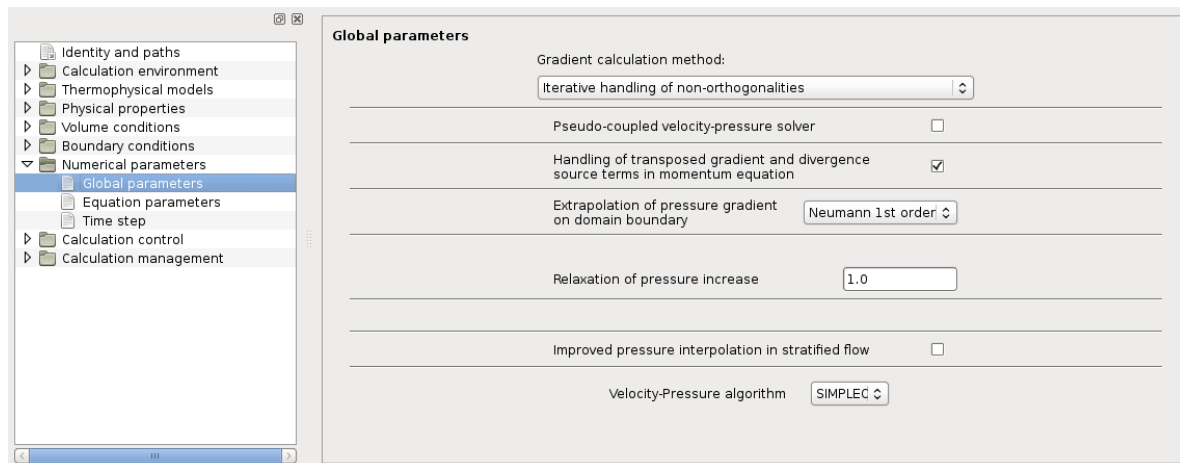


Figure 20: Global resolution parameters

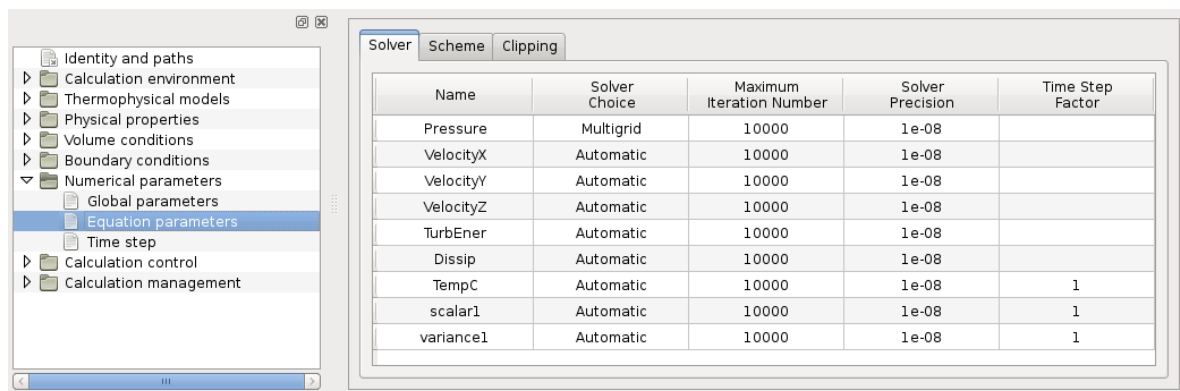


Figure 21: Numerical parameters for the main variables resolution

this function are automatically concatenated, and may be later joined using the mesh joining options.

Geometric transformations are defined using a homogeneous coordinates transformation matrix. Such a matrix has 3 lines and 4 columns, with the 3 first columns describing a rotation/scaling factor, and the last column describing a translation. A 4th line is implicit, containing zeroes off-diagonal, and 1 on the diagonal. The advantages of this representation is that any rotation/translation/scaling combination may be expressed by matrix multiplication, while simple rotations or translations may still be defined easily.

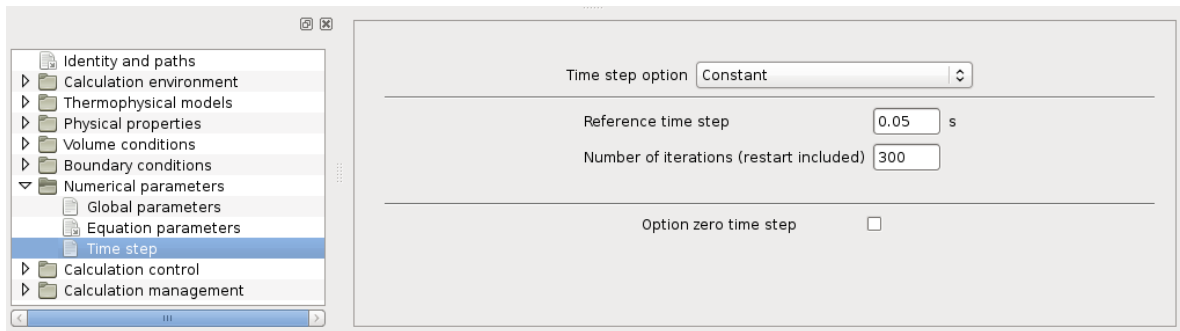


Figure 22: Time step settings

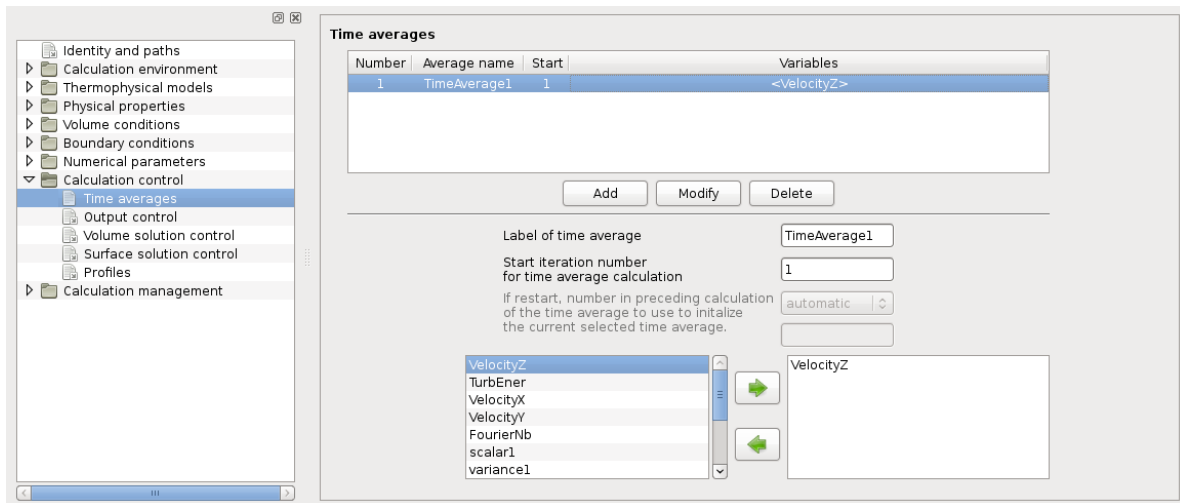


Figure 23: Management of time averaged variables

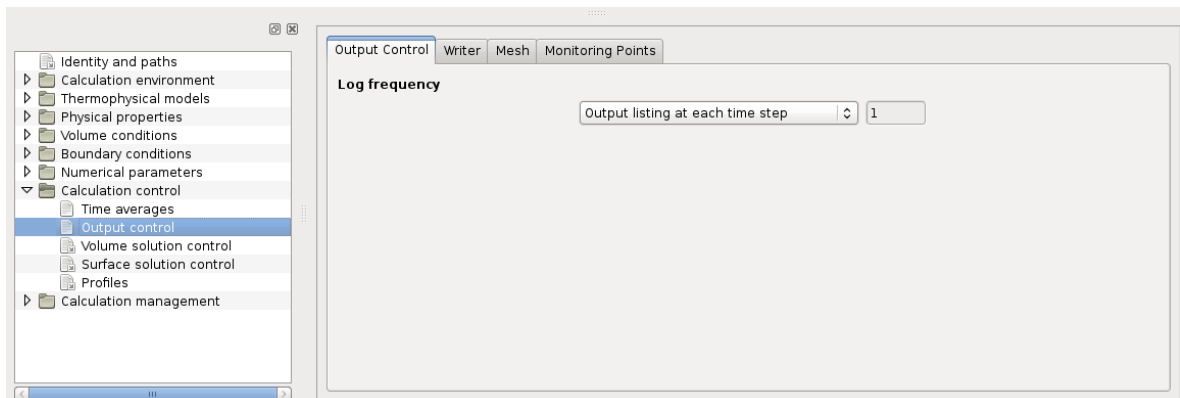


Figure 24: Parameters of chronological logging options

## 6.3 Non-default variables initialisation

The non-default variables initialisation is performed in the subroutine `cs_user_initialization` (called only during the calculation initialisation).

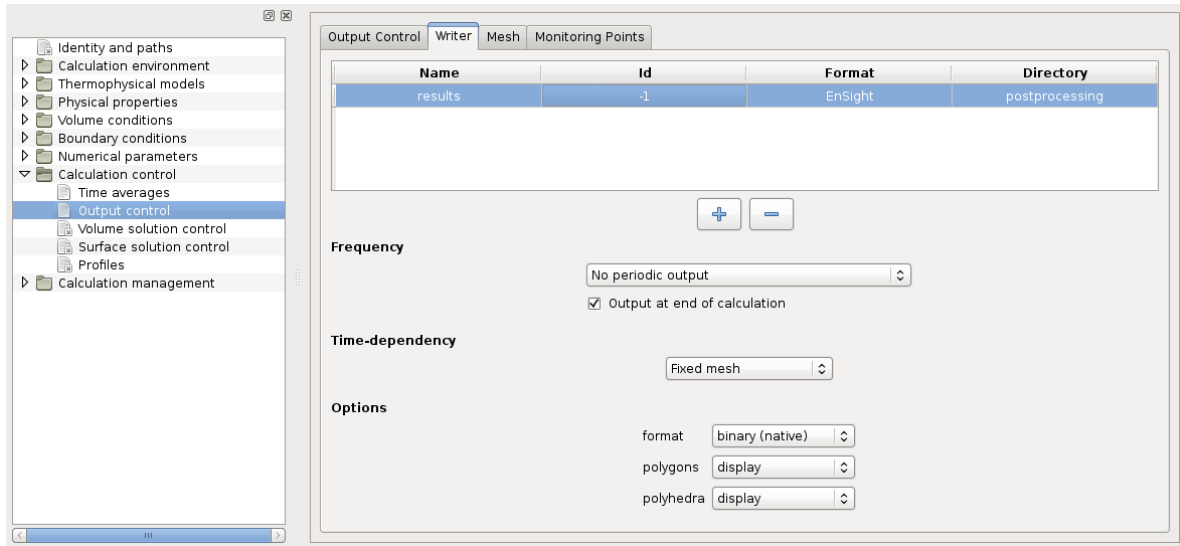


Figure 25: Management of postprocessing writers

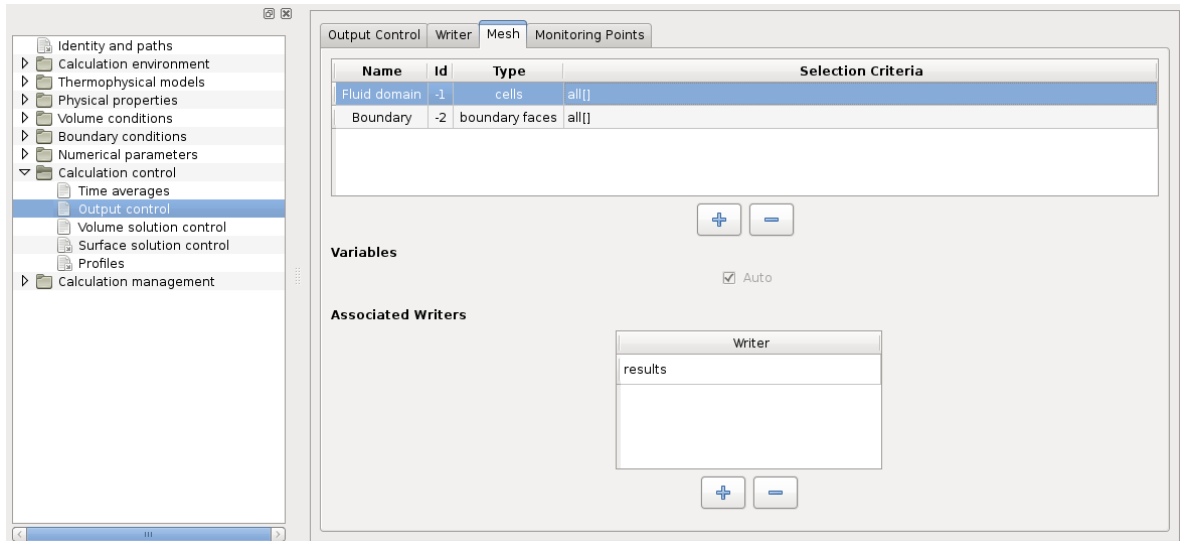


Figure 26: Management of postprocessing meshes

At the calculation beginning, the variables are initialised automatically by the code. Velocities and scalars are set to 0 (or *scamax* or *scamin* if 0 is outside the acceptable scalar variation range), and the turbulent variables are estimated from *uref* and *almax*.

For the  $k$  in  $k - \varepsilon$ ,  $R_{ij} - \varepsilon$ , v2f or  $k - \omega$  models:

$\text{rtp}(\text{i}1, \text{i}k\text{iph}) = 1.5 * (0.02 * \text{uref}) ** 2$  (in  $R_{ij} - \varepsilon$ ,  $R_{ij} = \frac{2}{3} k \delta_{ij}$ )

For the  $\varepsilon$  in  $k - \varepsilon$ ,  $R_{ij} - \varepsilon$  or v2f models:

$\text{rtp}(\text{i}1, \text{i}e\text{iph}) = \text{rtp}(\text{i}1, \text{i}k\text{iph}) ** 1.5 * \text{cmu} / \text{almax}$

For  $\omega$  in the  $k - \omega$  model:

$\text{rtp}(\text{i}1, \text{i}o\text{mgip}) = \text{rtp}(\text{i}1, \text{i}k\text{iph}) ** 0.5 / \text{almax}$

For  $\varphi$  and  $\bar{f}$  in the v2f models:

$\text{rtp}(\text{i}1, \text{i}p\text{hiph}) = 2/3$

$\text{rtp}(\text{i}1, \text{i}f\text{biph}) = 0$

For  $\alpha$  in EBRSM and BL-v2/k models:

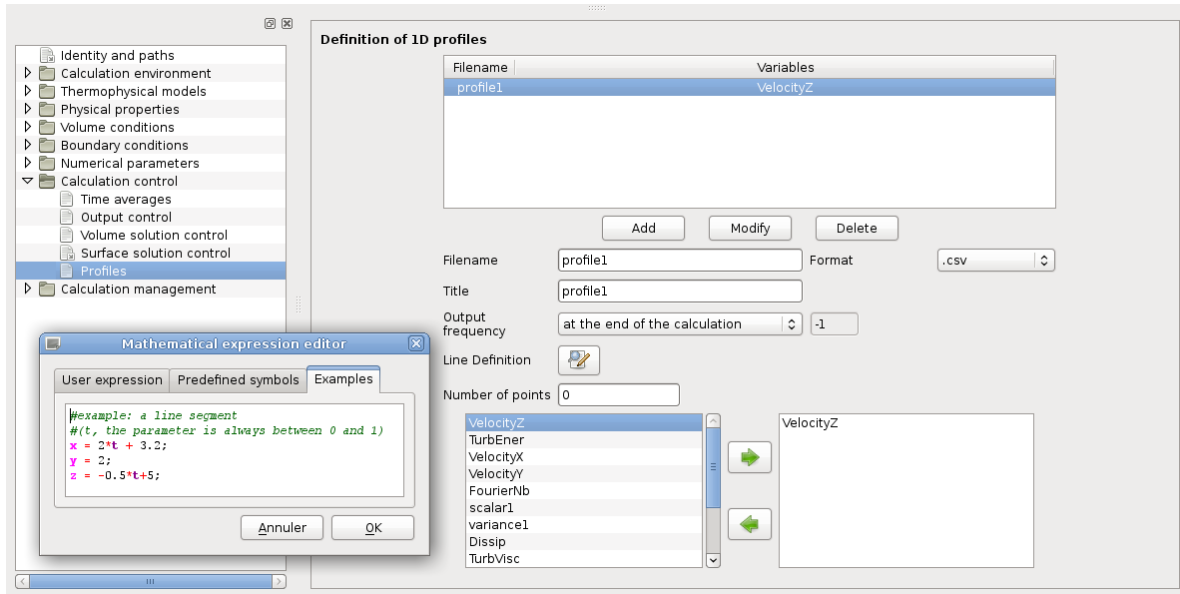


Figure 27: Management of 1D profiles of the solution

```
rtp(iel,ial) = 1
```

For  $\tilde{\nu}_t$  in Spalart-Allmaras model:

```
rtp(iel,inusa)=sqrt(1.5)*0.02*uref*almax
```

The subroutine `cs_user_initialization` allows if necessary to initialise certain variables to values closer to their estimated final values, in order to obtain a faster convergence.

This subroutine allows also to make a non-standard initialisation of physical parameters (density, viscosity, ...), to impose a local value of the time step, or to modify some parameters (time step, variable specific heat, ...) in the case of a calculation restart.

#### NOTE: VALUE OF THE TIME STEP

- For calculations with constant and uniform time step (`idtvar=0`), the value of the time step is `dtref`, given in the parametric file of the interface or `cs_user_parameters.f90`.
- For calculations with a non-constant time step (`idtvar=1` or `2`) which is not a calculation restart, the value of `dtref` given in the parametric file of the interface or in `cs_user_parameters.f90` is used to initialise the time step.
- For calculations with a non-constant time step (`idtvar=1` or `2`) which is a restart of a calculation whose time step type was different (for instance, restart using a variable time step of a calculation run using a constant time step), the value of `dtref`, given in the parametric file of the interface or in `cs_user_parameters.f90`, is used to initialise the time step.
- For calculations with non-constant time step (`idtvar=1` or `2`) which is a restart of a calculation whose time step type was the same (for instance, restart with `idtvar=1` of a calculation run with `idtvar=1`), the time step is read from the restart file and the value of `dtref` given in the parametric file of the interface, or in `cs_user_parameters.f90`, is not used.

It follows, that for a calculation with a non-constant time step (`idtvar=1` or `2`) which is a restart of a calculation in which `idtvar` had the same value, `dtref` does not allow to modify the time step. The

user subroutine `cs_user_initialization` allows to modify the array `dt` which contains the value of the time step read from the restart file (array whose size is `ncelet`, defined at the cell centres whatever the chosen time step type is).

## 6.4 Manage boundary conditions

The boundary conditions can be specified in the Graphical User Interface (GUI) under the heading “Boundary conditions” or in the user subroutine `cs_user_boundary_conditions` called every time step. With the GUI, each region and the type of boundary condition associated to it are defined in Figure 28. Then, the parameters of the boundary condition are specified in Figure 29. The colors of the boundary faces may be read directly from a “listing” file created by the Preprocessor. This file can be generated directly by the interface under the heading “Definition of boundary regions → Add from Preprocessor listing → Import groups and references from Preprocessor listing”, see Figure 28.

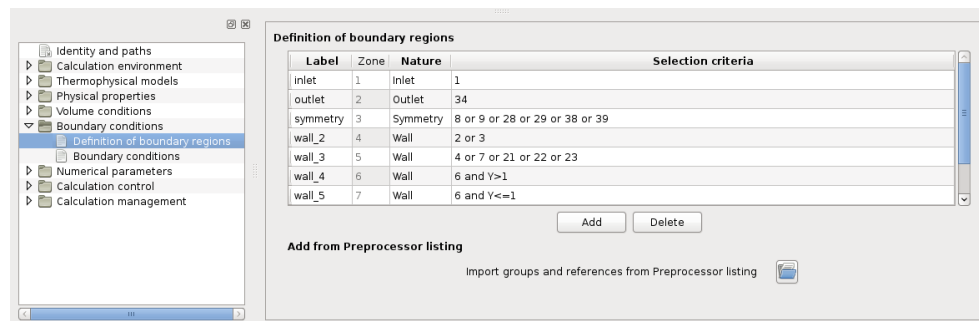


Figure 28: Definition of the boundary conditions

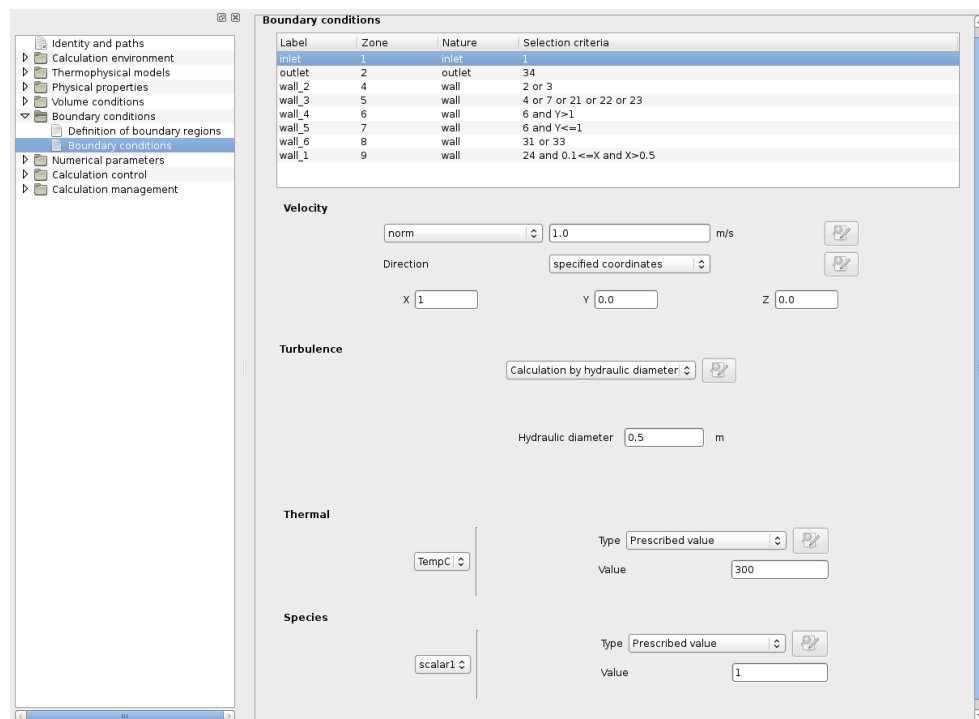


Figure 29: Parameters of the boundary conditions

`cs_user_boundary_conditions` is the second compulsory subroutine for every calculation launched without interface (except in the case of specific physics where the corresponding boundary condition user subroutine must be used).

When using the interface, only complex boundary conditions (input profiles, conditions varying in time, ...) need to be defined with `cs_user_boundary_conditions`. In the case of a calculation launched without the interface, all the boundary conditions must appear in `cs_user_boundary_conditions`.

`cs_user_boundary_conditions` is essentially constituted of loops on boundary face subsets. Several sequences of `call getfbr ('criterion', nlelt, lstelt)` (cf. §3.9.4) allow to select the boundary faces with respect to their group(s), their color(s) or geometric criteria. If needed, geometric and physical variables are also available to the user. These allow him to select the boundary faces using other criteria.

For more details about the treatment of boundary conditions, the user may refer to the theoretical and computer documentation [11] of the subroutine `condli` (for wall conditions, see `clptur`) (to access this document on a workstation, use `code_saturne info --guide theory`).

From the user point of view, the boundary conditions are fully defined by three arrays<sup>21</sup>: `itypfb(nfabor)`, `icodcl(nfabor,nvar)` and `rcodcl(nfabor,nvar,3)`.

- `itypfb(ifac)` defines the type of the face `ifac` (input, wall, ...).
- `icodcl(ifac,ivar)` defines the type of boundary condition for the variable `ivar` on the face `ifac` (Dirichlet, flux ...).
- `rcodcl(ifac,ivar,.)` gives the numerical values associated with the type of boundary condition (value of the Dirichlet, of the flux ...).

In the case of standard boundary conditions (see §6.4.1), it is sufficient to complete `itypfb(ifac)` and parts of the array `rcodcl`; the array `icodcl` and most of `rcodcl` are filled automatically. For non-standard boundary conditions (see §6.4.2), the arrays `icodcl` and `rcodcl` must be fully completed.

## 6.4.1 Coding of standard boundary conditions

The standard keywords used by the indicator `itypfb` are: `ientre`, `iparoi`, `iparug`, `isymet`, `isolib` and `iindef`.

- If `itypfb=ientre`: inlet face.
  - Zero-flux condition for pressure and Dirichlet condition for all other variables. The value of the Dirichlet must be given in `rcodcl(ifac,ivar,1)` for every value of `ivar`, except for `ivar=ipr`. The other values of `rcodcl` and `icodcl` are filled automatically.
- If `itypfb=iparoi`: smooth solid wall face, impermeable and with friction.
  - the eventual sliding wall velocity of the face is found in `rcodcl(ifac,ivar,1)` (`ivar` being `iu`, `iv` or `iw`). The initial values of `rcodcl(ifac,ivar,1)` are zero for the three velocity components (and therefore are to be specified only if the velocity is not equal to zero).  
*WARNING: the wall sliding velocity must belong to the boundary face plane. For safety, the code only uses the projection of this velocity on the face. As a consequence, if the velocity specified by the user does not belong to the face plane, the wall sliding velocity really taken into account will be different.*
  - For scalars, two kinds of boundary conditions can be defined:

---

<sup>21</sup>except with Lagrangian

↪ Imposed value at the wall. The user must write

```
icodcl(ifac,ivar)=5
rcodcl(ifac,ivar,1)=imposed value
```

↪ Imposed flux at the wall. The user must write

```
icodcl(ifac,ivar)=3
rcodcl(ifac,ivar,3)=flux imposed value (depending on the variable, the user
may refer to the case icodcl=3 of § 6.4.2 for the flux definition).
```

↪ If the user does not fill these arrays, the default condition is zero flux.

- If `itypfb=iparug`: rough solid wall face, impermeable and with friction.

→ the eventual moving velocity of the wall tangent to the face is given by `rcodcl(ifac,ivar,1)` (`ivar` being `iu`, `iv` or `iw`). The initial value of `rcodcl(ifac,ivar,1)` is zero for the three velocity components (and therefore must be specified only in the case of the existence of a slipping velocity).

*WARNING: the wall moving velocity must be in the boundary face plane. By security, the code uses only the projection of this velocity on the face. As a consequence, if the velocity specified by the user is not in the face plane, the wall moving velocity really taken into account will be different.*

→ The dynamic roughness must be specified in `rcodcl(ifac,iu,3)`. The values of `rcodcl(ifac,iv,3)` and `rcodcl(ifac,iw,3)` are not used.

→ For scalars, two kinds of boundary conditions can be defined:

↪ Imposed value at the wall. The user must write

```
icodcl(ifac,ivar)=6
rcodcl(ifac,ivar,1)=imposed value
rcodcl(ifac,ivar,3)=thermal roughness value
```

↪ Imposed flux at the wall. The user must write

```
icodcl(ifac,ivar)=3
rcodcl(ifac,ivar,3)=flux imposed value (for the flux definition according to
the variable, the user may refer to the case icodcl=3 of the paragraph 6.4.2).
```

↪ If the user does not complete these arrays, the default condition is zero flux.

- If `itypfb=isymet`: symmetry face (or wall without friction).

→ Nothing to be written in `icodcl` and `rcodcl`.

- If `itypfb=isolib`: free outlet face (or more precisely free inlet/outlet with forced pressure)

→ The pressure is always treated with a Dirichlet condition, calculated with the constraint  $\frac{\partial}{\partial n} \left( \frac{\partial P}{\partial \tau} \right) = 0$ . The pressure is set to  $P_0$  at the first `isolib` face met. The pressure calibration is always done on a single face, even if there are several outlets.

→ If the mass flow is coming in, the velocity is set to zero and a Dirichlet condition for the scalars and the turbulent quantities is used (or zero-flux condition if no Dirichlet value has been specified).

→ If the mass flow is going out, zero-flux condition are set for the velocity, the scalars and the turbulent quantities.

→ Nothing is written in `icodcl` or `rcodcl` for the pressure or the velocity. An optional Dirichlet condition can be specified for the scalars and turbulent quantities.

- If `itypfb=iindef`: undefined type face (non-standard case).

→ Coding is done in a non-standard way by filling both arrays `rcodcl` and `icodcl` (see § 6.4.2).



#### NOTES

- Whatever is the value of the indicator `itypfb(ifac)`, if the array `icodcl(ifac,ivar)` is modified by the user (*i.e.* filled with a non-zero value), the code will not use the default conditions for the variable `ivar` at the face `ifac`. It will take into account only the values of `icodcl` and `rcodcl` provided by the user (these arrays must then be fully completed, like in the non-standard case).

For instance, for a normal symmetry face where scalar 1 is associated with a Dirichlet condition equal to 23.8 (with an infinite exchange coefficient):

```
itypfb(ifac)=isymet
icodcl(ifac,isca(1))=1
rcodcl(ifac,isca(1),1)=23.8
```

(`rcodcl(ifac,isca(1),2)=rinfin` is the default value, therefore it is not necessary to specify a value)  
The boundary conditions for the other variables are automatically defined.

- The user can define new types of boundary faces. He only must choose a value  $N$  and to fully specify the boundary conditions (see §6.4.2). He must specify `itypfb(ifac)=N` where  $N$  range is 1 to `ntypmx` (maximum number of boundary face types), and of course different from the values `ientre`, `iparoi`, `iparug`, `isymet`, `isolib` and `iindef` (the values of these variables are given in the `paramx` module). This allows to easily isolate some boundary faces, in order for instance to calculate balances.

## 6.4.2 Coding of non-standard boundary conditions

If a face does not correspond to a standard type, the user must fill completely the arrays `itypfb`, `icodcl` and `rcodcl`. `itypfb(ifac)` is then equal to `iindef` or another value defined by the user (see note at the end of § 6.4.1). The arrays `icodcl` and `rcodcl` must be filled as follows:

- If `icodcl(ifac,ivar)=1`: Dirichlet condition at the face `ifac` for the variable `ivar`.
  - `rcodcl(ifac,ivar,1)` is the value of the variable `ivar` at the face `ifac`.
  - `rcodcl(ifac,ivar,2)` is the value of the exchange coefficient between the outside and the fluid for the variable `ivar`. An infinite value (`rcodcl(ifac,ivar,2)=rinfin`) indicates an ideal transfer between the outside and the fluid (default case).
  - `rcodcl(ifac,ivar,3)` is not used.
  - `rcodcl(ifac,ivar,1)` has the units of the variable `ivar`, *i.e.*:
    - ↪  $m/s$  for the velocity
    - ↪  $m^2/s^2$  for the Reynolds stress
    - ↪  $m^2/s^3$  for the dissipation
    - ↪  $Pa$  for the pressure
    - ↪  $^{\circ}C$  for the temperature
    - ↪  $J.kg^{-1}$  for the enthalpy
    - ↪  $^{\circ}C^2$  for temperature fluctuations
    - ↪  $J^2.kg^{-2}$  for enthalpy fluctuations
  - `rcodcl(ifac,ivar,2)` has the following units (defined in such way that when multiplying the exchange coefficient by the variable, the given flux has the same units as the flux defined below when `icodcl=3`):
    - ↪  $kg.m^{-2}.s^{-1}$  for the velocity
    - ↪  $kg.m^{-2}.s^{-1}$  for the Reynolds stress
    - ↪  $s.m^{-1}$  for the pressure
    - ↪  $W.m^{-2}.^{\circ}C^{-1}$  for the temperature
    - ↪  $kg.m^{-2}.s^{-1}$  for the enthalpy
- If `icodcl(ifac,ivar)=2`: radiative outlet at the face `ifac` for the variable `ivar`. It reads  $\frac{\partial Y}{\partial t} + C \frac{\partial Y}{\partial n} = 0$ , where  $C$  is a to be defined celerity of radiation.

EDF R&D	<b>Code_Saturne version 3.2.0 practical user's guide</b>	Code_Saturne documentation Page 75/202
---------	--	--

- `rcodcl(ifac,ivar,3)` is not used.
- `rcodcl(ifac,ivar,1)` is the flux value of `ivar` at the cell center  $I'$ , projection of the center of the adjacent cell on the straight line perpendicular to the boundary face and crossing its center, at the previous time step. It corresponds to:
- `rcodcl(ifac,ivar,2)` is CFL number based on the parameter  $C$ , the distance to the boundary  $I'F$  and the time step:  $CFL = \frac{Cdt}{I'F}$ ,
- If `icodcl(ifac,ivar)=3`: flux condition at the face `ifac` for the variable `ivar`.
  - `rcodcl(ifac,ivar,1)` and `rcodcl(ifac,ivar,2)` are not used.
  - `rcodcl(ifac,ivar,3)` is the flux value of `ivar` at the wall. This flux is negative if it is a source for the fluid. It corresponds to:
    - ↪  $-(\lambda_T + C_p \frac{\mu_t}{\sigma_T}) \underline{\nabla} T \cdot \underline{n}$  for a temperature (in  $W/m^2$ )
    - ↪  $-(\frac{\lambda_T}{C_p} + \frac{\mu_t}{\sigma_h}) \underline{\nabla} h \cdot \underline{n}$  for an enthalpy (in  $W/m^2$ ).
    - ↪  $-(\lambda_\varphi + \frac{\mu_t}{\sigma_\varphi}) \underline{\nabla} \varphi \cdot \underline{n}$  in the case of another scalar  $\varphi$  (in  $kg.m^{-2}.s^{-1}.[\varphi]$ , where  $[\varphi]$  are the units of  $\varphi$ ).
    - ↪  $-\Delta t \underline{\nabla} P \cdot \underline{n}$  for the pressure (in  $kg.m^{-2}.s^{-1}$ ).
    - ↪  $-(\mu + \mu_t) \underline{\nabla} U_i \cdot \underline{n}$  for a velocity component (in  $kg.m^{-1}.s^{-2}$ ).
    - ↪  $-\mu \underline{\nabla} R_{ij} \cdot \underline{n}$  for a  $R_{ij}$  tensor component (in  $W/m^2$ ).
- If `icodcl(ifac,ivar)=4`: symmetry condition, for the symmetry faces or wall faces without friction. This condition can only be used for velocity components ( $\underline{U} \cdot \underline{n} = 0$ ) and the  $R_{ij}$  tensor components (for other variables, a zero-flux condition type is usually used).
- If `icodcl(ifac,ivar)=5`: friction condition, for wall faces with friction. This condition can not be applied to the pressure.
  - ↪ For the velocity and (if necessary) the turbulent variables, the values at the wall are calculated from theoretical profiles. In the case of a sliding wall, the three components of the sliding velocity are given by (`rcodcl(ifac,iu,1)`, `rcodcl(ifac,iv,1)`, and `rcodcl(ifac,iw,1)`).  
*WARNING: the wall sliding velocity must belong to the boundary face plane. For safety, the code uses only the projection of this velocity on the face. Therefore, if the velocity vector specified by the user does not belong to the face plane, the wall sliding velocity really taken into account will be different.*
  - ↪ For other scalars, the condition `icodcl=5` is similar to `icodcl=1`, but with a wall exchange coefficient calculated from a theoretical law. Therefore, the values of `rcodcl(ifac,ivar,1)` and `rcodcl(ifac,ivar,2)` must be specified: see [11].
- If `icodcl(ifac,ivar)=6`: friction condition, for the rough-wall faces with friction. This condition can not be used with the pressure.
  - ↪ For the velocity and (if necessary) the turbulent variables, the values at the wall are calculated from theoretical profiles. In the case of a sliding wall, the three components of the sliding velocity are given by (`rcodcl(ifac,iu,1)`, `rcodcl(ifac,iv,1)`, and `rcodcl(ifac,iw,1)`).  
*WARNING: the wall sliding velocity must belong to the boundary face plane. For safety, the code uses only the projection of this velocity on the face. Therefore, if the velocity vector specified by the user does not belong to the face plane, the wall sliding velocity really taken into account will be different.*
  - The dynamic roughness height is given by `rcodcl(ifac,iu,3)` only.

↪ For the other scalars, the condition `icodcl=6` is similar to `icodcl=1`, but with a wall exchange coefficient calculated from a theoretical law. The values of `rcodcl(ifac,ivar,1)` and `rcodcl(ifac,ivar,2)` must therefore be specified: see [11]. The thermal roughness height is then given by `rcodcl(ifac,ivar,3)`.

- If `icodcl(ifac,ivar)=9`: free outlet condition for the velocity. This condition is only applicable to velocity components.  
If the mass flow at the face is negative, this condition is equivalent to a zero-flux condition.  
If the mass flow at the face is positive, the velocity at the face is set to zero (but not the mass flow).  
`rcodcl` is not used.

- If `icodcl(ifac,ivar)=14`: generalized symmetry boundary condition for vectors (Marangoni effect for the velocity for instance). This condition is only applicable to vectors and set a Dirichlet boundary condition on the normal component and a Neumann condition on the tangential components.

If the three components are `ivar1`, `ivar2`, `ivar3`, the required values are:

- `rcodcl(ifac,ivar1,1)`: Dirichlet value in the  $x$  direction.
- `rcodcl(ifac,ivar2,1)`: Dirichlet value in the  $y$  direction.
- `rcodcl(ifac,ivar3,1)`: Dirichlet value in the  $z$  direction.
- `rcodcl(ifac,ivar1,3)`: flux value for the  $x$  direction.
- `rcodcl(ifac,ivar2,3)`: flux value for the  $y$  direction.
- `rcodcl(ifac,ivar3,3)`: flux value for the  $z$  direction.

Therefore, the code automatically compute the boundary condition to impose to the normal and to the tangential components.

#### NOTE

- A standard `isolib` outlet face amounts to a Dirichlet condition (`icodcl=1`) for the pressure, a free outlet condition (`icodcl=9`) for the velocity and a Dirichlet condition (`icodcl=1`) if the user has specified a Dirichlet value or a zero-flux condition (`icodcl=3`) for the other variables.

### 6.4.3 Checking of the boundary conditions

The code checks the main compatibilities between the boundary conditions. In particular, the following rules must be respected:

- On each face, the boundary conditions of the three velocity components must belong to the same type. The same is true for the components of the  $R_{ij}$  tensor.
- If the boundary conditions for the velocity belong to the “sliding” type (`icodcl=4`), the conditions for  $R_{ij}$  must belong to the “symmetry” type (`icodcl=4`), and vice versa.
- If the boundary conditions for the velocity belong to the “friction” type (`icodcl=5` or `6`), the boundary conditions for the turbulent variables must belong to the “friction” type, too.
- If the boundary condition of a scalar belongs to the “friction” type, the boundary condition of the velocity must belong to the “friction” type, too.

In case of mistakes, if the post-processing output is activated (which is the default setting), a special error output, similar to the mesh format, is produced in order to help correcting boundary condition definitions.

EDF R&D	<i>Code_Saturne</i> version 3.2.0 practical user's guide	<i>Code_Saturne</i> documentation Page 77/202
---------	--	---

#### 6.4.4 Sorting of the boundary faces

In the code, it may be necessary to have access to all the boundary faces of a given type. To ease this kind of search, an array made of sorted faces is automatically filled (and updated at each time step):

`itrifb(nfavor)`.

`ifac=itrifb(i)` is the number of the  $i^{\text{th}}$  face of type 1.

`ifac=itrifb(i+n)` is the number of the  $i^{\text{th}}$  face of type 2, if there are  $n$  faces of type 1.

... etc.

Two auxiliary arrays of size `ntypmx` are also defined.

`idebty(ityp)` is the index corresponding to the first face of type `ityp` in the array `itrifb`.

`ifinty(ityp)` is the index corresponding to the last face of type `ityp` in the array `itrifb`.

Therefore, a value `ifac0` found between `idebty(ityp)` and `ifinty(ityp)` is associated to each face `ifac` of type `ityp=itypfb(ifac)`, so that `ifac=itrifb(ifac0)`.

If there is no face of type `ityp`, the code set

`ifinty(ityp)=idebty(ityp)-1`,

which enables to bypass, for all the missing `ityp`, the loops such as

`do ii=idebty(ityp),ifinty(ityp)`.

The values of all these indicators are displayed at the beginning of the code execution listing.

#### 6.4.5 Boundary conditions with LES

##### 6.4.5.1 Vortex method

The subroutine `usvort` allows to generate the non-stationary inlet boundary conditions for the LES by the vortex method. The method is based on the generation of vortices in the 2D inlet plane with help from the pre-defined functions. The fluctuation normal to the inlet plane is generated by a Langevin equation. It is in the subroutine `usvort` where the parameters of this method are given.

*subroutine called at each time step*

To allow the application of the vortex method, an indicator must be informed of the method in the user subroutine `cs_user_parameters.f90` (`ivrtex=1`)

The subroutine `usvort` contains 3 separate parts:

- The 1st part defines the number of inlets concerned with the vortex method (`nnttt`) and the number of vortex for each inlet (`nvort`), where `ient` represents the number of inlets.
- The 2nd part (`iappel=1`) defines the boundary faces at which the vortex method is applicable. The `irepvo` array is informed by `ient` which defines the number of inlets concerned with the vortex (essentially, the vortex method can be applied with many independant inlets).
- The 3rd section defines the main parameters of the method at each inlet. With the complexity of any given geometry, 4 cases are distinguished (the first 3 use the data file `ficvor` and in the final case only 1 initial velocity and energy are imposed.):
  - \* `icas=1`, For the outlet of a rectangular pipe; 1 boundary condition is defined for each side of the rectangle taking into account their interaction with the vortex.
  - \* `icas=2`, For the outlet of a circular pipe; the entry face is considered as a wall (as far as interaction with the vortex is concerned)
  - \* `icas=3`, For inlets of any geometry; no boundary conditions are defined at the inlet face (i.e no specific treatment on the interaction between the vortex and the boundary)
  - \* `icas=4`, similar to `icas=3` except the data file is not used (`ficvor`); the outflow parameters are estimated by the code from the global data (initial velocity, level of turbulence and dissipation), information which is supplied by the user.

When the geometry allows, cases 1 and 2 are used. Case 4 is only used if it is not possible to use the other 3.

In the first 3 cases, the 2 base vectors in the plane of each inlet must be defined (vectors **dir1** and **dir2**). The 3rd vector is automatically calculated by the code, defined as a product of **dir1** and **dir2**. **dir1** and **dir2** must be chosen imperatively to give (**cen**, **dir1**, **dir2**) an orthogonal reference of the inlet plane and so **dir3** is oriented in the entry domain. If **icas=2**, the **cen** position must be the center of gravity of the rectangle or disc.

The reference points (**cen**, **dir1**, **dir2**, **dir3**) which define the values of the variable in the **ficvor** file.

In the case where **icas=4**, the vectors **dir1** and **dir2** are generated by the code.

If **icas=1**, the boundary conditions at the rectangle's edges must be defined. They are defined in the array **iclvor**. **iclvor(ii,ient)** represents the standard boundary conditions at the edge **II** ( $1 \leq II \leq 4$ ) of the inlet **ient**. The code for the boundary conditions is as follows:

- \* **iclvor=1** for a wall
- \* **iclvor=2** for symmetry
- \* **iclvor=3** for periodicity of translation (the face corresponding to periodicity will automatically be taken as 3)

The 4 edges are numbered relative to the directions **dir1** and **dir2** as shown in Figure 30:

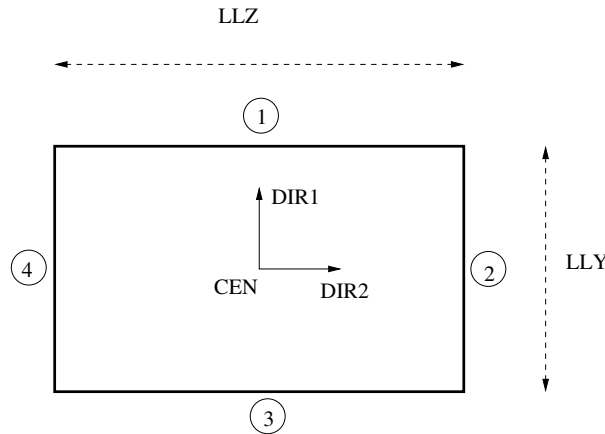


Figure 30: Numbering of the edges of a rectangular inlet(**icas=1**) treated by the vortex method

If **icas=1**, the user must define **llx** and **lly** which give the lengths of the rectangular pipe in the directions **dir1** and **dir2**.

If **icas=2**, **lld** represents the diameter of the circular pipe. If **icas=4**, **udebit**, **kdebit** and **edebit** are defined for each inlet, these give respectively, initial speed, turbulent energy level and the dissipation level. These can be used to obtain their magnitude using the correlations in the user routine **cs\_user\_boundary\_conditions** for fully developed flow in a pipe.

The independent parameters are defined as follows:

- \* **itmpl** represents the indicator of the advancement in time of the vortex. If **itmpli=1**, the vortex will be regenerated after a fixed time of **tmplim** second (defined as **itmpli=1**). If **itmpli=2**, following the data indicated in **ficvor** file, the vortex will have a variable life span equal to  $5C_\mu \frac{k^{\frac{3}{2}}}{\varepsilon U}$ , where  $C_\mu = 0,09$  and  $k$ ,  $\varepsilon$  and  $U$  represent respectively, turbulent energy, turbulent dissipation and the convective velocity in the direction normal to the inlet plane.
- \* **xsgmvo** represents the support functions used in the vortex method. They are representative of the eddy sizes entered in the vortex method. **isgmvo** is used to define their

EDF R&D	<b>Code_Saturne version 3.2.0 practical user's guide</b>	Code_Saturne documentation Page 79/202
---------	--	--

size: if **isgmvo**=1, **xsgmvo** will be constant across the inlet face and is defined in **usvort**, if **isgmvo**=2, **xsgmvo** will be variable and equal to the mixing length of the standard  $k-\varepsilon$  model ( $C_\mu^{\frac{3}{4}} \frac{k^{\frac{3}{2}}}{\varepsilon}$ ), if **isgmvo**=3, **xsgmvo** will be equal to the maximum of  $L_t$  et  $L_K$  where  $L_t$  and  $L_K$  are the  $\frac{\partial U}{\partial y} \frac{\partial U}{\partial y}$  Taylor and Kolmogorov co-efficients ( $L_T = (5\nu \frac{k}{\varepsilon})^{\frac{1}{2}}$ ,  $L_K = 200(\frac{\nu^3}{\varepsilon})^{\frac{1}{4}}$ ).

- \* **idepvo** gives the vortex displacement method in the 2D inlet plane (the vortex method is a Lagrangian method in which the eddy centres are replaced by a set velocity). If **idepvo**=1, the velocity displacement referred to by **ud** which is the vortex following a random sampling (a sample number **r**, is taken for each vortex, at each time step and for each direction and the center of the vortex is replaced by the 2 principle directions,  $\mathbf{rud}\Delta t$  where  $\Delta t$  is the time step of the calculation). If **idepvo**=2, the vortex will be convected by itself (with the speed given by the time step before the vortex method)

A data file, **ficvor**, must be defined in the cases of **icas**=1,2,3, for each inlet. The data file must contain the following data in order  $(x, y, U, \frac{\partial U}{\partial y}, k, \varepsilon)$ . The number of lines of the file is given by the integer **ndat**.  $x$  and  $y$  are the co-ordinates in the inlet plane defined by the vectors **dir1** and **dir2**.  $U$ ,  $k$  and  $\varepsilon$  are respectively, the average speed normal to the inlet, the turbulent energy and the turbulent dissipation.  $\frac{\partial U}{\partial y}$  is the derivative in the direction normal to the inlet boundary in the cases, **icas**=1, **icas**=2. Where **icas**=3 and **icas**=4 this variable is not applied (it is given the value 0) so the Langevin equations, used to generate fluctuations normal to the inlet plane, is de-activated (the fluctuations normal to the inlet is 0 on both these cases). Note that the application of many different test of the Langevin equation doesn't have a notable influence on the results and that, by contrast it simply increases the computing time per iteration and so it decreases the random sampling which slows down the pressure solver. The interpolation used in the vortex method is defined by the function **phidat**. An example is given at the end of **usvort** where the user can define the interpolation required. In the **phidat** function, **xx** and **yy** are the co-ordinates by which the value of **phidat** is calculated. **xdat** and **ydat** are the co-ordinates in the **ficvor** file. **vardat** is the value of the **phidat** function with the co-ordinates **xdat** and **ydat** (given in the **ficvor** file). Note that using an indicator **iii** accelerates the calculations (the user need not modify or delete). The user must also define the parameter **isuivo** which indicates if the vortex were started at 0 or if the file must be re-read (**ficmvo**).

## **WARNING**

- Be sure that the **ficvor** file and the interpolation in the user function **phidat** are compatible (in particular that all the entry region is covered by **ficvor**)
- If the user wants to use a 1D profile in the **dir2** direction, set  $x=0$  in the **ficvor** file and define the interpolation in **phidat**.

### **6.4.5.2 Synthetic Eddy Method**

The user file **cs\_user\_les\_inflow.f90** allows to generate the non-stationary boundary conditions for the LES by the Synthetic Eddy Method. The basic principle of this method is illustrated in figure 31: the turbulent fluctuations at the inlet are generated by a set of synthetic eddies advected across the inlet boundaries. The eddies evolve in a virtual "box" surrounding the inlet boundaries and each of them contributes to the normalized velocity fluctuations, depending on its relative position with the inlet faces and on a form function characterizing the shape of the eddies. By this way, the Synthetic Eddy Method provides a coherent flow with a target mean velocity and target Reynolds stresses at LES inlet.

**WARNING:** As for laminar or RANS inlets, the type of boundary for LES inlets is **ientre**. It has to be specified in the GUI or in the **cs\_user\_boundary\_conditions** subroutine. On the contrary, if

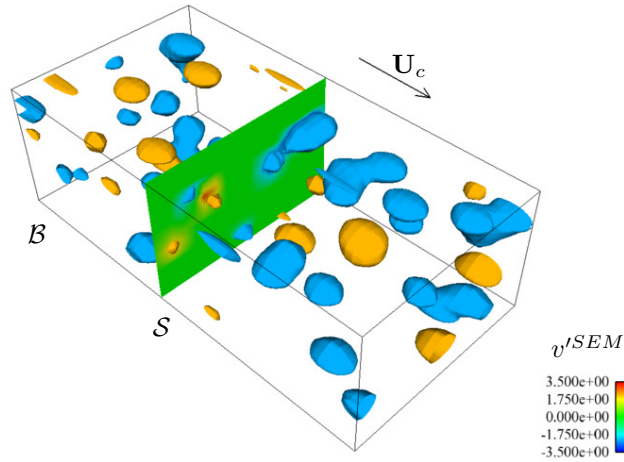


Figure 31: Illustration of the principle of the Synthetic Eddy Method, with  $\mathcal{S}$  the inlet boundary,  $\mathcal{B}$  the virtual box and  $\mathbf{U}_c$  the advection velocity of the eddies

Dirichlet values are given for these faces in the GUI or in the `cs_user_boundary_conditions` subroutine (`rcodcl(ifac,ivar,1)` array), they are erased by those provided by the Synthetic Eddy Method.

In the current version of *Code\_Saturne*, the Synthetic Eddy Method is not available through the GUI but only through the `cs_user_les_inflow.f90` user file. The user file contains 3 subroutines:

- `cs_user_les_inflow_init` (mandatory): global definition of synthetic turbulence inlets
- `cs_user_les_inflow_define` (mandatory): specific definition of each synthetic turbulence inlet
- `cs_user_les_inflow_advanced` (not mandatory): advanced definition of each synthetic turbulence inlet

`cs_user_les_inflow_init`: this subroutine defines some global parameters shared by all LES inlets. These parameters are:

- `nent`: number of LES inlet boundaries
- `isuisy`: in case of a restart calculation, it indicates if the synthetic turbulence is re-initialize (0) or read from the previous calculation (1). In that case, the checkpoint folder must contain the `les_inflow` restart file. This file is generated during a computation with synthetic turbulence, at the same physical times as the main and auxiliary restart files.

`cs_user_les_inflow_define`: this subroutine defines the specific parameters of each LES inlet. These parameters are:

- `typent`: type of LES inflow method. The Synthetic Eddy Method corresponds to `typent=3`. For the sake of comparison, other methods can be selected through this user file (see remark 2).
- `nelent`: number of synthetic eddies in the “box”. This parameter might be adjusted, depending on the case (in particular the size of the inlet plane and the level of turbulence). As a general rule, the greater is the better since an insufficient number can lead to an intermittent signal while some numerical tests have shown that this parameter does not have a great influence beyond a threshold value. Given the inlet of size  $h^2$  of a shear flow at a given Reynolds number  $Re = u_\tau h / \nu$ , an appropriate number of eddies can be evaluated by  $(Re/50)^3$  ( $Re$  and 50 approximates respectively the size, in wall unit, of the largest and the smallest synthetic eddy. Note the latter can depend on the grid size, see remark 1).



EDF R&D	<b>Code_Saturne version 3.2.0 practical user's guide</b>	Code_Saturne documentation Page 81/202
---------	--	--

- **iverbo**: level of verbosity in the listing. **iverbo=1** provides mainly informations about the size of the eddies and the size of the “box” surrounding the inlet boundary.
- **nfbent** and **lfbent**: number and list of boundary faces composing the LES inlet boundary.
- **vitent**: reference mean velocity at inlet. This parameter imposes the target mean velocity at inlet. A finer (non homogeneous) definition of the mean velocity can be done in the **cs\_user\_les\_inflow\_advanced** subroutine (see below).
- **enrent**: reference turbulence kinetic energy  $k$  at inlet. This parameter imposes the target Reynolds stresses  $R_{ij}$  at inlet, computed by  $R_{ij} = \frac{2}{3}k\delta_{ij}$  (isotropy). A finer (non isotropic and/or non homogeneous) definition of the Reynolds stresses can be done in the **cs\_user\_les\_inflow\_advanced** subroutine (see below).
- **dspent**: reference dissipation rate  $\varepsilon$  at inlet. This parameter is used to compute the size of the synthetic eddies (see remark 1). A finer (non homogeneous) definition of the dissipation rate can be done in the **cs\_user\_les\_inflow\_advanced** subroutine (see below).

**cs\_user\_les\_inflow\_advanced**: this optional subroutine enables to give an accurate (non homogeneous) specification of inflow statistics: mean velocity (**uvwent** array), Reynolds stresses (**rijent** array) and dissipation rate (**epsent** array). In that case, this accurate specification replaces the one given in **cs\_user\_les\_inflow\_define** subroutine (**vitent**, **enrent** and **dspent** variables).

**REMARK 1**: The specification of the dissipation rate  $\varepsilon$  at inlet is used to compute the size  $\sigma_i$  of the synthetic eddies in the  $i$  cartesian direction. One have:

$$\sigma_i = \max \left\{ C \frac{\left(\frac{3}{2}R_{ii}\right)^{3/2}}{\varepsilon}, \Delta \right\}, \quad C = 0.5.$$

$\Delta$  is a reference size of the grid, in order to assume that all synthetic eddies are discretized. In the implementation of *Code\_Saturne*, it is computed at each inlet boundary face  $F$  as:

$$\Delta = 2 \max_{i \leq 3, V \in \mathcal{V}} \left\{ |x_i^V - x_i^C| \right\}$$

with  $\mathcal{V}$  the subset of the vertices of the boundary face  $F$  and  $C$  the cell adjacent to  $F$ .

**REMARK 2**: For the sake of comparison, others LES inflow methods are available through the **cs\_user\_les\_inflow.f90** user file, in addition to the Synthetic Eddy Method:

- The Batten method corresponds to **typent=2** in **cs\_user\_les\_inflow\_define** subroutine. With this method, the inflow velocity signal is the superposition of several Fourier modes. The number of modes is indicated through the **nelent** keyword. As for Synthetic Eddy Method, the mean velocity, the turbulent kinetic energy and the dissipation rate have to be specified at inlet: either giving their reference values (**vitent**, **enrent** and **dspent**) in the **cs\_user\_les\_inflow\_define** subroutine, either providing an accurate local description in the **cs\_user\_les\_inflow\_advanced** subroutine.
- **typent=1**: turbulent fluctuations are given by a gaussian noise. The mean velocity and Reynolds stresses have to be specified (in **cs\_user\_les\_inflow\_define** or in **cs\_user\_les\_inflow\_advanced**). The other parameters of the user subroutines are useless. The turbulent fluctuations provided by this method are much less realistic than those provided by the Synthetic Eddy Method or the Batten method. Especially for low Reynolds number flows, this could lead to the rapid dissipation of this fluctuations and the laminarization of the flow.
- **typent=0**: No fluctuation. This method does not require any parameter. It should be reserved to regions where the flow is laminar.



## 6.5 Manage the variable physical properties

### 6.5.1 Basic variable physical properties

When the fluid properties are not constant, the user is offered the choice to define the variation laws in the Graphical User Interface (GUI) or in the subroutine `cs_user_physical_properties` which is called at each time step. In the GUI, in the item “Fluid properties” under the heading “Physical properties”, the variation laws are defined for the fluid density, viscosity, specific heat, thermal conductivity and scalar diffusivity through the use of a formula editor, see Figure 32 and Figure 33.

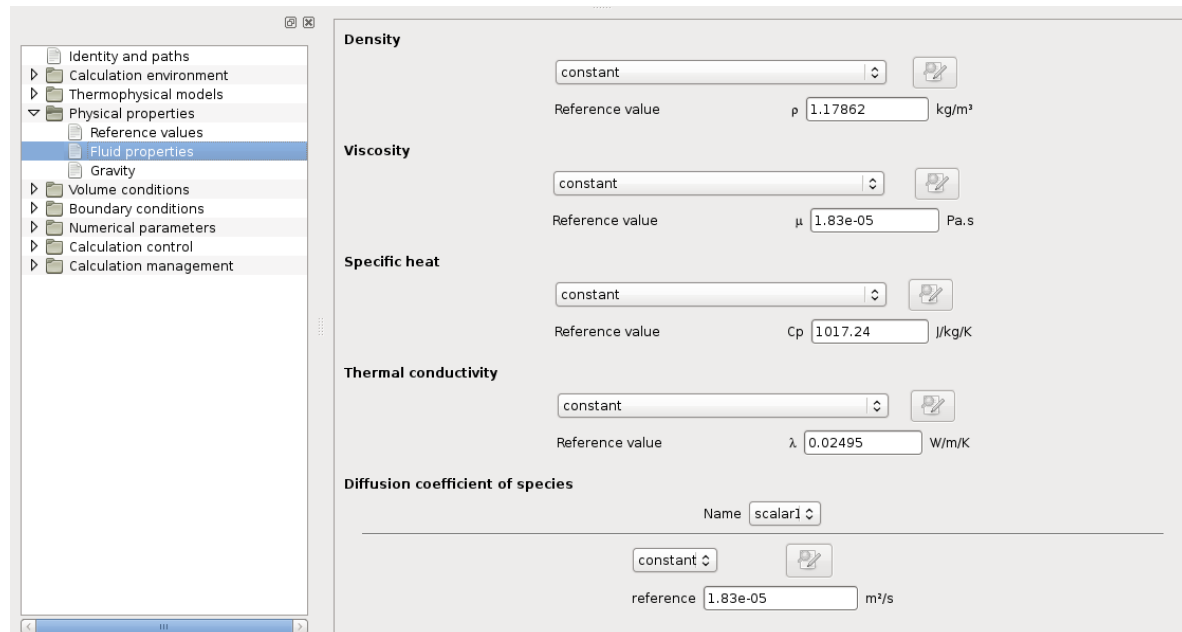


Figure 32: Physical properties - Fluid properties

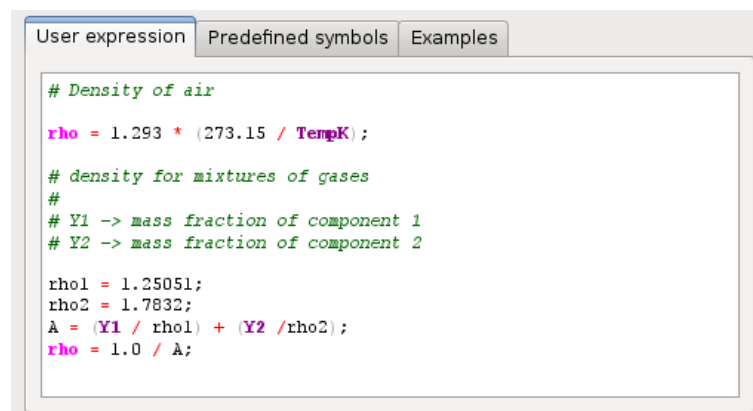


Figure 33: Definition of a user law for the density

If necessary, all the variation laws related to the fluid physical properties are written in the subroutine `cs_user_physical_properties`.

The validity of the variation laws must be checked, particularly when non-linear laws are defined (for

instance, a third-degree polynomial law may produce negative density values).

### **WARNING**

- If one wishes to impose a variable density or variable viscosity in `usphyv`, it must be flagged either in the interface or in `cs_user_parameters.f90` (`irovar=1`, `ivivar=1`).
- In order to impose a physical property ( $\rho$ ,  $\mu$ ,  $\lambda$ ,  $C_p$ )<sup>22</sup>, a reference value should be provided in the interface or in `cs_user_parameters.f90` (in particular for  $\rho$ , the pressure will be function of  $\rho_0 g z$ )
- By default, the  $C_p$  coefficient and the diffusivity for the scalars `iscal` ( $\lambda_T$  for the temperature) are considered as constant in time and uniform in space, with the values `cp0` and `visls0(iscal)` specified in the interface or in `cs_user_parameters.f90`.  
To assign a variable value to  $C_p$ , the user **must** specify it in the interface (with a user law) or assign the value 1 to `icp` in `cs_user_parameters.f90`, and fill for each cell `iel` the array `propce(iel,ipccp)` in `cs_user_physical_properties`. Completing the array `propce(iel,ipccp)` while `icp=0` induces array overwriting problems and produces wrong results.
- In the same way, to have variable diffusivities for the scalars `iscal`, the user **must** specify it in the interface (with a user law) or give the value 1 to `ivisls(iscal)` in `cs_user_parameters.f90`, and complete for each cell `iel` the array `propce(iel,ipcvsl)` in `cs_user_physical_properties`. Completing `propce(iel,ipcvsl)` while `ivisls(iscal)=0` induces memory overwriting problems and produces wrong results.

*Example:* If scalars 1 and 3 have a constant and uniform diffusivity, and if scalars 2 and 4 have a variable diffusivity, the following values must be set in `cs_user_parameters.f90` (or thanks to the GUI):

`ivisls(1)=0`, `ivisls(2)=1`, `ivisls(3)=0` and `ivisls(4)=1`.

The indicators `ivisls(2)` and `ivisls(4)` are then modified automatically by the code in order to return the rank corresponding to the diffusivity of each scalar in the list of physical properties<sup>23</sup>. The arrays `propce(iel,ipcvsl)` must then be completed with `ipcvsl=ipproc(ivisls(2))` and `ipcvsl=ipproc(ivisls(4))` (in `cs_user_physical_properties` or thanks to the GUI).

*Note:* The indicators `ivisls` must not be completed in the case of user scalars representing the average of the square of the fluctuations of another scalar, because the diffusivity of a user scalar `jj` representing the average of the square of the fluctuations of a user scalar `kk` comes directly from the diffusivity of this last scalar. In particular, the diffusivity of the scalar `jj` is variable if the diffusivity of `kk` is variable.

## **6.5.2 Modification of the turbulent viscosity**

The subroutine `usvist` is used to modify the calculation of the turbulent viscosity, *i.e.*  $\mu_t$  in  $kg.m^{-1}.s^{-1}$  (this piece of information, at the mesh cell centres, is conveyed by the variable `propce(iel,ipcvst)`, with `ipcvst = ipproc(ivisct)`). The subroutine is called at the beginning of every time step, after the calculation of the physical parameters of the flow and of the “conventional” value of  $\mu_t$  corresponding to the chosen turbulence model (indicator `iturb`).

*WARNING:* The calculation of the turbulent viscosity being a particularly sensible stage, a wrong use of `usvist` may seriously distort the results.

## **6.5.3 Modification of the variable $C$ of the dynamic LES model**

*Subroutine called every time step in the case of LES with the dynamic model.*

<sup>22</sup>except for some specific physics

<sup>23</sup>they are no longer equal to 1 but stay positive so that `ivisls>0` is synonymous with variable diffusivity

The subroutine `ussmag` is used to modify the calculation of the variable  $C$  of the LES sub-grid scale dynamic model.

Let us first remind that the LES approach introduces the notion of filtering between large eddies and small motions. The solved variables are said to be filtered in an “implicit” way. Sub-grid scale models (“dynamic” models) introduce in addition an explicit filtering.

The notations used for the definition of the variable  $C$  used in the dynamic models of *Code\_Saturne* are specified below. These notations are the ones assumed in the document [3], to which the user may refer to for more details.

The value of  $a$  filtered by the explicit filter (of width  $\widetilde{\Delta}$ ) is called  $\widetilde{a}$  and the value of  $a$  filtered by the implicit filter (of width  $\overline{\Delta}$ ) is called  $\overline{a}$ . We define:

$$\begin{aligned}\overline{S}_{ij} &= \frac{1}{2} \left( \frac{\partial \overline{u}_i}{\partial x_j} + \frac{\partial \overline{u}_j}{\partial x_i} \right) & ||\overline{S}|| &= \sqrt{2\overline{S}_{ij}\overline{S}_{ij}} \\ \alpha_{ij} &= -2\overline{\Delta}^2 ||\overline{S}|| \overline{S}_{ij} & \beta_{ij} &= -2\overline{\Delta}^2 ||\overline{S}|| \overline{S}_{ij} \\ L_{ij} &= \widetilde{\overline{u}_i \overline{u}_j} - \overline{\widetilde{u}_i \widetilde{u}_j} & M_{ij} &= \alpha_{ij} - \beta_{ij}\end{aligned}\quad (4)$$

In the framework of LES, the total viscosity (molecular + sub-grid) in  $kg.m^{-1}.s^{-1}$  may be written in *Code\_Saturne*:

$$\begin{aligned}\mu_{total} &= \mu + \mu_{sub-grid} & \text{if } \mu_{sub-grid} > 0 \\ &= \mu & \text{otherwise} \\ \text{with } \mu_{sub-grid} &= \rho C \overline{\Delta}^2 ||\overline{S}||\end{aligned}\quad (5)$$

$\overline{\Delta}$  is the width of the implicit filter, defined at the cell  $\Omega_i$  by  
 $\overline{\Delta} = XLESFL * (ALES * |\Omega_i|)^{BLES}$ .

In the case of the Smagorinsky model (`iturb=40`),  $C$  is a constant which is worth  $C_s^2$ .  $C_s^2$  is the so-called Smagorinsky constant and is stored in the variable `csmago`.

In the case of the dynamic model (`iturb=41`),  $C$  is variable in time and in space. It is determined by  

$$C = \frac{M_{ij}L_{ij}}{M_{kl}M_{kl}}.$$

In practice, in order to increase the stability, the code does not use the value of  $C$  obtained in each cell, but an average with the values obtained in the neighbouring cells (this average uses the extended neighbourhood and corresponds to the explicit filter). By default, the value calculated by the code is

$$C = \frac{\widetilde{M_{ij}L_{ij}}}{\widetilde{M_{kl}M_{kl}}}$$

The subroutine `ussmag` allows to modify this value. It is for example possible to calculate the local average after having calculated the ratio

$$C = \left[ \frac{\widetilde{M_{ij}L_{ij}}}{\widetilde{M_{kl}M_{kl}}} \right]$$

**WARNING:** The subroutine `ussmag` can be activated only when the dynamic model is used.

## 6.6 User source terms

Let us assume that the user source terms modify the equation of a variable  $\varphi$  in the following way:

$$\rho \frac{\partial \varphi}{\partial t} + \dots = \dots + S_{impl} \times \varphi + S_{expl}$$

EDF R&D	<b>Code_Saturne version 3.2.0 practical user's guide</b>	Code_Saturne documentation Page 85/202
---------	--	--

The example is valid for a velocity component, for a turbulent variable ( $k$ ,  $\varepsilon$ ,  $R_{ij}$ ,  $\omega$ ,  $\varphi$  or  $\bar{f}$ ) and for a scalar (or for the average of the square of the fluctuations of a scalar), because the syntax of all the sub-routines `ustsns`, `ustske`, `ustsri`, `ustsv2`, `ustskw`, `ustssa` and `ustssc` in the `cs_user_source_terms` file is similar.

In the finite volume formulation, the solved system is then modified as follows:

$$\left( \frac{\rho_i \Omega_i}{\Delta t_i} - \Omega_i S_{impl,i} \right) (\varphi_i^{(n+1)} - \varphi_i^{(n)}) + \dots = \dots + \Omega_i S_{impl,i} \varphi_i^{(n)} + \Omega_i S_{expl,i}$$

The user needs therefore to provide the following values:

$$\mathbf{crvimp}_i = \Omega_i S_{impl,i}$$

$$\mathbf{crvexp}_i = \Omega_i S_{expl,i}$$

In practice, it is essential for the term  $\left( \frac{\rho_i \Omega_i}{\Delta t_i} - \Omega_i S_{impl,i} \right)$  to be positive. To ensure this property, the equation really taken into account by the code is the following:

$$\left( \frac{\rho_i \Omega_i}{\Delta t_i} - \text{Min}(\Omega_i S_{impl,i}; 0) \right) (\varphi_i^{(n+1)} - \varphi_i^{(n)}) + \dots = \dots + \Omega_i S_{impl,i} \varphi_i^{(n)} + \Omega_i S_{expl,i}$$

To make the “implication” effective, the source term decomposition between the implicit and explicit parts will be done by the user who must ensure that  $\mathbf{crvimp}_i = \Omega_i S_{impl,i}$  is always negative (otherwise the solved equation remains right, but there will not be “implication”).

*Note:* Without the GUI, if the velocity components coupling algorithm is selected (`ivelco=1`), the user subroutine of the `cs_user_source_terms` file which must be filled in to add a user source term is `ustsnv` (and not `ustsns` which is only used the segregated algorithm). However, the formulation is still the same knowing that  $\mathbf{crvimp}_i$  and  $\mathbf{crvexp}_i$  are now global arrays for the three velocity components.

*WARNING:* When the second-order in time is used along with the extrapolation of the source terms<sup>24</sup>, it is no longer possible to test the sign of  $S_{impl,i}$ , because of coherence reasons (for more details, the user may refer to the theoretical and computer documentation [11] of the subroutine `preduv`). The user must therefore make sure it is always positive (or take the risk to affect the calculation stability).

#### PARTICULAR CASE OF A LINEARISED SOURCE TERM

In some cases, the added source term is not linear, but the user may want to linearise it using a first-order Taylor development, in order to make it partially implicit.

Let us consider an equation of the type:

$$\rho \frac{\partial \varphi}{\partial t} = F(\varphi)$$

We want to make it implicit using the following method:

$$\begin{aligned} \frac{\rho_i \Omega_i}{\Delta t} (\varphi_i^{(n+1)} - \varphi_i^{(n)}) &= \Omega_i \left[ F(\varphi_i^{(n)}) + (\varphi_i^{(n+1)} - \varphi_i^{(n)}) \frac{dF}{d\varphi}(\varphi_i^{(n)}) \right] \\ &= \Omega_i \frac{dF}{d\varphi}(\varphi_i^{(n)}) \times \varphi_i^{(n+1)} + \Omega_i \left[ F(\varphi_i^{(n)}) - \frac{dF}{d\varphi}(\varphi_i^{(n)}) \times \varphi_i^{(n)} \right] \end{aligned}$$

The user must therefore specify:

$$\mathbf{crvimp}_i = \Omega_i \frac{dF}{d\varphi}(\varphi_i^{(n)})$$

$$\mathbf{crvexp}_i = \Omega_i \left[ F(\varphi_i^{(n)}) - \frac{dF}{d\varphi}(\varphi_i^{(n)}) \times \varphi_i^{(n)} \right]$$

<sup>24</sup>indicator `isno2t` for the velocity, `isto2t` for the turbulence and `isso2t` for the scalars

EDF R&D	<b>Code_Saturne version 3.2.0 practical user's guide</b>	Code_Saturne documentation Page 86/202
---------	--	--

*Example:*

If the equation is  $\rho \frac{\partial \varphi}{\partial t} = -K\varphi^2$ , the user must set:

$$\begin{aligned} \text{crvimp}_i &= -2K\Omega_i \varphi_i^{(n)} \\ \text{crvexp}_i &= K\Omega_i [\varphi_i^{(n)}]^2 \end{aligned}$$

## 6.6.1 In Navier-Stokes

The source term in Navier-Stokes can be filled in thanks to the GUI or the `cs_user_source_terms` user file. Without the GUI, the subroutine `ustsns` is used to add user source terms to the Navier-Stokes equations (at each time step) when no coupling of the velocity components is considered (`ivelco=0`). When the velocity components are coupled (`ivelco=1`), the subroutine `ustsnv` which must be used.

`ustsns` is called three times every time step, once for each velocity component (`ivar` is successively `iu`, `iv` and `iw`). At each passage, the user must complete if necessary the arrays `crvimp` and `crvexp` expressing respectively the implicit and explicit part of the source term. If no other source terms apart from `ivar=iu` for example, are required, `crvimp` and `crvexp` must be read over and their 2 other components, `ivar=iv(ihpas)` and `ivar=iw` must be cancelled.

`ustsnv` is called only one time per time step and for each cell `iel`, the vector `crvexp(.,iel)` (explicit part) and the matrix `crvimp(.,.,iel)` (implicit part) must be filled in for the whole velocity vector.

## 6.6.2 For $k$ and $\varepsilon$

*Subroutine called every time step, for the  $k - \varepsilon$  and the  $v2f$  models.*

The subroutine `ustske` is used to add source terms to the transport equations related to the turbulent kinetics energy  $k$  and to the turbulent dissipation  $\varepsilon$ . This subroutine is called every time step (the treatment of the two variables  $k$  and  $\varepsilon$  is made simultaneously). The user is expected to provide the arrays `crkimp` and `crkexp` for  $k$ , and `creimp` and `creexp` for  $\varepsilon$ . These arrays are similar to the arrays `crvimp` and `crvexp` given for the velocity in the user subroutine `ustsns`. The way of making implicit the resulting source terms is the same as the one presented in `ustsns`. For  $\varphi$  and  $\bar{f}$  in the  $v2f$  model, see `ustsv2`, §6.6.4.

## 6.6.3 For $R_{ij}$ and $\varepsilon$

*Subroutine called every time step, for the  $R_{ij} - \varepsilon$  models.*

The subroutine `ustsri` is used to add source terms to the transport equations related to the Reynolds stress variables  $R_{ij}$  and to the turbulent dissipation  $\varepsilon$ . This subroutine is called 7 times every time step (once for each Reynolds stress component and once for the dissipation). The user must provide the arrays `crvimp` and `crvexp` for the variable `ivar` (referring successively to `ir11`, `ir22`, `ir33`, `ir12`, `ir13`, `ir23` and `iep`). These arrays are similar to the arrays `crvimp` and `crvexp` given for the velocity in the user subroutine `ustsns`. The method for impliciting the resulting source terms is the same as that presented in `ustsns`.

## 6.6.4 For $\varphi$ and $\bar{f}$

*Subroutine called every time step, for the  $v2f$  models.*

The subroutine `ustsv2` is used to add source terms to the transport equations related to the variables  $\varphi$  and  $\bar{f}$  of the  $v2f$   $\varphi$ -model. This subroutine is called twice every time step (once for  $\varphi$  and once for  $\bar{f}$ ). The user is expected to provide the arrays `crvimp` and `crvexp` for `ivar` referring successively to `iphi` and `ifb`. Concerning  $\varphi$ , these arrays are similar to the arrays `crvimp` and `crvexp` given for the

EDF R&D	<b>Code_Saturne version 3.2.0 practical user's guide</b>	Code_Saturne documentation Page 87/202
---------	--	--

velocity in the user subroutine **ustsns**. Concerning  $\bar{f}$ , the equation is slightly different:

$$L^2 \text{div}(\underline{\nabla}(\bar{f})) = \bar{f} + \dots + S_{impl} \times \bar{f} + S_{expl}$$

In the finite volume formulation, the solved system is written as:

$$\int_{\partial\Omega_i} \underline{\nabla}(\bar{f})^{(n+1)} dS = \frac{1}{L_i^2} \left( \Omega_i \bar{f}_i^{(n+1)} + \dots + \Omega_i S_{impl,i} \bar{f}_i^{(n+1)} + \Omega_i S_{expl,i} \right)$$

The user must then specify:

$$\text{crvimp}_i = \Omega_i S_{impl,i}$$

$$\text{crvexp}_i = \Omega_i S_{expl,i}$$

The way of making implicit the resulting source terms is the same as the one presented in **ustsns**.

### 6.6.5 For $k$ and $\omega$

*Subroutine called every time step, for the  $k - \omega$  SST model.*

The subroutine **ustskw** is used to add source terms to the transport equations related to the turbulent kinetics energy  $k$  and to the specific dissipation rate  $\omega$ . This subroutine is called every time step (the treatment of the two variables  $k$  and  $\omega$  is made simultaneously). The user is expected to provide the arrays **crkimp** and **crkexp** for the variable  $k$ , and the arrays **crwimp** and **crwexp** for the variable  $\omega$ . These arrays are similar to the arrays **crvimp** and **crvexp** given for the velocity in the user subroutine **ustsns**. The way of making implicit the resulting source terms is the same as the one presented in **ustsns**.

### 6.6.6 For $\tilde{\nu}_t$

*Subroutine called every time step, or Spalart-Allmaras model.*

The subroutine **ustssa** is used to add source terms to the transport equations related to the turbulent viscosity  $\nu_t$  for the Spalart-Allmaras model. This subroutine is called every time step. The user is expected to provide the arrays **crkimp** and **crkexp** for the variable  $\tilde{\nu}_t$ . These arrays are similar to the arrays **crvimp** and **crvexp** given for the velocity in the user subroutine **ustsns**. The way of making implicit the resulting source terms is the same as the one presented in **ustsns**.

### 6.6.7 For user scalars

*Subroutine called every time step.*

The source term in the transport equations related to the user scalars (passive or not, average of the square of the fluctuations of a scalar, ...) can be filled in thanks to the GUI or the **cs\_user\_source\_terms** user file. Without the GUI, the subroutine **ustssc** is used to add source terms to the transport equations related to the user scalars. In the same way as **ustsns**, this subroutine is called every time step, once for each user scalar. The user must provide the arrays **crvimp** and **crvexp** related to each scalar. **cvimp** and **crvexp** must be set to 0 for the scalars on which it is not wished for the user source term to be applied (the arrays are initially set to 0 at each inlet in the subroutine).

## 6.7 Pressure drops (head losses) and porosity

### 6.7.1 Head losses

Pressure drops can be defined in the Graphical User Interface (GUI) or in the subroutine **uskpdc** (called three times every time step). In the GUI, under the heading “Volume conditions”, the item “Volume regions definition” allows to define areas where pressure drops occur, see an example in fig

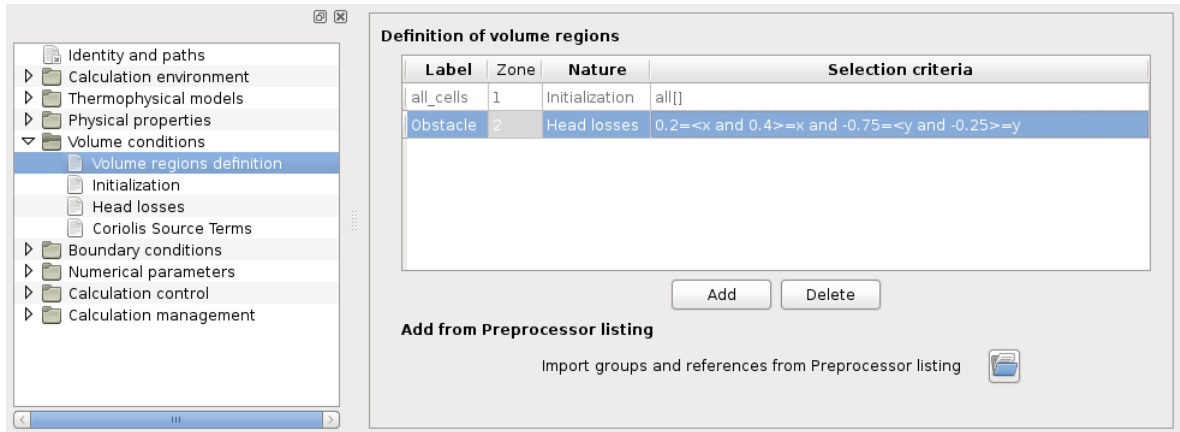


Figure 34: Creation of head losses region

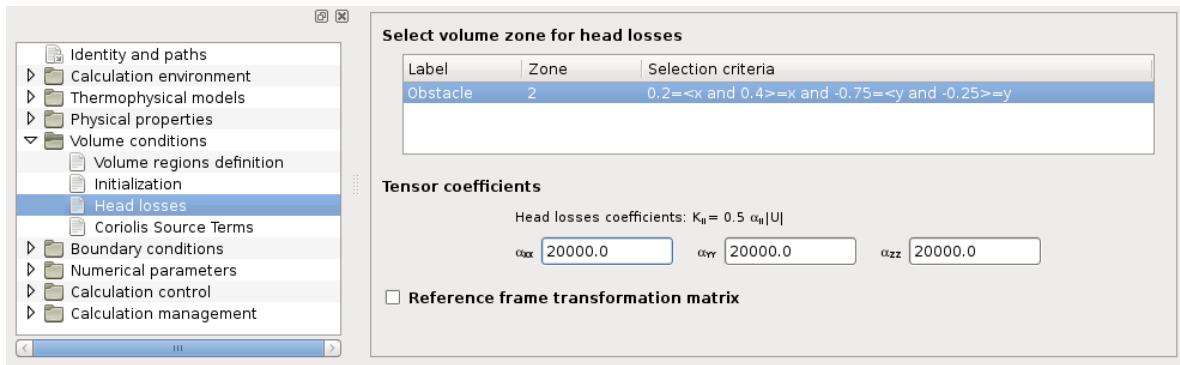


Figure 35: Head losses coefficients

34. The item “Head losses” allows to specify the head loss coefficients, see Figure 35. The tensor representing the pressure drops is supposed to be symmetric and positive.

If necessary, the pressure drops are written in the subroutine `uskpdc`.

- During the first call, all the cells are checked to know the number of cells in which a pressure drop is present. This number is called `nceptdp` in `uskpdc` (and corresponds to `nceptdc`). It is used to lay out the arrays related to the pressure drops. If there is no pressure drop, `nceptdp` must be equal to zero (it is the default value, and the rest of the subroutine is then useless).
- During the second call, all the cells are checked again to complete the array `icepdp` whose size is `nceptdp`. `icepdp(ielpdc)` is the number of the  $ielpdc^{th}$  cell containing pressure drops.
- During the third call, all the cells containing pressure drops are checked in order to complete the array containing the components of the tensor of pressure drops `ckupdp(nceptdp,6)`. This array is so that the equation related to the velocity may be written:

$$\rho \frac{\partial}{\partial t} \underline{u} = \dots - \rho \underline{K}_{pdc} \cdot \underline{u}$$

The tensor components are given in the following order (in the general reference frame): `k11`, `k22`, `k33`, `k12`, `k23`, `k13` with `k12`, `k23` and `k13` being zero if the tensor is diagonal.

The three calls are made every time step, so that variable pressure drop zones or values may be treated.

### 6.7.2 Porosity

The management of the porosity is not yet available in the GUI. To define the porosity, the user must fill in the subroutine `usporo` and set the keyword `iporos` to 1 in the `cs_user_parameters` file. This subroutine is called every time step.

When using the subroutine `usporo`, the user is expected to fill in the array `porosi` for each cell in order to give the porosity.

## 6.8 Management of the mass sources

The subroutine `ustsma` is used to add a density source term in some cells of the domain (called at each time step). The mass conservation equation is then modified as follows:

$$\frac{\partial \rho}{\partial t} + \text{div}(\rho \underline{u}) = \Gamma$$

$\Gamma$  is the mass source term expressed in  $kg.m^{-3}.s^{-1}$ .

The presence of a mass source term modifies the evolution equation of the other variables, too. Let  $\varphi$  be a any solved variable apart from the pressure (velocity component, turbulent energy, dissipation, scalar, ...). Its evolution equation becomes:

$$\rho \frac{\partial \varphi}{\partial t} + \dots = \dots + \Gamma(\varphi_i - \varphi)$$

$\varphi_i$  is the value of  $\varphi$  associated with the mass entering or leaving the domain. After discretisation, the equation may be written:

$$\rho \frac{\varphi^{(n+1)} - \varphi^{(n)}}{\Delta t} + \dots = \dots + \Gamma(\varphi_i - \varphi^{(n+1)})$$

For each variable  $\varphi$ , there are two possibilities:

- We can consider that the mass is added (or removed) with the ambient value of  $\varphi$ . In this case  $\varphi_i = \varphi^{(n+1)}$  and the equation of  $\varphi$  is not modified.
- Or we can consider that the mass is added with an imposed value  $\varphi_i$  (this solution is physically correct only when the mass is effectively added,  $\Gamma > 0$ ).

This subroutine is called three times every time step.

- During the first call, all the cells are checked to know the number of cells containing a mass source term. This number is called `ncesmp` in `ustsma` (and corresponds to `ncetsm`). It is used to lay out the arrays related to the mass sources. If there is no mass source, `ncesmp` must be equal to zero (it is the default value, and the rest of the subroutine is then useless).
- During the second call, all the cells are checked again to complete the array `icetsm` whose dimension is `ncesmp`. `icetsm(ieltsm)` is the number of the `ieltsm`<sup>th</sup> cell containing a mass source.
- During the third call, all the cells containing mass sources are checked in order to complete the arrays `itypsm(ncesmp,nvar)` and `smacel(ncesmp,nvar)`:
  - `itypsm(ieltsm,ivar)` is the flow type associated with the variable `ivar` in the `ieltsm`<sup>th</sup> cell containing a mass source.
  - `itypsm=0`:  $\varphi_i = \varphi^{(n+1)}$  condition



itypsm=1: imposed  $\varphi_i$  condition  
itypsm is not used for ivar=ipr  
- smacel(ieltsm,ipr) is the value of the mass source term  $\Gamma$ , in  $kg.m^{-3}.s^{-1}$ .  
- smacel(ieltsm,ivar), for ivar different from ipr, is the value of  $\varphi_i$  for the variable ivar in the ielstm<sup>th</sup> cell containing a mass source.

#### NOTES

- If itypsm(ieltsm,ivar)=0, smacel(ieltsm,ivar) is not used.
- If  $\Gamma = \text{smacel}(\text{ieltsm}, \text{ipr}) < 0$ , mass is removed from the system, and Code\_Saturne considers automatically a  $\varphi_i = \varphi^{(n+1)}$  condition, whatever the values given to itypsm(ieltsm,ivar) and smacel(ieltsm,ivar) (the extraction of a variable is done at ambient value).

The three calls are made every time step, so that variable mass source zones or values may be treated.

For the variance, do not take into account the scalar  $\varphi_i$  in the environment where  $\varphi \neq \varphi_i$  generates a variance source.

## 6.9 User law editor of the GUI

A formula interpreter is embedded in Code\_Saturne, which can be use throught the GUI. In order to call the formula editor of the GUI, click on the button:



The formula editor is a window with three tab:

- User expression

This tab is the formula editor. At the opening of the window only the required symbols are display. The syntax colourization shows to the user symbols which are required symbols, functions, or user variables. Each expression must be closed by a semicolon (“;”). The required symbols must be present in the final user law. A syntax checker is used when the user click on the OK button.

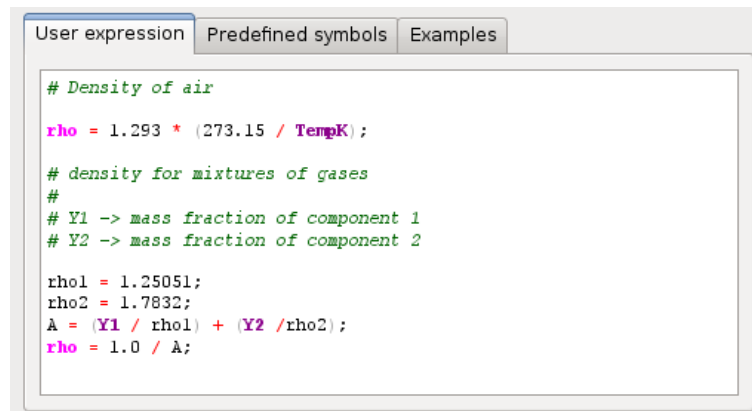


Figure 36: Example of the user law editor

- Predefined symbols

There are three types of symbol

Useful functions:

**cos**: cosine

**sin**: sine

**tan**: tangent

**exp**: exponential

**sqrt**: square root

**log**: napierian logarithm

**acos**: arc cosine

**asin**: arc sine

**atan(x)**: arc tangent (arc tangent of x in radians; the return value is in the range  $[-\pi/2, \pi/2]$ )

**atan2(y,x)**: arc tangent (arc tangent of y/x in radians; the return value is in the range  $[-\pi, \pi]$ )

**cosh**: hyperbolic cosine

**sinh**: hyperbolic sine

**tanh**: hyperbolic tangent

**abs**: absolute value

**mod**: modulo

**int**: floor

**min**: minimum

**max**: maximum

**interp1d**: 1D linear interpolation

Useful constants:

**pi** = 3.14159265358979323846

**e** = 2.718281828459045235

Operators and statements:

**+**    **-**    **\***    **/**    **^**

**!**    **<**    **>**    **<=**    **>=**    **==**    **!=**    **&&**    **||**

**while if else print**

- Examples

This tab displays examples of formula, which could be copy and paste.

External data

Through the predefined function **interp1d** it is possible to use external data inside a user law. This function allows to make linear interpolation from a data file (**.dat** or **.csv**) by selecting two columns in this file for the abscissa and ordinate of a function. In a "**.dat**" data file, tabular values are stored in plain-text form and are separated by a blank. In a comma-separated values ("**.csv**") file, values are separated by a comma **,**. For both format, lines beginning by **#** are considered as a commentary.

To use this function:

**interp1d(file.csv, i, j, x)**

with :

- file.csv, the name of the file
- i, the column number for the abscissa
- j, the column number for the ordinate
- x, the number for which the linear interpolation is looked

The function return a linear interpolation for the x value if it is in the range of the abscissa definition, or a linear extrapolation if it is out of this range.

NB : the abscissa value must be in the growing order.

## 7 Results analysis

### 7.1 Definition of post-processing and mesh zones

The functions defined in `cs_user_postprocess.c`, namely `cs_user_postprocess_writers`, `cs_user_postprocess_meshes`, and `cs_user_postprocess_activate` allow for the definition of post-processing output formats and frequency, and for the definition of surface or volume sections, in the form of lists of `nlfac` internal faces (`lstfac`) and `nlfab` boundary faces (`lstfab`), or of `nlcel` cells (`lstcel`), in order to generate chronological outputs in *EnSight*, *MED* or *CGNS* format.

One or several writers can be associated with each post-processing mesh, or “part” created. The arguments of the function `cs_post_define_writer` are as follows:

- **writer\_id**: id the the associated writer.  
negative ids are reserved (-1 for the main output), but the matching writer's options may be redefined by calls to this function.
- **case\_name**: basic name of the associated case.  
*WARNING*: depending on the chosen format, this name may be shortened (maximum number of characters: 32 for *MED*, 19 for *EnSight*) or modified automatically (white-spaces or forbidden characters will be replaced by '\_')
- **dir\_name**: name of the output directory
- **fmt\_name**: choice of the output format:
  - *EnSight Gold* (*EnSight* also accepted)
  - *MED*
  - *CGNS*
  - *CCM* (only for the full volume and boundary meshes)

The options are not case-sensitive, so *EnSight* or *CGNS* are valid, too.

- **fmt\_opts**: character string containing a list of options related to the format, separated by commas; for the *EnSight Gold* format, these options are:
  - *binary* for a binary format version (for *EnSight*, default)
  - *big\_endian* to force outputs to be in *big-endian* mode (for *EnSight*).
  - *text* for a text format version (for *EnSight*).
  - *adf* for ADF file type (for *CGNS*).
  - *hdf5* for HDF5 file type (for *CGNS*, normally the default if HDF5 support is available).

- *discard\_polygons* to prevent from exporting faces with more than four edges (which may not be recognized by some post-processing tools); such faces will therefore not appear in the post-processing mesh.
- *discard\_polyhedra* to prevent from exporting elements which are neither tetrahedra, prisms, pyramids nor hexahedra (which may not be recognized by some post-processing tools); such elements will therefore not appear in the post-processing mesh.
- *divide\_polygons* to divide faces with more than four edges into triangles, so that any post-processing tool can recognize them
- *divide\_polyhedra* to divide elements which are neither tetrahedra, prisms, pyramids nor hexahedra into simpler elements (tetrahedra and pyramids), so that any post-processing tool can recognize them
- *split\_tensor* to export the components of a tensor variable as a series of independent variables (a variable is recognised as a tensor if its dimension is 6 or 9); not implemented yet.
- **time\_dep**: indicates if the post-processing (i.e. visualization) meshes (or “parts”) are:
  - **FVM\_WRITER\_FIXED\_MESH** fixed (usual case)
  - **FVM\_WRITER\_TRANSIENT\_COORDS** deformable (the vertex positions may vary over time)
  - **FVM\_WRITER\_TRANSIENT\_CONNECT** modifiable (the lists of cells or faces defining these meshes can be changed over time).
- **output\_at\_end**: force output at calculation end if not 0
- **frequency\_n**: default output frequency in time steps associated with this writer, or  $< 0$  (the output may be forced or prevented at any time step using the function `cs_user_postprocess_activate`)
- **frequency\_t**: default output frequency in seconds associated with this writer, or  $< 0$  (has priority over **frequency\_n**, and the output may be forced or prevented at any time step using the function `cs_user_postprocess_activate`)

In order to allow the user to add an output format to the main output format, or to add a mesh to the default output, the lists of standard and user meshes and writers are not separated. Negative numbers are reserved for the non-user items. For instance, the mesh numbers -1 and -2 correspond respectively to the global mesh and to boundary faces, generated by default, and the writer -1 corresponds to the default post-processing writer.

The user chooses the numbers corresponding to the post-processing meshes and writers he wants to create. These numbers must be positive integers. It is possible to associate a user mesh with the standard post-processing case (-1), or to ask for outputs regarding the boundary faces (-2) associated with a user writer.

For safety, the output frequency and the possibility to modify the post-processing meshes are associated with the writers rather than with the meshes. This logic avoids unwanted generation of inconsistent post-processing outputs. For instance, *EnSight* would not be able to read a case in which one field is output to a given part every 10 time steps, while another field is output to the same part every 200 time steps.

The possibility to modify a mesh over time is limited by the most restrictive writer which is associated with. For instance, if writer 1 allows the modification of the mesh topology (argument **time\_dep** = **FVM\_WRITER\_TRANSIENT\_CONNECT** in the call to `cs_post_define_writer`) and writer 2 allows no modification (**time\_dep** = **FVM\_WRITER\_FIXED\_MESH**), a user post-processing mesh associated with the writers 1 and 2 will not be modifiable, but a mesh associated only with the writer 1 will be modifiable. The modification may be done by using the advanced `cs_post_define_volume_mesh_by_func()` or `cs_post_define_surface_mesh_by_func()`, associated with a user-defined selection function based on time-varying criteria (such as field values being above a given threshold). If the **time\_dep** argument is

set to **true**, the mesh will be redefined using the selection function at each output time step for every modifiable mesh.

It is possible to output variables which are normally automatically output on the main volume or boundary meshes to a user mesh which is a subset of one of these by setting the **auto\_variables** argument of one of the **cs\_post\_define...** mesh to **true**.

It is also possible to define an alias of a post-processing mesh. An alias shares all the attributes of its parent mesh (without duplication), except its number. This may be used to output different variables on a same mesh with 2 different writers: the choice of output variables is based on the mesh, so if  $P_a$  is associated with writer  $W_a$ , all that is needed is to define an alias  $P_b$  to  $P_a$  and associate it with writer  $W_b$  to allow a different output variable selection with each writer. An alias may be created using the **pstalm** subroutine.

Modification of a post-processing mesh or its alias over time is always limited by the most restrictive "writer" to which its meshes have been associated (parts of the structures being shared in memory). It is possible to define as many aliases as are required for a true mesh, but an alias cannot be defined for another alias.

It is not possible to mix cells and faces in the same mesh (most of the post-processing tools being perturbed by such a case)<sup>25</sup>.

For a better understanding, the user may refer to the examples given in **cs\_user\_postprocess\_meshes**. We can note that the white-spaces in the beginning or in the end of the character strings given as arguments of the functions called are suppressed automatically.

The additional variables to post-process on the defined meshes will be specified in the subroutine **usvpst** in the **cs\_user\_postprocess\_var.f90** file.

*WARNING In the parallel case, some meshes may not contain any local elements on a given processor. This is not a problem at all, as long as the mesh is defined for all processors (empty or not). It would in fact not be a good idea at all to define a post-processing mesh only if it contains local elements, global operations on that mesh would become impossible, leading to probable deadlocks or crashes.*

### 7.1.1 Management of the post-processing intermediate outputs

By default, a post-processing frequency is defined for each writer, as defined using the GUI or through the **The cs\_user\_postprocess\_writers** function of the **cs\_user\_postprocess.c** file. For each writer, the user may define if an output is automatically generated at the end of the calculation, even if the last time step is not a multiple of the required time step number of physical time.

For finer control, the **cs\_user\_postprocess\_activate** function of the **cs\_user\_postprocess.c** file may be used to specify when post-processing outputs will be generated.

For example, a user who wants to generate post-processing outputs (also called "chronological outputs") at the time step number 36 and around the physical time  $t=12$  seconds may use the following test:

```
if (nt_max_abs == 36) {
    int writer_id = 0;
    cs_post_activate_writer(writer_id, false);
}
/* If the current time step is the 36th */
/* all writers. */
/* activate writers. */
/* End of the test on the physical time. */
```

<sup>25</sup>actually, faces adjacent to selected cells and belonging to face or cell groups may be selected when the **add\_groups** of **cs\_post\_define...** mesh is set to **true**, so as to maintain group information, but those faces will only be written for formats supporting this (such as MED), and will only bear groups, not variable fields

## 7.2 Definition of the variables to post-process

For the mesh parts defined using the GUI or in `cs_user_postprocess.c`, the `usvpst` subroutine of `cs_user_postprocess_var.f90` file may be used to specify the variables to post-process (called for each “part”, at every active time step of an associated “writer”, see `cs_user_postprocess.c`).

The output of a given variable is generated by means of a call to the subroutine `post_write_var`, whose arguments are:

- **nummai**: current “part” number (input argument in `usvpst`).
  - **namevr**: name to give to the variable.
  - **idimt**: dimension of the variable (3 for a vector, 1 for a scalar).
  - **ientla**: indicates if the stored arrays are “interlaced” or not:
    - 0: not interlaced, in the form  $\{x_1, x_2, \dots, x_n, y_1, y_2, \dots, y_n, z_1, z_2, \dots, z_n\}$  (case of all variables defined in `rtp`).
    - 1: interlaced, in the form  $\{x_1, y_1, z_1, x_2, y_2, z_2, \dots, x_n, y_n, z_n\}$  (case of the geometric parameters, like `xyzcen`, `surfbo`, ...).
- For a scalar variable, this argument does not matter.
- **ivarpr**: indicates if the variable is defined on the “parent” mesh or locally:
    - 0: variable generated by the user in the given work arrays `tracel`, `trafac`, and `trafbr` (whose size is respectively the number of cells, internal faces and boundary faces of the “part”,  $\times 3$ ). The arrays `lstcel`, `lstfac`, and `lstfbr` can be used to get the numbers corresponding to the cells, internal faces and boundary faces associated with the “part” and to generate the appropriate post-processing variable.
    - 1: variable already defined in the main mesh (“parent” mesh of the “parts”), for example the variables in the `rtp` array. Instructions in the report which list `lstcel`, `lstfac`, and `lstfbr` will be treated directly by the sub routine, avoiding unused copies and simplifying the code
  - **ntcabs**: absolute current time step number. If a negative value is given (usually -1), the variable will be regarded as time-independent (and we will have to make sure this call is only made once).
  - **ttcabs**: current physical time value. It is not taken into account if `ntcabs` < 0.
  - **tracel**: array containing the values of the variable at the cells. If `ivarpr` = 1, this argument will be replaced by the position of the beginning of the array on which the variable is defined, for instance `rtp(1, iu(1))` for the velocity.
  - **trafac**: equivalent of `tracel` for the internal faces.
  - **trafbr**: equivalent of `tracel` for the boundary faces.

The user may refer to the example, which presents the different ways of generating an output of a variable.

*Note: To generate outputs of different variables on the same mesh with different frequencies, it is recommended to create an alias of this mesh and to associate it with a different “writer” using the GUI or in `cs_user_postprocess.c`.*

## 7.3 Modification of the variables at the end of a time step

The subroutine `cs_user_extra_operations` is called at the end of every time step. It is used to print or modify any variable at the end of every time step.

Several examples are given in the directory `EXAMPLES`:

- Calculation of a thermal balance at the boundaries and in the domain (including the mass source terms)
- Modification of the temperature in a given area starting from a given time
- Extraction of a 1D profile (which is also possible with the GUI, see Figure 27)
- Printing of a moment
- Usage of utility subroutines in the case of a parallel calculation (calculation of a sum on the processors, of a maximum, ...)

*WARNING: As all the variables (solved variables, physical properties, geometric parameters) can be modified in this subroutine, a wrong use may distort totally the calculation.*

The thermal balance example is particularly interesting.

- It can be easily adapted to another scalar (only three simple modifications to do, as indicated in the subroutine).
- It shows how to make a sum on all the sub-domains in the framework of a parallel calculation (see the calls to the subroutines `par*`).
- It shows the precautions to take before doing some operations in the framework of periodic or parallel calculations (in particular when we want to calculate the gradient of a variable or to have access to values at the neighbouring cells of a face).
- Finally it must not be forgotten that the resolution with temperature (and not enthalpy) as a solved variable is questionable when the specific heat is not constant.

## 7.4 Non-standard management of the chronological record files

The interface and the subroutine `cs_user_parameters.f90` allow to manage the “automatic” chronological record files in an autonomous way: position of the probes, printing frequency and related variables. The results are written in a different file for each variable. These files are written in text (readable by *xmgrace* or *gnuplot*) or CSV format and contain the profiles corresponding to every probe. This type of output format may not be well adapted if, for instance, the number of probes is too high. The subroutine `ushist`, called at each time step, allows then to personalise the output format of the chronological record files. The example in the directory works as follows:

- Positioning of the probes (only at the first passage): the index `ii` varies between 1 and the number of probes. The coordinates `xx`, `yy` and `zz` of each probe are given. The subroutine `findpt` gives then the number `icapt(ii)` of the cell center which is the closest to the defined probe.
- Opening of the output files (only at the first pass): in the example, the program opens a different file for all the `nvar` variables. `ficush(j)` contains the name of the  $J^{\text{th}}$  file and `impush(j)` its unit number (`impush` is initialised by default so that the user has at his disposal specific unit numbers and does not run the risk to overwrite an already open file).

- Writing to the files: in the version given as example, the program writes the time step number, the physical time step (based on the standard time step in the case of a variable time step) and the value of the selected variable at the different probes.
- Closing of the files (only at the last time step).

**WARNING:** The use of `ushist` neither erases nor replaces the parameters given in the interface or in `cs_user_parameters.f90`. Therefore, in the case of the use of `ushist`, and to avoid the creation of useless files, the user should set `ncapt=0` in the interface or in `cs_user_parameters.f90` to deactivate the automatic production of chronological records.

In addition, `ushist` generates supplementary result files.

## 8 Advanced modelling setup

### 8.1 Use of a specific physics

Specific physics such as dispersed phase, atmospheric flows, gas combustion, pulverised fuel combustion, electrical model and compressible model can be added by the user from the interface, or by using the subroutine `usppmo` of the `cs_user_parameters` file (called only during the calculation initialisation). With the interface, when a specific physics is activated in Figure 37, additional items or headings may appear (see for instance Sections 8.6.4 and 8.2.0.1).

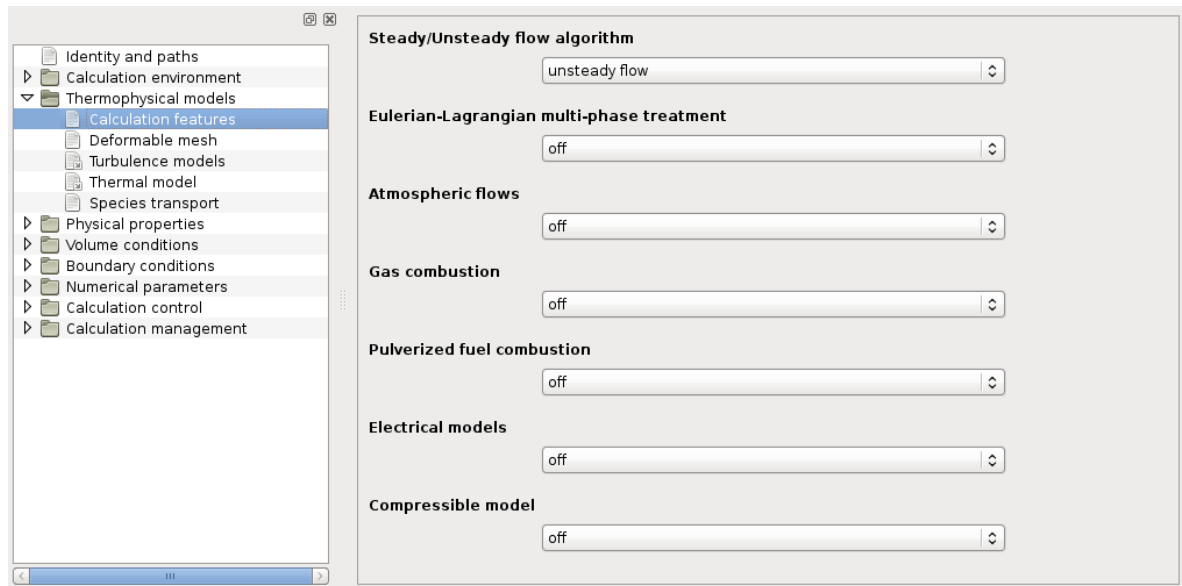


Figure 37: Specific physics models selection

When the interface is not used, `usppmo` is one of the three subroutines which must be obligatory completed by the user in order to use a specific physics module (only heavy fuel combustion is not available with the GUI). At the moment, *Code\_Saturne* allows to use two “pulverised coal” modules (with Lagrangian coupling or not) and one “pulverised heavy fuel” module, two “gas combustion” modules, two “electrical” modules, a “compressible” module and an “atmospheric” module. To activate one of these modules, the user must complete one (and only one) of the indicators `ippmod(i....)` in the subroutine `usppmo`. By default, all the indicators `ippmod(i....)` are initialised at -1, which means that no specific physics is activated.

- Diffusion flame in the framework of “3 points” rapid complete chemistry: indicator `ippmod(icod3p)`



- `ippmod(icod3p) = 0` adiabatic conditions
- `ippmod(icod3p) = 1` permeatic conditions (enthalpy transport)
- `ippmod(icod3p) = -1` module not activated
- Eddy Break Up pre-mixed flame: indicator `ippmod(icoebu)`
  - `ippmod(icoebu) = 0` adiabatic conditions at constant richness
  - `ippmod(icoebu) = 1` permeatic conditions at constant richness
  - `ippmod(icoebu) = 2` adiabatic conditions at variable richness
  - `ippmod(icoebu) = 3` permeatic conditions at variable richness
  - `ippmod(icoebu) = -1` module not activated
- Libby-Williams pre-mixed flame: indicator `ippmod(icolwc)`
  - `ippmod(icolwc)=0` two peak model with adiabatic conditions.
  - `ippmod(icolwc)=1` two peak model with permeatic conditions.
  - `ippmod(icolwc)=2` three peak model with adiabatic conditions.
  - `ippmod(icolwc)=3` three peak model with permeatic conditions.
  - `ippmod(icolwc)=4` four peak model with adiabatic conditions.
  - `ippmod(icolwc)=5` four peak model with permeatic conditions.
  - `ippmod(icolwc)=-1` module not activated.
- Multi-coals and multi-classes pulverised coal combustion: indicator `ippmod(iccoal)` The number of different coals must be less than or equal to `ncharm = 3`. The number of particle size classes `nclpch(icha)` for the coal `icha`, must be less than or equal to `ncpcmx = 10`.
  - `ippmod(iccoal) = 0` imbalance between the temperature of the continuous and the solid phases
  - `ippmod(iccoal) = 1` otherwise
  - `ippmod(iccoal) = -1` module not activated
- Multi-classes pulverised heavy fuel combustion: indicator `ippmod(icfuel)`
  - `ippmod(icfuel) = 0` module activated
  - `ippmod(icfuel) = -1` module not activated
- Lagrangian modelling of multi-coals and multi-classes pulverised coal combustion: indicator `ippmod(icpl3c)` The number of different coals must be less than or equal to `ncharm = 3`. The number of particle size classes `nclpch(icha)` for the coal `icha`, must be less than or equal to `ncpcmx = 10`.
  - `ippmod(icpl3c) = 1` coupling with the Lagrangian module, with transport of  $H_2$
  - `ippmod(icpl3c) = -1` module not activated
- Electric arcs module (Joule effect and Laplace forces): indicator `ippmod(ielarc)`
  - `ippmod(ielarc) = 1` determination of the magnetic field by means of the Ampere's theorem (not available)
  - `ippmod(ielarc) = 2` determination of the magnetic field by means of the vector potential
  - `ippmod(ielarc) = -1` module not activated
- Joule effect module (Laplace forces not taken into account): indicator `ippmod(ieljou)`
  - `ippmod(ieljou) = 1` use of a real potential

- `ippmod(ieljou) = 2` use of a complex potential
- `ippmod(ieljou) = 3` use of real potential and specific boundary conditions for transformers.
- `ippmod(ieljou) = 4` use of complex potential and specific boundary conditions for transformers.
- `ippmod(ieljou) = -1` module not activated
- Compressible module: indicator `ippmod(icompf)`
  - `ippmod(icompf) = 0` module activated
  - `ippmod(icompf) = -1` module not activated
- atmospheric flow module: indicator `ippmod(iatmos)`
  - `ippmod(iatmos) = -1` module not activated
  - `ippmod(iatmos) = 0` standard modelling
  - `ippmod(iatmos) = 1` dry atmosphere
  - `ippmod(iatmos) = 2` humid atmosphere

*WARNING: Only one specific physics module can be activated at the same time.*

In the framework of the gas combustion modelling, the user may impose his own enthalpy-temperature tabulation (conversion law). He needs then to give the value zero to the indicator `indjon` (the default value being 1). For more details, the user may refer to the following note (thermochemical files).

#### NOTE: THE THERMO-CHEMICAL FILES

The user must not forget to place in the directory DATA the thermochemical file `dp_FCP.xml`, `dp_C3P`, `dp_C3PSJ` or `dp_ELE` (depending on the specific physics module he activated) Some example files are placed in the directory DATA/REFERENCE at the creation of the study case. Their content is described below.

- Example of file for the gas combustion:
  - if the enthalpy-temperature conversion data base JANAF is used: `dp_C3P` (see array 1).
  - if the user provides his own enthalpy-temperature tabulation (there must be three chemical species and only one reaction): `dp_C3PSJ` (see array 2). This file replaces `dp_C3P`.
- Example of file for the pulverised coal combustion: `dp_FCP.xml` (see the example in the directory DATA/REFERENCE, this file can be filled in thanks to the GUI).
- Example of file for the heavy fuel combustion: `DP_FUE_new` (see the example in DATA/REFERENCE).
- Example of file for the electric arcs: `dp_ELE` (see array 3).

Lines	Examples of values	Variables	Observations
1	5	<b>ngaze</b>	Number of current species
2	10	<b>npo</b>	Number of points for the enthalpy-temperature tabulation
3	300.	<b>tmin</b>	Temperature inferior limit for the tabulation
4	3000.	<b>tmax</b>	Temperature superior limit for the tabulation
5			Empty line
6	CH4 O2 CO2 H2O N2	<b>nomcoe(ngaze)</b>	List of the current species
7	.35 .35 .35 .35 .35	<b>kabse(ngaze)</b>	Absorption coefficient of the current species
8	4	<b>nato</b>	Number of elemental species
9	.012 1 0 1 0 0	<b>wmolat(nato), atgaze(ngaze,nato)</b>	Molar mass of the elemental species (first column)
10	.001 4 0 0 2 0		Composition of the current species as a function of the elemental species ( <b>ngaze</b> following columns)
11	.016 0 2 2 1 0		
12	.014 0 0 0 0 2		
13	3	<b>ngazg</b>	Number of global species Here, <b>ngazg</b> = 3 (Fuel, Oxidiser and Products)
14	1. 0. 0. 0. 0.	<b>compog(ngaze,ngazg)</b>	Composition of the global species as a fonction of the current species of the line 6 In the order: Fuel (line 15), Oxidiser (line 16) and Product (line 17)
15	0. 1. 0. 0. 3.76		
16	0. 0. 1. 2. 7.52		
17	1	<b>nrgaz</b>	Number of global reactions Here <b>nrgaz</b> = 1 (always equal to 1 in this version)
18	1 2 -1 -9.52 10.52	<b>igfuel(nrgaz), igoxy(nrgaz), stoeg(ngazg,nrgaz)</b>	Numbers of the global species concerned by the stoichiometric ratio (first 2 integers) Stoichiometry in reaction global species. Negative for the reactants (here "Fuel" and "Oxidiser") and positive for the products (here "Products")

Table 1: Example of file for the gas combustion when JANAF is used: **dp\_C3P**

Lines	Examples of values	Variables	Observations
1	6	<b>npo</b>	Number of tabulation points
2	50. -0.32E+07 -0.22E+06 -0.13E+08	<b>th(npo),</b> <b>ehgazg(1,npo),</b> <b>ehgazg(2,npo),</b> <b>ehgazg(3,npo)</b>	Temperature(first column), mass enthalpy of fuel, oxidiser and products (columns 2,3 and 4) from line 2 to line <b>npo</b> +1
3	250. -0.68E+06 -0.44E+05 -0.13E+08		
4	450. 0.21E+07 0.14E+06 -0.13E+08		
5	650. 0.50E+07 0.33E+06 -0.12E+08		
6	850. 0.80E+07 0.54E+06 -0.12E+08		
7	1050. 0.11E+08 0.76E+06 -0.11E+08		
8	.00219 .1387 .159	<b>wmolg(1),</b> <b>wmolg(2),</b> <b>wmolg(3)</b>	Molar mass of fuel, oxidiser and products
9	.11111	<b>fs(1)</b>	Mixing rate at the stoichiometry (relating to Fuel and Oxidiser)
10	0.4 0.5 0.87	<b>ckabsg(1),</b> <b>ckabsg(2),</b> <b>ckabsg(3)</b>	Absorption coefficient of fuel, oxidiser and products
11	1. 2.	<b>xco2, xh2o</b>	Molar coefficients of $CO_2$ and $H_2O$ in the products (radiation using Modak)

Table 2: Example of file for the gas combustion when the user provides his own enthalpy-temperature tabulation (there must be three species and only one reaction): **dp\_C3PSJ** (this file replaces **dp\_C3P**)

Lines	Examples of values	Variables	Observations
1	# Fichier ASCII format libre ...		Free comment
2	# Les lignes de commentaires ...		Free comment
3	# ...		Free comment
4	# Proprietes de l'Argon ...		Free comment
5	# ...		Free comment
6	# Nb d'especes NGAZG et Nb ...		Free comment
7	# NGAZG NPO ...		Free comment
8	1 238	<b>ngazg</b> <b>npo</b>	Number of species Number of given temperature points for the tabulated physical properties ( <b>npo</b> ≤ <b>npot</b> set in <b>ppthch</b> ) So there will be <b>ngazg</b> blocks of <b>npo</b> lines each
9	# ...		Free comment
14	0	<b>ixkabe</b>	Radiation options for <b>xkabe</b>
15	# ...		Free comment
16	# Proprietes ...		Free comment
17	# T H ...		Free comment
18	# Temperature Enthalpie ...		Free comment
19	# ...		Free comment
20	# K J/kg ...		Free comment
21	# ...		Free comment
22	300. 14000. ...	<b>h</b> <b>roel</b> <b>cpel</b> <b>sigel</b> <b>visel</b> <b>xlabel</b> <b>xkabel</b>	Tabulation in line of the physical properties as a function of the temperature in Kelvin for each of the <b>ngazg</b> species Enthalpy in J/kg Density in kg/m3 Specific heat in J/(kg K) Electric conductivity in Ohm/m Dynamic viscosity in kg/(m s) Thermal conductivity in W/(m K) Absorption coefficient (radiation)

Table 3: Example of file for the electric arcs module: **dp\_ELE**

## 8.2 Pulverised coal and gas combustion module (needs update)

### 8.2.0.1 Initialisation of the variables

For coal combustion, it is possible to initialise the specific variables in the Graphical User Interface (GUI) or in the subroutine `cs_user_initialization`. In the GUI, when a coal combustion physics is selected in the item “Calculation features” under the heading “Thermophysical models”, an additional item appears: “Pulverized coal combustion”. In this item the user can define coal types, its composition, the oxydant and reactions parameters, see Figure 38 to Figure 42.

The screenshot shows the 'Coal' tab selected in the 'Coal combustion' section. It contains a table for 'Number of coal types' with two rows: 'Coal 1' and 'Coal 2'. To the right of the table are 'Add' and 'Delete' buttons. Below this is a 'Classes' section with tabs for 'Coal', 'Coke', 'Ashes', 'Devolatilisation', and 'Heterogeneous combust'. The 'Coal' tab is selected, showing a table for 'Number of classes' with two rows: 'Class 1' with an 'Initial diameter' of 0,002 and 'Class 2' with an 'Initial diameter' of 0,0001. To the right of the table are 'Add' and 'Delete' buttons.

Figure 38: Thermophysical models - Pulverized coal combustion, coal classes

The screenshot shows the 'Coke' tab selected in the 'Coke composition' section. It contains three input fields for 'Composition over C on dry', 'Composition over H on dry', and 'Composition over O on dry', each with a value of 0.0 and a '%' symbol. Below this is a 'Coke properties' section with an input field for 'PCI' with a value of 0.0 and a unit of J/kg.

Figure 39: Pulverized coal combustion, coke

If the user deals with gas combustion or if he (or she) does not want to use the GUI for coal combustion, the subroutine `cs_user_initialization` must be used (only during the calculation initialisation). In this section, “specific physics” will refer to gas combustion or to pulverised coal combustion.

These subroutines allow the user to initialise some variables specific to the specific physics activated

Figure 40: Pulverized coal combustion, coal composition

Figure 41: Pulverized coal combustion, reaction parameters

via `usppmo`. As usual, the user may have access to several geometric variables to discriminate between different initialisation zones if needed.

*WARNING: in the case of a specific physics modelling, all the variables will be initialised here, even the potential user scalars: `cs_user_initialization` is no longer used.*

- in the case of the EBU pre-mixed flame module, the user can initialise in every cell `iel`: the mixing rate `rtp(iel,isca(ifm))` in variable richness, the fresh gas mass fraction `rtp(iel,isca(iygfm))` and the mixture enthalpy `rtp(iel,isca(ihm))` in permeatic conditions
- in the case of the rapid complete chemistry diffusion flame module, the user can initialise in every cell `iel`: the mixing rate `rtp(iel,isca(ifm))`, its variance `rtp(iel,isca(ifp2m))` and the mixture mass enthalpy `rtp(iel,isca(ihm))` in permeatic conditions

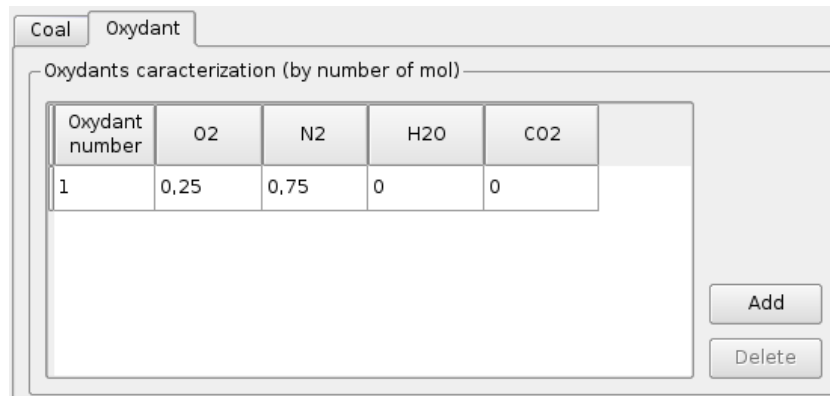


Figure 42: Pulverized coal combustion, oxydant

- in the case of the pulverised coal combustion module, the user can initialise in every cell `iel`:
  - the transport variables related to the solid phase
    - `rtp(iel,isca(ixch(icla)))` the reactive coal mass fraction related to the class `icla` (`icla` from 1 to `nclacp` which is the total number of classes, *i.e.* for all the coal type)
    - `rtp(iel,isca(ixck(icla)))` the coke mass fraction related to the class `icla`
    - `rtp(iel,isca(inp(icla)))` the number of particles related to class `icla` per kg of air-coal mixture
    - `rtp(iel,isca(ih2(icla)))` the mass enthalpy related to the class `icla` in permeatic conditions
  - `rtp(iel,isca(ihm))` the mixture enthalpy
  - the transport variables related to the gas phase
    - `rtp(iel,isca(if1m(icha)))` the mean value of the tracer 1 representing the light volatile matters released by the coal `icha`
    - `rtp(iel,isca(if2m(icha)))` the mean value of the tracer 2 representing the heavy volatile matters released by the coal `icha`
    - `rtp(iel,isca(if3m))` the mean value of the tracer 3 representing the carbon released as CO during coke burnout
    - `rtp(iel,isca(if4p2m))` the variance associated with the tracer 4 representing the air (the mean value of this tracer is not transported, it can be deduced directly from the three others)
    - `rtp(iel,isca(ifp3m))` the variance associated with the tracer 3

## 8.2.1 Boundary conditions

In this section, “specific physics” refers to gas combustion or to pulverised coal combustion. For coal combustion, it is possible to manage the boundary conditions in the Graphical User Interface (GUI). When the coal combustion physics is selected in the heading “Thermophysical models”, specific boundary conditions are activated for inlets, see Figure 43. The user fills for each type of coal previously defined (see § 8.2.0.1) the initial temperature and initial composition of the inlet flow, as well as the mass flow rate.

For gas combustion or if the GUI is not used for coal combustion, the use of `cs_user_boundary_conditions` (called at every time step) is as mandatory as `cs_user_parameters.f90` and `usppmo` to run a calculation involving specific physics. The way of using them is the same as using in the framework of standard calculations, that is, run several loops on the boundary faces lists (cf. §3.9.4) marked out by

Boundary conditions

Label	Zone	Nature	Selection criteria
wall	3	wall	5
cold_inlet	1	inlet	2
hot_inlet	2	inlet	6
outlet	5	outlet	7

Flows and temperatures

Oxydant and coal

Mass flow rate and temperature for oxydant

Norm
0.03183 m/s

Oxydant number 1
Temperature 1273.15 K

Direction
Normal to the inlet

Mass flow rate and temperature of coals

Coal number	Flow (kg/s)	Temperature (K)
Coal 1	1	1273,15
Coal 2	1	1273,15

Ratio of mass distribution for each class of coal

	Coal 1	Coal 2
Class 1	30	100
Class 2	70	

Figure 43: Boundary conditions for the combustion of coal

their colors, groups, or geometrical criterion, where the type of face, the type of boundary condition for each variable and eventually the value of each variable are defined.

*WARNING: In the case of a specific physics modelling, all the boundary conditions for every variable must be defined here, even for the eventual user scalars: `cs_user_boundary_conditions` is not used at all.*



EDF R&D	<b>Code_Saturne version 3.2.0 practical user's guide</b>	Code_Saturne documentation Page 106/202
---------	--	---

In the case of a specific physics modelling, a zone number `izone`<sup>26</sup> (for instance the color `icoul`) is associated with every boundary face, in order to gather together all the boundary faces of the same type. In comparison to `cs_user_boundary_conditions`, the main change from the user point of view concerns the faces whose boundary conditions belong to the type `itypfb=ientre`:

- for the EBU pre-mixed flame module:
    - the user can choose between the “burned gas inlet” type (marked out by the burned gas indicator `ientgb(izone)=1`) and the “fresh gas inlet” type (marked out by the fresh gas indicator `ientgf(izone)=1`)
    - for each inlet type (fresh or burned gas), a mass flow or a velocity must be imposed:
      - to impose the mass flow,
        - the user gives to the indicator `qimp(izone)` the value 1,
        - the mass flow value is set in `qimp(izone)` (positive value, in  $\text{kg s}^{-1}$ )
        - finally he imposes the velocity vector direction by giving the components of a direction vector in `rcodcl(ifac,iu)`, `rcodcl(ifac,iv)` and `rcodcl(ifac,iw)`
  - WARNING:**
    - the variable `qimp(izone)` refers to the mass flow across the whole zone `izone` and not across a boundary face (specifically for the axi-symmetric calculations, the inlet surface of the mesh must be broken up)
    - the variable `qimp(izone)` deals with the inflow across the area `izoz` and only across this zone; it is recommended to pay attention to the boundary conditions.
    - the velocity direction vector is neither necessarily normed, nor necessarily incoming.
  - to impose a velocity, the user must give to the indicator `qimp(izone)` the value 0 and set the three velocity components (in  $\text{m.s}^{-1}$ ) in `rcodcl(ifac,iu)`, `rcodcl(ifac,iv)` and `rcodcl(ifac,iw)`
  - finally he specifies for each gas inlet type the mixing rate `fment(izone)` and the temperature `tkent(izone)` in Kelvin
- for the “3 points” diffusion flame module:
  - the user can choose between the “oxydiser inlet” type marked out by `ientox(izone)=1` and the “fuel inlet” type marked out by `ientfu(izone)=1`
  - concerning the input mass flow or the input velocity, the method is the same as for the EBU pre-mixed flame module
  - finally, the user sets the temperatures `tinoxy` for each oxydiser inlet and `tinfue`, for each fuel inlet

*Note: In the standard version, only the cases with only one oxydising inlet type and one fuel inlet type can be treated. In particular, there must be only one input temperature for the oxidiser (`tinoxy`) and one input temperature for the fuel (`tinfuel`).*
- for the pulverised coal module:
  - the inlet faces can belong to the “primary air and pulverised coal inlet” type, marked out by `ientcp(izone)=1`, or to the “secondary or tertiary air inlet” type, marked out by `ientat(izone)=1`
  - in a way which is similar to the process described in the framework of the EBU module, the user chooses for every inlet face to impose the mass flow or not (`qimp(izone)=1` or 0). If the mass flow is imposed, the user must set the air mass flow value `qimpat(izone)`, its direction in `rcodcl(ifac,iu)`, `rcodcl(ifac,iv)` and `rcodcl(ifac,iw)` and the incoming air temperature `timpat(izone)` in Kelvin. If the velocity is imposed, he has to set `rcodcl(ifac,iu)`, `rcodcl(ifac,iv)` and `rcodcl(ifac,iw)`.

<sup>26</sup>`izone` must be less than the maximum number of boundary zone allowable by the code, `nozppm`. This is fixed at 2000 in `pppvar`; not to be modified

→ if the inlet belongs to the “primary air and pulverised coal” type (`ientcp(izone) = 1`) the user must also define for each coal type `icha`: the mass flow `qimpcp(izone,icha)`, the granulometric distribution `distch(izone,icha,iclacp)` related to each class `iclacp`, and the injection temperature `timpcp(izone,icha)`

## 8.2.2 Initialisation of the options of the variables

In the case of coal combustion, time averages, chronological records and listings follow-ups can be set in the Graphical User Interface (GUI) or in the subroutines `usebu1`, `usd3p1`, `uslwc1`, `uscpi1` and `uscpl1`. In the GUI, under the heading “Calculation control”, additional variables appear in the list in the items “Time averages” and “Profiles”, as well as in the item Volume solution control”, see Figure 44 and Figure 45.

Figure 44: Calculation control - Time averages

In this section, “specific physics” refers to gas combustion or pulverised coal combustion.

For gas combustion or if the GUI is not used for coal combustion, the 3 subroutines `usebu1`, `usd3p1`, `uslwc1`, `uscpi1` and `uscpl1` can be used to complete `cs_user_parameters.f90` for the considered specific physics. These subroutines are called at the calculation start. They allow to:

- generate, for the variables which are specific to the activated specific physics module, chronological outputs (indicators `ichrvr(ipp)`), follow-ups in the listings (indicator `ilisvr(ipp)`) and to activate chronological records at the probes defined in `cs_user_parameters.f90` (indicators `ihisvr(ipp)`).

The way of doing it is the same as in `cs_user_parameters.f90` and the writing frequencies of these outputs are set by `cs_user_parameters.f90`. The values of the indicators `ipp` are `ipp=ipppro(ipproc(ivar))`, with `ivar` the number of the specific physics variable. Concerning the main variables (velocity, pressure, etc ...) the user must still complete `cs_user_parameters.f90`

Solution control			
Name	Print in listing	Post-processing	Probes
Fr_HET_02	<input checked="" type="checkbox"/>	<input checked="" type="checkbox"/>	1 2 3 4
Enthalpy	<input checked="" type="checkbox"/>	<input checked="" type="checkbox"/>	1 2 3 4
NP_CP01	<input checked="" type="checkbox"/>	<input checked="" type="checkbox"/>	1 2 3 4
NP_CP02	<input checked="" type="checkbox"/>	<input checked="" type="checkbox"/>	1 2 3 4
NP_CP03	<input checked="" type="checkbox"/>	<input checked="" type="checkbox"/>	1 2 3 4
XCH_CP01	<input checked="" type="checkbox"/>	<input checked="" type="checkbox"/>	1 2 3 4
XCH_CP02	<input checked="" type="checkbox"/>	<input checked="" type="checkbox"/>	1 2 3 4
XCH_CP03	<input checked="" type="checkbox"/>	<input checked="" type="checkbox"/>	1 2 3 4
XCK_CP01	<input checked="" type="checkbox"/>	<input checked="" type="checkbox"/>	1 2 3 4
XCK_CP02	<input checked="" type="checkbox"/>	<input checked="" type="checkbox"/>	1 2 3 4
XCK_CP03	<input checked="" type="checkbox"/>	<input checked="" type="checkbox"/>	1 2 3 4
ENT_CP01	<input checked="" type="checkbox"/>	<input checked="" type="checkbox"/>	1 2 3 4
ENT_CP02	<input checked="" type="checkbox"/>	<input checked="" type="checkbox"/>	1 2 3 4
ENT_CP03	<input checked="" type="checkbox"/>	<input checked="" type="checkbox"/>	1 2 3 4
Fr_MV101	<input checked="" type="checkbox"/>	<input checked="" type="checkbox"/>	1 2 3 4
Fr_MV102	<input checked="" type="checkbox"/>	<input checked="" type="checkbox"/>	1 2 3 4
Fr_MV201	<input checked="" type="checkbox"/>	<input checked="" type="checkbox"/>	1 2 3 4
Fr_MV202	<input checked="" type="checkbox"/>	<input checked="" type="checkbox"/>	1 2 3 4
Var_AIR	<input checked="" type="checkbox"/>	<input checked="" type="checkbox"/>	1 2 3 4
Temp_GAZ	<input checked="" type="checkbox"/>	<input checked="" type="checkbox"/>	1 2 3 4
ROM_GAZ	<input checked="" type="checkbox"/>	<input checked="" type="checkbox"/>	1 2 3 4
YM_CHx1m	<input checked="" type="checkbox"/>	<input checked="" type="checkbox"/>	1 2 3 4
YM_CHx2m	<input checked="" type="checkbox"/>	<input checked="" type="checkbox"/>	1 2 3 4

Figure 45: Calculation control - Volume solution control

if he wants to get chronological records, printings in the listing or chronological outputs. The variables which can be activated by the user for each specific physics are listed below. The calculation variables `ivar` (defined at the cell `iel` by `rtp(iel,ivar)`) and the properties `iprop` (defined at the cell `iel` by `propce(iel,iproc(iprop))`) are listed now:

→ EBU pre-mixed flame modelling:

- Calculation variables `rtp(iel,ivar)`
  - `ivar = isca(iygfm)` fresh gas mass fraction
  - `ivar = isca(ifm)` mixing rate
  - `ivar = isca(ihm)` enthalpy, if transported
- Properties `propce(iel,iproc(iprop))`
  - `iprop = itemp` temperature
  - `iprop = iym(1)` fuel mass fraction
  - `iprop = iym(2)` oxidiser mass fraction
  - `iprop = iym(3)` product mass fraction
  - `iprop = ickabs` absorption coefficient, when the radiation modelling is activated

EDF R&D	<b><i>Code_Saturne</i> version 3.2.0 practical user's guide</b>	<i>Code_Saturne</i> documentation Page 109/202
---------	---	--

`iprop = it3m` and `it4m` “ $T^3$ ” and “ $T^4$ ” terms, when the radiation modelling is activated

→ rapid complete chemistry diffusion flame modelling:

everything is identical to the “EBU” case, except the fresh gas mass fraction which is replaced by the variance of the mixing rate `ivar=isca(ifp2m)`

→ pulverised coal modelling with 3 combustibles:

*variables shared by the two phases:*

- Calculation variables `rtp(iel,ivar)`
  - `ivar = isca(ihm)`: gas-coal mixture enthalpy
  - `ivar = isca(imm1)`: molar mass of the gas mixture

*variables specific to the dispersed phase:*

- Calculation variables `rtp(iel,ivar)`
  - `ivar = isca(ixck(icla))`: coke mass fraction related to the class `icla`
  - `ivar = isca(ixch(icla))`: reactive coal mass fraction related to the class `icla`
  - `ivar = isca(inp(icla))`: number of particles of the class `icla` per kg of air-coal mixture
  - `ivar = isca(ih2(icla))`: mass enthalpy of the coal of class `icla`, if we are in permeatic conditions
- Properties `propce(iel,ipproc(iprop))`
  - `iprop = imm1`: molar mass of the gas mixture
  - `iprop = itemp2(icla)`: temperature of the particles of the class `icla`
  - `iprop = irom2(icla)`: density of the particles of the class `icla`
  - `iprop = idiam2(icla)`: diameter of the particles of the class `icla`
  - `iprop = igmdch(icla)`: disappearance rate of the reactive coal of the class `icla`
  - `iprop = igmdv1(icla)`: mass transfer caused by the release of light volatiles from the class `icla`
  - `iprop = igmdv2(icla)`: mass transfer caused by the release of heavy volatiles from the class `icla`
  - `iprop = igmhet(icla)`: coke disappearance rate during the coke burnout of the class `icla`
  - `iprop = ix2(icla)`: solid mass fraction of the class `icla`

*variables specific to the continuous phase:*

- Calculation variables `rtp(iel,ivar)`
  - `ivar = isca(if1m(icha))`: mean value of the tracer 1 representing the light volatiles released by the coal `icha`
  - `ivar = isca(if2m(icha))`: mean value of the tracer 2 representing the heavy volatiles released by the coal `icha`
  - `ivar = isca(if3m)`: mean value of the tracer 3 representing the carbon released as CO during coke burnout
  - `ivar = isca(if4pm)`: variance of the tracer 4 representing the air
  - `ivar = isca(if3p2m)`: variance of the tracer 3
- Properties `propce(iel,ipproc(iprop))`
  - `iprop = itemp1`: temperature of the gas mixture
  - `iprop = iym1(1)`: mass fraction of  $CH_{X1m}$  (light volatiles) in the gas mixture
  - `iprop = iym1(2)`: mass fraction of  $CH_{X2m}$  (heavy volatiles) in the gas mixture
  - `iprop = iym1(3)`: mass fraction of CO in the gas mixture
  - `iprop = iym1(4)`: mass fraction of  $O_2$  in the gas mixture
  - `iprop = iym1(5)`: mass fraction of  $CO_2$  in the gas mixture
  - `iprop = iym1(6)`: mass fraction of  $H_2O$  in the gas mixture

`iprop = iym1(7)`: mass fraction of  $N_2$  in the gas mixture

- set the relaxation coefficient of the density `srrom`, with  
 $\rho^{n+1} = \text{srrom} * \rho^n + (1 - \text{srrom})\rho^{n+1}$   
(the default value is `srrom = 0.8`. At the beginning of a calculation, a sub-relaxation of 0.95 may reduce the numerical “shocks”).
- set the dynamic viscosity `dift10`. By default `dift10 = 4.25 kgm-1s-1` (the dynamic diffusivity being the ratio between the thermal conductivity  $\lambda$  and the mixture specific heat  $C_p$  in the equation of enthalpy).
- set the value of the constant `cebu` of the Eddy Break Up model (only in `usebu1`. By default `cebu=2.5`)

## 8.3 Heavy fuel oil combustion module

### 8.3.1 Initialisation of transported variables

To initialise or modify (in case of a continuation) values of transported variables and of the time step, the standard subroutine `cs_user_initialization` is used.

Physical properties are stored in `propce` (cell center) and `propfb`. For instance, `propce(iel, ipproc(irom))` is `rom(iel)`, the mean density (in  $kg.m^{-3}$ ).

Physical properties (`rom`, `viscl`, `cp`, ...) are computed in `ppphyv` and are not to be modified here.

The `cs_user_initialization-fuel.f90` example illustrates how the user may initialise quantities related to gaseous species and droplets compositions in addition to the chosen turbulent model.

### 8.3.2 Boundary conditions

Boundary conditions are defined as usual on a per-face basis in `cs_user_boundary_conditions`. They may be assigned in two ways:

- . for “standard” boundary conditions (inlet, free outlet, wall, symmetry): a code is defined in the array `itypfb` (of dimensions equal to the number of boundary faces). This code will then be used by a non-user subroutine to assign the conditions.
- . for “non-standard” conditions: see details given in `cs_user_boundary_conditions-fuel.f90` example.

## 8.4 Radiative thermal transfers in semi-transparent gray media

### 8.4.1 Initialisation of the radiation main parameters

The main radiation parameters can be initialise in the Graphical User Interface (GUI) or in the user subroutine `usray1`. In the GUI, under the heading “Thermophysical models”, when one of the two thermal radiative transfers models is selected, see Figure 46, additional items appear. The user is asked to choose the number of directions for angular discretisation, to define the absorption coefficient and select if the radiative calculation are restarted or not, see Figure 47 and Figure 49. When “Advanced options” is selected for both models Figure 48 or Figure 50 appear, the user must fill the resolution frequency and verbosity levels. In addition, the activation of the radiative transfer leads to the creation of an item “Surface solution control” under the heading “Calculation control”, see Figure 51, where radiative transfer variables can be selected to appear in the output listing.

If the GUI is not used, `usray1` is one of the two subroutine which must be completed by the user for all calculations including radiative thermal transfers. It is called only during the calculation initialisation.

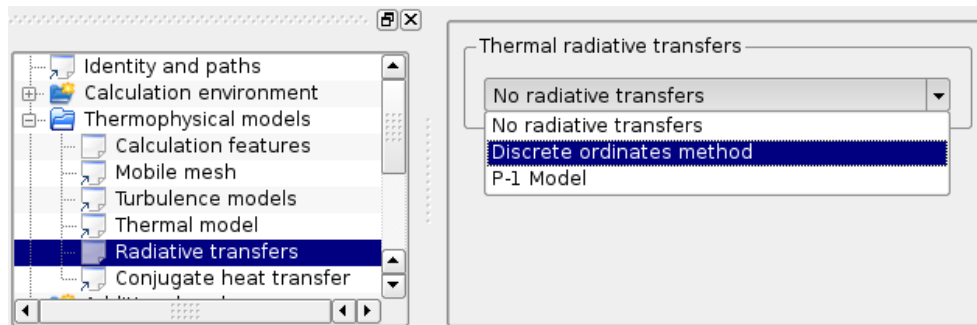


Figure 46: Radiative transfers models

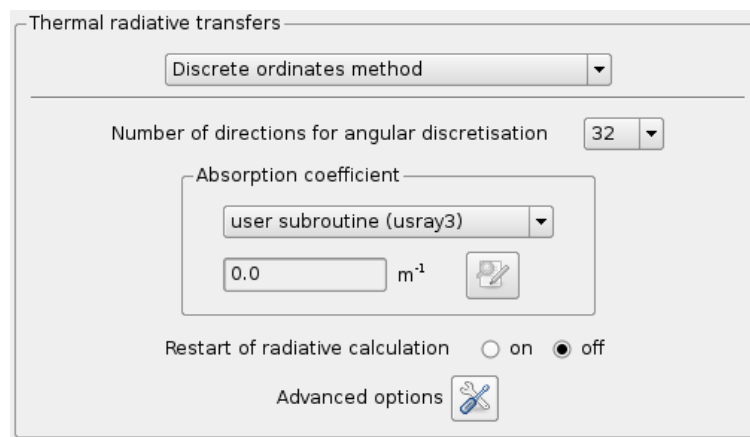


Figure 47: Radiative transfers - parameters of the DO method

It is composed of three headings. The first one is dedicated to the activation of the radiation module, only in the case of classic physics.

*WARNING: when a calculation is ran using a specific physics module, this first heading must not be completed. The radiation module is then activated or not, according to the parameter file related to the considered specific physics.*

In the second heading the basic parameters of the radiation module are indicated.

Finally, the third heading deals with the selection of the post-processing graphic outputs. The variables to treat are splitted into two categories: the volumetric variables and those related to the boundary faces.

For more details about the different parameters, the user may refer to the keyword list (§ 9).

## 8.4.2 Radiative transfers boundary conditions

These informations can be filled by the user through the Graphical User Interface (GUI) or by using the subroutine `usray2` (called every time step). If the interface is used, when one of the “Radiative transfers” options is selected in Figure 46, it activates specific boundary conditions each time a “Wall” is defined, see Figure 52. The user can then choose between 3 cases. The parameters the user must specify are displayed for one of them in Figure 53.

When the GUI is not used, `usray2` is the second subroutine necessary for every calculation which

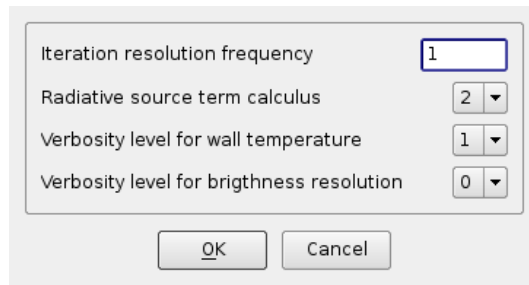


Figure 48: Radiative transfers - advanced parameters of the DO method

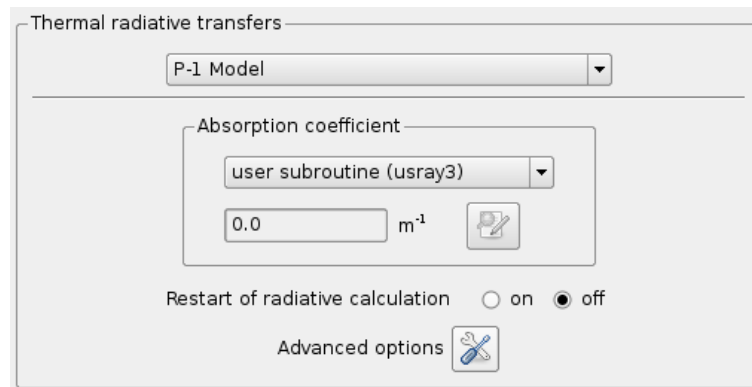


Figure 49: Radiative transfers - parameters of the P-1 model

includes radiative thermal transfers. It is used to give all the necessary parameters concerning, in the one case, the wall temperature calculation, and in the other, the coupling between the thermal scalar (temperature or enthalpy), and the radiation module at the calculation domain boundaries. It must be noted that the boundary conditions concerning the thermal scalar which may have been defined in the subroutine `cs_user_boundary_conditions` will be modified by the radiation module according to the data given in `usray2` (cf. §3.9.4).

A zone number must be given to each boundary face <sup>27</sup>and, specifically for the walls, a boundary condition type and an initialisation temperature (in Kelvin). The initialisation temperature is only used to make the solving implicit at the first time step. The zone number allows to assign an arbitrary integer to a set of boundary faces having the same radiation boundary condition type. This gathering is used by the calculation, and in the listing to print some physical values (mean temperature, net radiative flux ...). An independent graphic output in *EnSight* format is associated with each zone and allows the display on the boundary faces of the variables selected in the third heading of the subroutine `usray1`.

A boundary condition type stored in the array `ISOTHP` is associated with each boundary face. There are five different types:

- `itpimp`: wall face with imposed temperature,
- `ipgrno`: for a grey or black wall face, calculation of the temperature by means of a flux balance,
- `iprefl`: for a reflecting wall face, calculation of the temperature by means of a flux balance. This is fixed at 2000 in `radiat` and cannot be modified.
- `ifgrno`: grey or black wall face to which a conduction flux is imposed,

<sup>27</sup>this must be less than the maximum allowable by the code, `nozrdm`. This is fixed at 2000 in `radiat` and cannot be modified.

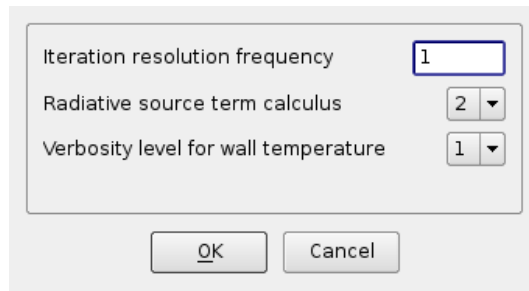


Figure 50: Radiative transfers - advanced parameters of the P-1 model

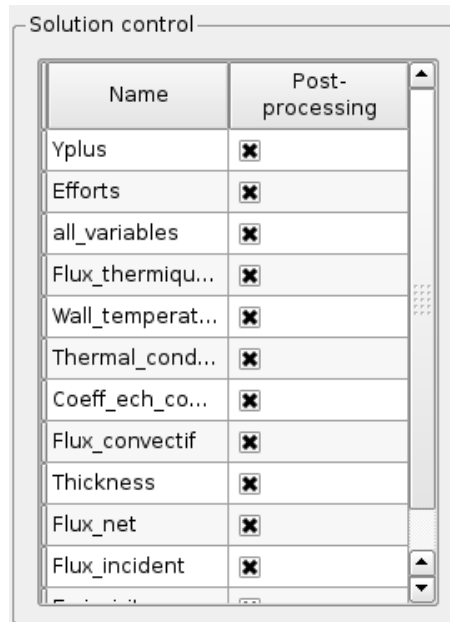


Figure 51: Calculation control - Radiative transfers post-processing output

- **ifrefl**: reflecting wall face to which a conduction flux is imposed, which is equivalent to impose this flux directly to the fluid.

Depending on the selected boundary condition type at every wall face, the code needs to be given some supplementary pieces of information:

- **itpimp**: the array **tintp** must be completed with the imposed temperature value and the array **epsp** must be completed with the emissivity value (strictly positive).
- **ipgrno**: must be given: an initialisation temperature in the array **tintp**, the wall emissivity (strictly positive, in **epsp**), thickness (in **epap**), thermal conductivity (in **xlamp**) and an external temperature (in **textp**) in order to calculate a conduction flux across the wall.
- **iprefl**: must be given: an initialisation temperature (in **tintp**), the wall thickness (in **epap**) and thermal conductivity (in **xlamp**) and an external temperature (in **textp**).
- **ifgrno**: must be given: an initialisation temperature (in **tintp**), the wall emissivity (in **epsp**) and the conduction flux (in  $W/m^2$  whatever the thermal scalar, enthalpy or temperature) in the array **rcodcl**. The value of **rcodcl** is positive when the conduction flux is directed from the



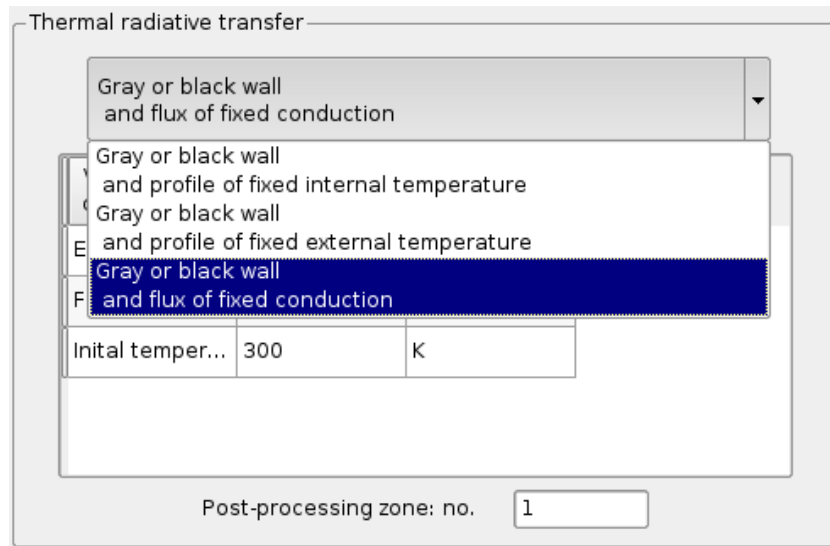


Figure 52: Boundary conditions - choice of wall thermal radiative transfers

inside of the fluid domain to the outside (for instance, when the fluid heats the walls). If the conduction flux is null, the wall is adiabatic.

- **ifrefl**: must be given: an initialisation temperature (in **tintp**) and the conduction flux (in  $W/m^2$  whatever the thermal scalar) in the array **rcodcl**. The value of **rcodcl** is positive when the conduction flux is directed from the inside of the fluid domain to the outside (for instance, when the fluid heats the walls). If the conduction flux is null, the wall is adiabatic. The flux received by **rcodcl** is directly imposed as boundary condition for the fluid.

*WARNING: it is mandatory to set a zone number to every boundary face, even those which are not wall faces. These zones will be used during the printing in the listing. It is recommended to gather together the boundary faces of the same type, in order to ease the reading of the listing.*

### 8.4.3 Absorption coefficient of the medium, boundary conditions for the luminance and calculation of the net radiative flux

When the absorption coefficient is not constant, the subroutine **usray3** is called instead at each time step. It is composed of three parts. In the first one, the user must provide the absorption coefficient of the medium in the array **CK**, for each cell of the fluid mesh. By default, the absorption coefficient of the medium is 0, which corresponds to a transparent medium.

*WARNING: when a specific physics is activated, it is forbidden to give a value to the absorption coefficient in this subroutine. In this case, it is calculated automatically, or given by the user via a thermo-chemical parameter file (**dp\_C3P** or **dp\_C3PSJ** for gas combustion, and **dp\_FCP** for pulverised coal combustion).*

The two following parts of this subroutine concern a more advanced use of the radiation module. It is about imposing boundary conditions to the equation of radiative transfer and net radiative flux calculation, in coherence with the luminance at the boundary faces, when the user wants to give it a particular value. In most cases, the given examples do not need to be modified.

Wall radiative characteristics	Value	Unit
Emissivity	0,8	
Conductivity	3	W/m/K
Thickness	0,1	m
Profile of ext...	300	K
Profile of int...	300	K

Post-processing zone: no.

Figure 53: Boundary conditions - example of wall thermal radiative transfer

## 8.4.4 Encapsulation of the temperature-enthalpy conversion

*Subroutine called every time step.*

The user subroutine `usray4` is used to call the user subroutine `usthht`. `usthht` is used to encapsulate a simple enthalpy-temperature conversion law and its inverse. The user can implement his own conversion formulas into it.

This subroutine is useless when the thermal scalar is the temperature.

*WARNING: when a specific physics is activated, it is forbidden to use this subroutine. In this case, `usray4` is replaced by `ppray4`, which is not a user subroutine.*

The value of the argument `mode` allows to know in which direction the conversion will be made:

- `mode = 1`: the fluid enthalpy in the cell must be converted into temperature (in Kelvin),
- `mode = -1`: the wall temperature (`text` or `tparoi`, in Kelvin) must be converted into enthalpy.

*WARNING: the value of `mode` is passed as argument and must not be modified by the user.*

## 8.4.5 Input of radiative transfer parameters

*The routine `usray5` is called twice. The first time is for boundary conditions. The second time is for the net radiation flux computation*

In this subroutine, during the first call (`iappel=1`), the boundary conditions are filled:

- the radiative intensity must be set in the array `cofrua` when the discrete ordinates model is used; an example is given in `usray5` for an isotropic radiation field on a grey wall. Proposed boundary conditions for the intensity in `usray5` are: symmetry, inlet/outlet, and wall boundary,
- the entering intensity for free boundaries is set to zero in `cofrua` (if the user has more information, he can improve it),

EDF R&D	<b>Code_Saturne version 3.2.0 practical user's guide</b>	Code_Saturne documentation Page 116/202
---------	--	---

- arrays `cofrua` and `cofrub` must be filled when the P-1 model is used. The boundary conditions proposed are the same as with the discrete ordinates model.

During the second call (`iappel=2`), the density of the net radiation flux must be calculated consistently with the boundary conditions of the intensity considering that the density of net flux is the balance between the radiative emitting part of a boundary face (and not the reflecting one) and the radiative absorbing part. The provided example is consistent with the example of the intensity boundary conditions given when the discrete ordinates model is used.

## 8.5 Conjugate heat transfer

### 8.5.1 Thermal module in a 1D wall

*subroutine called at every time step*

This subroutine takes into account the affected thermal inertia by a wall. Some boundary faces are treated as a solid wall with a given thickness, on which the code resolves an unidimensional equation for the heat conduction. The coupling between the 1D module and the fluid works in a similar way to the coupling with the SYRTHES. In construction, the user is not able to account for the heat transfer between different parts of the wall. A physical analysis of each problem, case by case is required to evaluate the relevance of its usage by way of a report of the simple conditions (temperature, zero-flux ) or a coupling with SYRTHES.

The use of this code requires that the thermal scalar is defined as (`iscalt > 0`).

*WARNING: The 1D thermal module is developed assuming the thermal scalar as a temperature. If the thermal scalar is an enthalpy, the code calls the subroutine `usthht` for each transfer of information between the fluid and the wall in order to convert the enthalpy to temperature and vice-versa. This function has not been tested and is firmly discouraged. If the thermal variable is the total (compressible) energy, the thermal module will not work.*

This procedure is called twice, at initialisation and again at each time step.

- The 1st call (initialisation) all the boundary faces that will be treated as a coupled wall are marked out. This figure is written noted as `nfkpt1d`. It applies dimension to the arrays in the thermal module. `nfkpt1d` will be at 0 if there are no coupled faces (it is in fact the default value, the remainder of the subroutine is not used in this case). The parameter `isuit1` also must be defined, this indicates if the temperature of the wall must be initialised or written in the file (stored in the variable `filmt1`).
- The 2nd call (initialisation) again concern the wall faces, it completes the `ifpt1d` array of dimension `nfpt1d`. `ifpt1d(ifbt1d)` is the number `ifbt1dth` boundary faces coupled with the thermal module of a 1D wall. The directional parameters are then completed for a pseudo wall associated to each face
  - `nppt1d(nfpt1d)`: number of cells in the 1D mesh associated to the pseudo wall.
  - `eppt1d(nfpt1d)`: thickness of the pseudo wall.
  - `rgpt1d(nfpt1d)`: geometry of the pseudo wall mesh (refined as a fluid if `rgt1d` is smaller than 1)
  - `tppt1d(nfpt1d)`: initialisation temperature of the wall (uniform in thickness). In the course of the calculation, the array stores the temperature of the solid at the fluid/solid interface.

Other than for re-reading a file (`ficmt1`), `tppt1d` is not used. `nppt1d`, `ifpt1d`, `rgpt1d` and `eppt1d` are compared to data from the follow-up file and they must be identical.

*WARNING: The test in `ifpt1d` implicitly assumes that the array is completed in ascending order (i.e `ifpt1d(ii)>ifpt1d(jj)` if `ii>jj`). This will be the case if the coupled faces are defined starting from the unique loop on the boundary faces (as in the example). If this is not the case, contact the development team to short circuit the test.*

- The 3rd call (at each time step) is for the confirmation that all the arrays involving physical parameter and external boundary conditions have been completed.
  - `iclt1d(nfpt1d)`: Typical boundary condition at the external (pseudo) wall: Dirichlet condition (`iclt1d=1`) or flux condition (`iclt1d=3`)
  - `tept1d(nfpt1d)`: External temperature of the pseudo wall in the Dirichlet case.
  - `hept1d(nfpt1d)`: External coefficient of transfer in the pseudo wall under Dirichlet conditions (in  $W.m^{-2}.K$ ).
  - `fept1d(nfpt1d)`: External heat flux in the pseudo wall under the flux conditions (in  $W.m^{-2}$ , negative value for energy entering the wall).
  - `xlmt1d(nfpt1d)`: Conductivity  $\lambda$  of the wall uniform in thickness (in  $W.m^{-1}.K^{-1}$ ).
  - `rcpt1d(nfpt1d)`: Volumetric heat capacity  $\rho C_p$  of the wall uniform in thickness (in  $J.m^{-3}.K^{-1}$ ).
  - `dtpt1d(nfpt1d)`: Physical time step associated with the solved 1D equation of the pseudo wall (which can be different from the time step in the calculation).

The 3<sup>rd</sup> call, done at each time step, allows to impose boundary conditions and physical values in time.

## 8.5.2 Fluid-Thermal coupling with SYRTHES

When the user wishes to couple *Code\_Saturne* with SYRTHES to include heat transfers, it can be done with the Graphical User Interface (GUI) or the `cs_syrthes_coupling` user function. In the GUI, to set such a coupling, a thermal scalar must be selected first in the item “Thermal scalar” under the heading “Thermophysical models”. Then the item “Conjugate heat transfer” will appear, see Figure 54. The zones where the coupling occurs must be defined and a projection axis can be specified in case of 2D coupling.

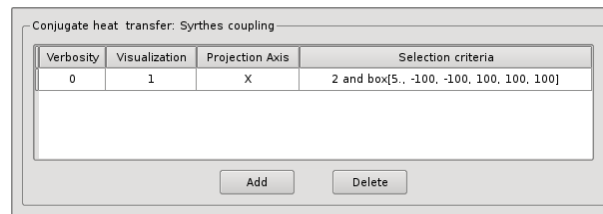


Figure 54: Thermophysical models - coupling with SYRTHES

If the function `cs_user_syrthes_coupling` is used, the user must specify the arguments passed to the '`cs_syr_coupling_define`' function. These arguments are:

- `syrthes_name` is the matching SYRTHES application name (useful only when more than one SYRTHES and one *Code\_Saturne* domain are present),
- `boundary_criteria` is the surface selection criteria,
- `volume_criteria` is the volume selection criteria,
- `projection_axis`: ' ' if the user wishes to use a 3D standard coupling, or specify 'x', 'y', or 'z' as the projection axis if a 2D coupling with SYRTHES is used,

- `verbosity` is the verbosity level.
- `visualization` is the visualization level.

Examples are provided in `cs_user_coupling.c`.

The user may also define global coupling options relative to the handling of time-stepping, by adapting the example `cs_user_coupling` in the `cs_user_coupling.c` file. In the case of multiple couplings, these options are global to all SYRTHES and *Code\_Saturne* couplings.

## 8.6 Particle-tracking (Lagrangian) Module

### 8.6.1 General information

- The particle-tracking (or Lagrangian) module enables the simulation of poly-dispersed particulate flows, by calculating the trajectories of individual particles, mainly characterized by their diameter and density (if no heat nor mass transfer between particle and fluid are activated).
- The standard use of the particle-tracking module follows the **Moments/PDF approach** : the instantaneous properties of the underlying flow needed to calculate the particle motion are reconstructed from the averaged values (obtained by Reynolds-Averaged Navier-Stokes simulation) by using stochastic processes. The statistics of interest are then obtained through Monte-Carlo simulation.
- As a consequence, it is important to emphasize that the most important (and physically meaningful) results of a particle-tracking calculation following the Moments/PDF approach are **statistics**. Volume and surface statistics, steady or unsteady, can be calculated. Individual particle trajectories (as 1D, *EnSight*-readable cases) and displacements (as *EnSight*-readable animations) can also be provided, but only for illustrative purposes.

### 8.6.2 Activating the particle-tracking module

The activation of the particle-tracking module is performed either:

- in the Graphical User Interface (GUI): Calculation features → Thermophysical models → Eulerian-Lagrangian multi-phase treatment → particles and droplets tracking
- or in the user subroutine `uslag1`, by setting the `iilag` parameter to a non-null value.

### 8.6.3 Basic guidelines for standard simulations

Except for cases in which the flow conditions depend on time, it is generally recommended to perform a first Lagrangian calculation whose aim is to reach a steady-state (i.e. to reach a time starting from which the relevant statistics do not depend on time anymore). In a second step, a calculation restart is done to calculate the statistics. When the single-phase flow is steady and the particle volume fraction is low enough to neglect the particles influence on the continuous phase behaviour, it is recommended to perform a Lagrangian calculation on a frozen field.

It is then possible to calculate steady-state volumetric statistics and to give a statistical weight higher than 1 to the particles, in order to reduce the number of simulated (“numerical”) particles to treat while keeping the right concentrations. Otherwise, when the continuous phase flow is stationary, but the two-coupling coupling must be taken into consideration, it is still possible to activate stationary statistics.

When the continuous phase flow is non-stationary, it is no longer possible to use stationary statistics. To have correct statistics at every moment in the whole calculation domain, it is imperative to have

EDF R&D	<b><i>Code_Saturne</i> version 3.2.0 practical user's guide</b>	<i>Code_Saturne</i> documentation Page 119/202
---------	---	--

an established particle seeding and it is recommended (when it is possible) not to impose statistical weights different from the unity.

Finally, when the so-called complete model is used for turbulent dispersion modelling, the user must make sure that the volumetric statistics are directly used for the calculation of the locally undisturbed fluid flow field.

When the thermal evolution of the particles is activated, the associated particulate scalars are always the inclusion temperature and the locally undisturbed fluid flow temperature expressed in degrees Celsius, whatever the thermal scalar associated with the continuous phase is (temperature or enthalpy). If the thermal scalar associated with the continuous phase is the temperature in Kelvin, the unit change is done automatically. If the thermal scalar associated with the continuous phase is the enthalpy, the enthalpy-temperature conversion subroutine `usthht` must be completed for `mode=1`, and must express temperatures in degrees Celsius. In all cases, the thermal backward coupling of the dispersed phase on the continuous phase is adapted to the thermal scalar transported by the fluid.

## 8.6.4 Prescribing the main modelling parameters (GUI and/or `uslag1`)

### USE OF THE GUI

In the GUI, the selection of the Lagrangian module activates the heading **Particle and droplets tracking** in the tree menu. The initialization is performed in the three items included in this heading:

- **Global settings.** Here, the user defines the kind of Euler/Lagrange multi-phase treatment, the main parameters, the specific physics associated with the particles and advanced numerical options, see Figure 55 to Figure 56.
- **Statistics.** Here, the user can select volume and boundary statistics to be post-processed.
- **Output.** the user defines the output frequency, post-processing options for particles and select the variables that will appear in the listing.

### USE OF THE SUBROUTINE `USLAG1`

When the GUI is not used, `uslag1` must be completed. This subroutine gathers in different headings all the keywords which are necessary to configure the Lagrangian module. The different headings refer to:

- the global configuration parameters
- the specific physical models describing the particle behaviour
- the backward coupling (influence of the dispersed phase on the continuous phase)
- the numerical parameters
- the volumetric statistics
- the boundary statistics
- the post-processing in trajectory mode

For more details about the different parameters, the user may refer to the keyword list (§ 9.8).

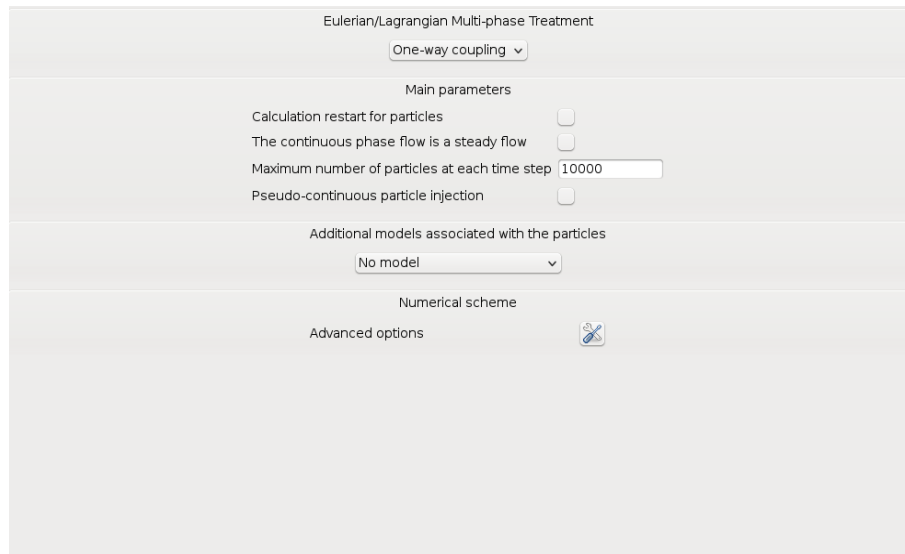


Figure 55: Lagrangian module - View of the Global Settings page

### 8.6.5 Prescribing particle boundary conditions (GUI and/or uslag2)

In the framework of the multiphase Lagrangian modelling, the management of the boundary conditions concerns the particle behaviour when there is an interaction between its trajectory and a boundary face. These boundary conditions may be imposed independently of those concerning the Eulerian fluid phase (but they are of course generally consistent). The boundary condition zones are actually redefined by the Lagrangian module (cf. §3.9.4), and a type of particle behaviour is associated with each one. The boundary conditions related to particles can be defined in the Graphical User Interface (GUI) or in the subroutine `uslag2`. More advanced user-defined boundary conditions can be prescribed in the user-subroutine `uslain`.

#### USE OF THE GUI

In the GUI, selecting the Lagrangian module in the activates the item **Particle boundary conditions** under the heading **Boundary conditions** in the tree menu. Different options are available depending on the type of standard boundary conditions selected (wall, inlet/outlet, etc...), see Figure 57.

#### USE OF THE SUBROUTINE `USLAG2`

The main numerical variables are described below.

**ifrlag(nfabor) [ia]:** In the Lagrangian module, the user defines **nfrlag** boundary zones from the color of the boundary faces, or more generally from their properties (colors, groups...), from the boundary conditions defined in `cs_user_boundary_conditions`, or even from their coordinates. To do so, the array **ifrlag(nfabor)** giving for each face **ifac** the number **ifrlag(ifac)** corresponding to the zone to which it belongs, is completed. The zone numbers (*i.e.* the values of **ifrlag(ifac)**) are chosen freely by the user, but must be strictly positive integers less than or equal to **nflagn** (parameter stored in `lagpar`, whose default value is 100). A zone type is associated with every zone; it will be used to impose global boundary conditions. *WARNING: it is essential that every boundary face belongs to a zone..*

**iusncl(nflagn) [ia]:** For all the **nfrlag** boundary zones previously identified, the number of classes

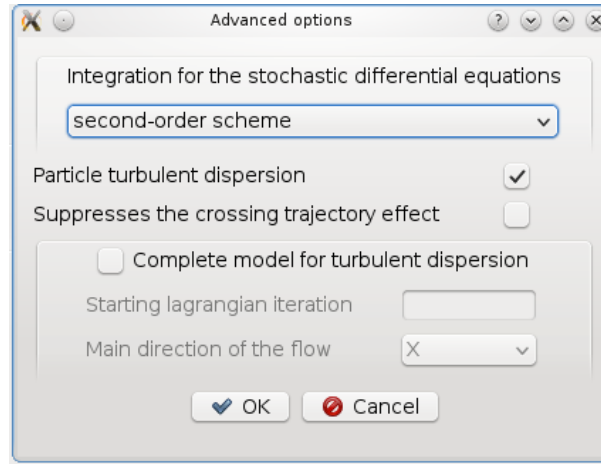


Figure 56: Lagrangian module - Global Settings, advanced numerical options

Lagrangian boundary conditions			
Label	Nature	Particle-boundary interaction	Number of classes
wall	wall	Particles rebound	0
injection	inlet	Particles inlet	0
out	outlet	Particles outlet	0
wall_1	wall	Deposition and elimination ▾	0
		Deposition Particles inlet Particles rebound Deposition and elimination	

Figure 57: Lagrangian module - boundary conditions

$\text{nbclas}^{28}$  of entering particles is given:  $\text{iusncl}(\text{izone}) = \text{nbclas}$ . By default, the number of particle classes is zero. The maximum number of classes is  $\text{nclagm}$  (parameter stored in `lagpar`, whose default value is 20)..

`iusclb(nflagm) [ia]`: For all the `nflag` boundary zones previously identified, a particle boundary condition type is given. The categories of particle boundary condition types are marked out by the keywords `ientrl`, `isortl`, `irebol`, `idepo1`, `idepo2`, `iencrl`).

- if  $\text{iusclb}(\text{izone}) = \text{ientrl}$ ,  $\text{izone}$  is a particle injection zone. For each particle class associated with this zone, information must be provided (see below). If a particle trajectory may cross an injection zone, then this particle leaves the calculation domain.
- if  $\text{iusclb}(\text{izone}) = \text{isortl}$ , the particles interacting with the zone  $\text{izone}$  leave definitely the calculation domain.
- if  $\text{iusclb}(\text{izone}) = \text{irebol}$ , the particles undergo an elastic rebound on the boundary zone  $\text{izone}$ .

<sup>28</sup>a class is a set of particles sharing the same physical properties and the same characteristics concerning the injection in the calculation domain



- if `iusclb(izone) = idepo1`, the particles settle definitely on the boundary zone `izone`. These particles leave the calculation domain and are definitely erased from the calculation
- if `iusclb(izone) = idepo2`, the particles settle definitely on the boundary zone `izone` and they are kept in the calculation domain: the particles do not disappear after touching the boundary zone. However, using `idepo2` type zones necessitates more memory than using `idepo1` type zones.
- if `iusclb(izone) = iencr1`, the particles which are coal particles (if `iphyla = 2`) can become fouled up on the zone `izone`. The slagging is a `idepo1` type deposit of the coal particle if a certain criterion is respected. Otherwise, the coal particle rebounds (`irebo1` type behaviour). This boundary condition type is available if `iencra = 1`. A limit temperature `tprenc`, a critical viscosity `visref` and the coal composition in mineral matters must be given in the subroutine `uslag1`.

`iuslag(nclagm, nflagm, ndlaim) [ia]`: Some pieces of information must be given for each particle class associated with an injection zone. The first part consists in integers contained in the array `iuslag`. There are at the most `ndlaim` integers. These pieces of information must be provided for each class `iclas` and each particle injection zone `izone`. They are marked out by means of “pointers”:

- `iuslag(iclas,izone,ijnbp)`: number of particles to inject in the calculation domain per class and per zone.
- `iuslag(iclas,izone,ijfre)`: injection period (expressed in number of time steps). If the period is null, then there is injection only at the first absolute Lagrangian time step (including the restart calculations).
- `iuslag(iclas,izone,ijuvw)`: type of velocity condition:
  - if `iuslag(iclas,izone,ijuvw) = 1`, the particle velocity vector is imposed, and its components must be given in the array `ruslag` (see below).
  - if `iuslag(iclas,izone,ijuvw) = 0`, the particle velocity is imposed perpendicular to the injection boundary face and with the norm `ruslag(iclas,izone,iuno)`.
  - if `iuslag(iclas,izone,ijuvw) = -1`, the particle injection velocity is equal to the fluid velocity at the center of the cell neighbouring the injection boundary face.
- `iuslag(iclas,izone,inuch1)`: when the particles are coal particles (`iphyla = 2`), this part of the array contains the coal index-number, between 1 and `ncharb` (defined by the user in the thermochemical file `dp_FCP`, with `ncharb ≤ ncharm = 3`).

`ruslag(nclagm, nflagm, ndlagm) [ra]`: Some pieces of information must be given for each particle class associated with an injection zone. The second and last part consists in real numbers contained in the array `ruslag`. There are at the most `ndlagm` such real numbers. These pieces of information must be provided for each class `iclas` and each particle injection zone `izone`. They are marked out by means of “pointers”:

- `ruslag(iclas,izone,iuno)`: norm of the injection velocity, useful if `iuslag(iclas,izone,ijuvw) = 0`.
- `ruslag(iclas,izone,iupt)`, `ruslag(iclas,izone,ivpt)`,  
`ruslag(iclas,izone,iwpt)`: components of the particle injection vector, useful if `iuslag(iclas,izone,ijuvw) = 1`.
- `ruslag(iclas,izone,idebt)`: allows to impose a particle mass flow. According to the number of injected particles, the particle statistical weight `tepa(npt,jrpoi)` is recalculated in order to respect the required mass flow (the number of injected particles does not change). When the mass flow is null, it is not taken into account.

EDF R&D	<b>Code_Saturne version 3.2.0 practical user's guide</b>	Code_Saturne documentation Page 123/202
---------	--	---

- `ruslag(iclas,izone,ipoit)`: particle statistical weight per class and per zone.
- `ruslag(iclas,izone,idpt)`: particle diameter. When the particles are coal particles (`iphyla = 2`), this diameter is provided by the thermochemical file `dp_FCP` via the array `diam20(iclg)`, where `iclg` is the “pointer” on the total class number (*i.e.* for all the coal types). When the standard deviation of the particle diameter is different from zero, this diameter becomes a mean diameter.
- `ruslag(iclas,izone,ivdpt)`: standard deviation of the injection diameter. To impose this standard deviation allows to respect granulometric distribution: the diameter of each particle is calculated from the mean diameter, the standard deviation and a Gaussian random number. In this case, it is strongly recommended to intervene in the subroutine `uslain` to restrict the diameter variation range, in order to avoid aberrant values. If this standard deviation is null, then the particle diameter is constant per class and per zone.
- `ruslag(iclas,izone,iropt)`: particle density. When the particles are coal particles (`iphyla = 2`), this density is set in the thermochemical file `dp_FCP` via the array `rho0ch(icha)`, where `icha` is the coal number.
- `ruslag(iclas,izone,itpt)`: particle injection temperature in °C. Useful if `iphyla = 1` and if `itpvar = 1`.
- `ruslag(iclas,izone,icpt)`: particle injection specific heat. Useful if `iphyla = 1` and if `itpvar = 1`. When the particles are coal particles (`iphyla = 2`), the specific heat is set in the thermochemical file `dp_FCP` via the array `cp2ch(icha)`.
- `ruslag(iclas,izone,iepsi)`: particle emissivity. Useful if `iphyla = 1` and if `itpvar = 1`, and if the radiation module is activated for the continuous phase (note: when `iphyla = 2`, the coal particle emissivity is given the value 1).
- `ruslag(iclas,izone,ihpt)`: particle injection temperature in °C when these particles are coal particles. The array `ruslag(iclas,izone,itpt)` is then no longer active. Useful if `iphyla = 2`.
- `ruslag(iclas,izone,imcht)`: mass of reactive coal. Useful if `iphyla = 2`.
- `ruslag(iclas,izone,imckt)`: mass of coke. This mass is null if the coal did not begin to burn before its injection. Useful if `iphyla = 2`.

`iusvis(nflagm) [ia]`: In order to display the variables at the boundaries defined in the subroutine `uslag1`, this array allows to select the boundary zones on which a display is wanted. To do so, a number is associated with each zone `izone`. If this number is strictly positive, the corresponding zone is selected; if it is null, the corresponding zone is eliminated. If several zones are associated with the same number, they will be displayed together in the same selection with *EnSight*. Each selection will be split in *EnSight* parts according to the geometric types of the present boundary faces (*i.e.* 'tria3', 'quad4' and 'nsided')..

## 8.6.6 Advanced particle-tracking set-up

In this section, some information is provided for a more advanced numerical set-up of a particle-tracking simulation.

### USER-DEFINED INLET PARTICLE PROFILE

The user subroutine `uslain` can be used to complete `uslag2` when the particles must be injected in the domain according to fine constraints (profile, position, ...): the arrays `ettp`, `tepa` and `itepa` can

EDF R&D	<b>Code_Saturne version 3.2.0 practical user's guide</b>	Code_Saturne documentation Page 124/202
---------	--	---

be modified here for the new particles (these arrays were previously completed automatically by the code from the data provided by the user in **uslag2**).

In the case of a more advanced use, it is possible to modify here all the arrays **ettp**, **tepa** and **itepa**. The particles already present in the calculation domain are marked out by an index varying between 1 and **nbpert**. The particles entering the calculation domain at the current iteration are marked out by an index varying between **nbpert**+1 and **nbpnew**.

#### USER-DEFINED VOLUME STATISTICS : **USLAST** AND **USLAEN**

An intervention in both subroutines **uslast** and **uslaen** is required if supplementary user volumetric statistics are wanted.

The subroutine **uslast** is called at the end of every Lagrangian iteration, it allows therefore the modification of variables related to the particles, or the extraction and preparation of data to display in the listing or the post-processing.

The volumetric statistics are calculated by means of the array **statis**. Two situations may happen:

- the calculation of the statistics is not stationary: **statis** is reset at every Lagrangian iteration;
- the calculation of the statistics is stationary: the array **statis** is used to store cumulated values of variables, which will be averaged at the end of the calculation in the subroutine **uslaen**.

According to the user parameter settings, it may happen that during the same calculation, the statistics will be non-stationary in a first part and stationary in second part.

In the subroutine **uslaen**, the variable whose volumetric statistic is wanted is stored in the array **statis**. In the framework of stationary statistics, the average itself is calculated in the subroutine **uslaen**. This average is obtained through the division of the cumulated value by:

- either the duration of the stationary statistics calculation stored in the variable **tstat**,
- or the number of particles in statistical weight.

This method of averaging is applied to every piece in the listing and to the post-processing outputs.

In this subroutine is calculated the average corresponding to the cumulated value obtained in the subroutine **uslast**. This subroutine is also used for the standard volumetric statistics. Several examples are therefore described.

#### USER-DEFINED STOCHASTIC DIFFERENTIAL EQUATIONS

An intervention in the subroutine **uslaed** is required if supplementary user variables are added to the particle state vector (arrays **ettp** and **ettpa**). This subroutine is called at each Lagrangian sub-step.

The integration of the stochastic differential equations associated with supplementary particulate variables is done in this subroutine.

When the integration scheme of the stochastic differential equations is a first-order (**nordre** = 1), this subroutine is called once every Lagrangian iteration, if it is a second-order (**nordre** = 2), it is called twice.

The solved stochastic differential equations must be written in the form:

$$\frac{d\Phi_p}{dt} = -\frac{\Phi_p - \Pi}{\tau_\phi}$$

where  $\Phi_p$  is the  $I$ th supplementary user variable (**nvls** in total) available in **ettp(nbpmax, jvls(i))** and in **ettpa(nbpmax, jvls(i))**,  $\tau_\phi$  is a quantity homogeneous to a characteristic time, and  $\Pi$  is a coefficient which may be expressed as a function of the other particulate variables contained in **ettp** and **ettpa**.

In order to do the integration of this equation, the following parameters must be provided:

EDF R&D	<b><i>Code_Saturne</i> version 3.2.0 practical user's guide</b>	<i>Code_Saturne</i> documentation Page 125/202
---------	---	--

- $\tau_\phi$ , equation characteristic time, in the array **aux11** for every particle,
- $\Pi$ , equation coefficient, in the array **aux12**. If the integration scheme is a first-order, then  $\Pi$  is expressed as a function of the particulate variables at the previous iteration, stored in the array **ettpa**. If the chosen scheme is a second-order, then  $\Pi$  is expressed at the first call of the subroutine (prediction step **nor** = 1) as a function of the variables at the previous iteration (stored in **ettpa**), then at the second call (correction step **nor** = 2) as a function of the predicted variables stored in the array **ettp**.

If necessary, the thermal characteristic time  $\tau_c$ , whose calculation can be modified by the user in the subroutine **uslatc**, is stored for each particle in the part **tempct(nbpmax,1)** of the array **tempct**.

#### USER-DEFINED PARTICLE RELAXATION TIME

The particle relaxation time may be modified in the subroutine **uslatp** according to the chosen formulation of the drag coefficient. The particle relaxation time, modified or not by the user, is available in the array **taup**.

#### USER-DEFINED PARTICLE THERMAL CHARACTERISTIC TIME

The particle thermal characteristic time may be modified in the subroutine **uslatc** according to the chosen correlation for the calculation of the Nusselt number. This subroutine is called at each Lagrangian sub-step. The thermal characteristic time, modified or not by the user, is available in the zone **tempct(nbpmax,1)** of the array **tempct**.

#### PARTICLE CLONING/MERGING TECHNIC

An intervention in the subroutine **uslaru** is required when the particle cloning/merging option is activated *via* the keyword **iroule**. The important function 'croule' must then be completed.

The aim of this technique is to reduce the number of particles to treat in the whole flow and to refine the description of the particle cloud only where the user wants to get more accurate volumetric statistics than in the rest of the calculation domain.

The values given to the importance function are strictly positive real numbers allowing to classify the zones according to their importance. The higher the value given to the importance function, the more important the zone.

For instance, when a particle moves from a zone of importance 1 to a zone of importance 2, it undergoes a cloning: the particle is replaced by two identical particles, whose statistical weight is the half of the initial particle. When a particle moves from a zone of importance 2 to a zone of importance 1, it undergoes a fusion: the particle survives to its passing through with a probability of 1/2. A random dawning is used to determine if the particle will survive or disappear.

In the same way, when a particle moves from a zone of importance 3 to a zone of importance 7, it undergoes a cloning. The particle is cloned in  $\text{Int}(7/3)=2$  or  $\text{Int}(7/3)+1=3$  particles with a probability of respectively  $1-(7/3-\text{Int}(7/3))=2/3$  and  $7/3-\text{Int}(7/3)=1/3$ . If the particle moves from a zone of importance 7 to a zone of importance 3, it undergoes a fusion: it survives with a probability of 3/7.

*WARNING: The importance function must be a strictly positive real number in every cell*

## 8.7 Compressible module

When the compressible module<sup>29</sup> is activated, it is recommended to:

<sup>29</sup>For more details concerning the compressible version, the user may refer to the theory guide [11] and the document "Implantation d'un algorithme compressible dans *Code\_Saturne*", Rapport EDF 2003, HI-83/03/016/A, P. Mathon, F. Archambeau et J.-M. Hérard.

EDF R&D	<b><i>Code_Saturne</i> version 3.2.0 practical user's guide</b>	<i>Code_Saturne</i> documentation Page 126/ <a href="#">202</a>
---------	---	---

- use the option “time step variable in time and uniform in space” (`idtvar=1`) with a maximum Courant number of 0.4 (`coumax=0.4`): these choices must be written in `cs_user_parameters.f90` or specify with the GUI.
- keep the convective numerical schemes proposed by default (*i.e.*: upwind scheme).

With the compressible algorithm, the density and the specific total energy are two new solved variables (`isca(irho)` and `isca(ienerg)`). The temperature variable deduced from the specific total energy variable is `isca(itempk)` for the compressible module.

Initialisation of the options of the variables, boundary conditions, initialisation of the variables and management of variable physical properties can be done with the GUI. We describe below the subroutines the user has to fill in without the GUI.

### 8.7.1 Initialisation of the options of the variables

*Subroutines called at each time step.*

Without the GUI, the subroutines `uscfx1` and `uscfx2` in `cs_user_parameters.f90` must be completed.

`uscfx1` allows to specify:

- `ieos`: equation of state (only perfect gas with a constant adiabatic coefficient, `ieos=1` is available, but the user can complete the subroutine `cfther`, which is not a user subroutine, to add new equations of state).
- `ivisls(itempk)`: molecular thermal conductivity, constant (0) or variable (1).
- `iviscv`: volumetric molecular viscosity, constant (0) or variable (1).

`uscfx2` allows to specify:

- `ivivar`: molecular viscosity, constant (0) or variable (1).
- `visls0(itempk)`: reference molecular thermal conductivity.
- `viscv0`: reference volumetric molecular viscosity.
- `xmasmr`: molar mass of the perfect gas (`ieos=1`).
- `icfgrp`: specify if the hydrostatic equilibrium must be accounted for in the boundary conditions.

### 8.7.2 Management of the boundary conditions

*Subroutine called at each time step.*

Without the GUI, the `cs_user_boundary_conditions` subroutine may be used with the compressible module to define specific boundary conditions (see the `cs_user_boundary_conditions-compressible` file in the directory `EXAMPLES` for examples of boundary conditions with the compressible module).

With the compressible module, the different available boundary condition types are the following:

- Inlet/outlet for which velocity and two thermodynamics variables are known.
- Subsonic inlet with imposed density and velocity.
- Subsonic outlet with imposed static pressure.
- Supersonic outlet.

- Wall (adiabatic or not).
- Symmetry.

It is advised to only use these predefined boundary conditions type for the compressible module.

### 8.7.3 Initialisation of the variables

*Subroutine called only at the initialisation of the calculation*

Without the GUI, the subroutine `cs_user_initialization` may be used for initialization of velocity, turbulence and passive scalars (see the `cs_user_initialization-compressible` file in the directory `EXAMPLES` for examples of initialisations with the compressible module). Concerning pressure, density, temperature and specific total energy, only 2 variables out of the 4 are independent. The user may then initialise the desired variable pair (apart from temperature-energy) and the two other variables will be calculated automatically by giving the right value to the variable `iccfth` used for the call to the subroutine `cfther`.

### 8.7.4 Management of variable physical properties

*Subroutine called at each time step.*

Without the GUI, all the laws of the fluid physical properties (molecular viscosity, molecular volumetric viscosity, molecular thermal conductivity, molecular diffusivity of the user-defined scalars) can be specified in the subroutine `uscfpv` of the `cs_user_physical_properties` file, which is then called at each time step. This subroutine replaces and is similar to `usphyv`.

The user should check that the defined laws are valid for the whole variation range of the variables. Moreover, as only the perfect gas with a constant adiabatic coefficient equation of state is available, it is not advised to give a law for the isobaric specific heat without modifying the equation of state in the subroutine `cfther` which is not a user subroutine.

## 8.8 Management of the electric arcs module

### 8.8.1 Activating the electric arcs module

The activation of the electric arcs module is performed either:

- in the Graphical User Interface (GUI): `Calculation features` → `Electrical models`
- or in the user subroutine `usppmo`, by setting the `ielarc` or `ieljou` parameter to a non-null value.

### 8.8.2 Initialisation of the variables

*subroutine called only at the initialisation of the calculation*

The subroutine `cs_user_initialization` allows the user to initialise some of the specific physics variables prompted via `usppmo`. It is called only during the initialisation of the calculation. The user has access, as usual, to many geometric variables so that the zones can be treated separately if needed.

The values of potential and its constituents are initialised if required.

It should be noted that the enthalpy is relevant.

- For the electric arcs module, the enthalpy value is taken from the temperature of reference `t0` (given in `cs_user_parameters.f90`) from the temperature-enthalpy tables supplied in the data file `dp_ELE`. The user must not intervene here.

EDF R&D	<b>Code_Saturne version 3.2.0 practical user's guide</b>	Code_Saturne documentation Page 128/202
---------	--	---

- For the Joule effect module, the value of enthalpy must be specified by the user . An example is given of how to obtain the enthalpy from the temperature of reference `t0`(given in `cs_user_parameters.f90`), the temperature-enthalpy law must be supplied. A code is suggested in the sub routine `usthht`(which is there for the determination of physical properties).

### 8.8.3 Variable physical properties

All the laws of the variation of physical data of the fluid are written (when necessary) in the subroutine `cs_user_physical_properties`. It is called at each time step.

*WARNING: For the electric module, it is here that all the physical variables are defined (including the relative cells and the eventual user scalars): `cs_user_physical_properties` is not used.*

The user should ensure that the defined variation laws are valid for the whole range of variables. Particular care should be taken with non-linear laws (for example, a 3<sup>rd</sup> degree polynomial law giving negative values of density)

*WARNING: in the electric module, all the physical properties are considered as variables and are therefore stored in the `propce` array. `cp0`, `viscls0`, `viscl0` are not used*

For the Joule effect, the user is required to supply the physical properties in the subroutine. Examples are given which are to be adapted by the user. If the temperature is to be determined to calculate the physical properties, the solved variable, enthalpy must be deduced. The preferred temperature-enthalpy law can be selected in the subroutine `usthht` (an example of the interpolation is given from the law table. This subroutine can be re-used for the initialisation of the variables(`cs_user_initialization`)) For the electric arcs module, the physical properties are interpolated from the data file `dp_ELE` supplied by the user. Modifications are generally not necessary.

### 8.8.4 Boundary Conditions

For the electric module, in `cs_user_boundary_conditions`, each boundary face should be associated with a number `izone`<sup>30</sup>(the color `icoul` for example) in order to group together all the boundary faces of the same type. In the report `cs_user_boundary_conditions`, the main change from the users point of view concerns the specification of the boundary conditions of the potential, which isn't implied by default. The Dirichlet and Neumann conditions must be imposed explicitly using `icodcl` and `rcodcl` (as would be done for the classical scalar).

Furthermore, if one wishes to slow down the power dissipation (Joule effect module) or the current (electric arcs module) from the imposed values (`puismp` and `couimp` respectively), they can be changed by the potential scalar as shown below:

- For the electric arcs, the imposed potential difference can be a fixed variable: for example, the cathode can be fixed at 0 and the potential at the anode contains the variable `dpot`. This variable is initialised in `useli1` (in `cs_user_parameters.f90`) by an estimated potential difference. If `ielcor=1` (see `useli1`), `dpot` is updated automatically during the calculation to obtain the required current.
- For the Joule module effect, `dpot` is again used with the same signification as in the electric arcs module. If `dpot` is not wanted in the setting of the boundary conditions, the variable `coejou` can be used. `coejou` is the coefficient by which the potential difference is multiplied to obtain the desired power dissipation. By default this begins at 1 and is updated automatically. If `ielcor=1` (see `useli1`), multiply the imposed potentials in `cs_user_boundary_conditions` by `coejou` at each time step to achieve the desired power dissipation.

<sup>30</sup>`izone` must be less than the maximum value allowed by the code, `nozzppm`. This is fixed at 2000 in `ppvar` and cannot be modified.



EDF R&D	<b>Code_Saturne version 3.2.0 practical user's guide</b>	Code_Saturne documentation Page 129/202
---------	--	---

*WARNING: In alternative current, attention should be paid to the values of potential imposed at the limits: the variable named "real potential" represents an effective value if the current is in single phase, and a "real part" if not.*

- For the Joule studies, a complex potential is sometimes needed (`ippmod(ieljou)=2`): this is the case in particular where the current has 3 phases. To have access to the phase of the potential, and not just to its amplitude, the 2 variables must be deleted: in *Code\_Saturne*, there are 2 arrays specified for this role, the real part and the imaginary part of the potential. For use in the code, these variables are named "real potential" and "imaginary potential". For an alternative sinusoidal potential  $Pp$ , the maximum value is noted as  $Pp_{\max}$ , the phase is noted as  $\phi$ , the real potential and the imaginary potential are respectively  $Pp_{\max} \cos\phi$  and  $Pp_{\max} \sin\phi$ .
- For the Joule studies in which one does not have access to the phases, the real potential (imaginary part =0) will suffice (`ippmod(ieljou)=1`): this is obviously the case with continuous current, but also with single phase alternative current. In *Code\_Saturne* there is only 1 variable for the potential, called "real potential". Pay attention to the fact that in alternate current, the "real potential" represents a effective value of potential,  $\frac{1}{\sqrt{2}} Pp_{\max}$  (in continuous current there is no such ambiguity).

#### ADDITIONS FOR TRANSFORMERS

The following additional boundary conditions must be defined for transformers:

- the intensity at each electrode
- the voltage on each termin of transformers. To achieve it, the intensity, the rvoltage at each termin, the Rvoltage, and the total intensity of the transformer are calculated.

Finally, a test is performed to check if the offset is zero or if a boundary face is in contact with the ground.

### 8.8.5 Initialisation of the variable options

The subroutine `usel11` (in `cs_user_parameters.f90`) is called at each time step. It allows:

- to give the coefficient of relaxation of the density `srrom`:  

$$\rho^{n+1} = \text{srrom} * \rho^n + (1 - \text{srrom})\rho^n$$
(for the electric arcs, the sub-relaxation is taken into account during the 2nd time step; for the Joule effect the sub relaxation is not accounted for unless the user specifies in `uselph`)
- indicates if the data will be fixed in the power dissipation or in the current, done in `ielcor`.
- target current fixed as `couimp` (electric arcs module) or the power dissipation `puism` (Joule module effect).
- Fix the initial value of potential difference `dpot`, the for the calculations with a single fixed parameter as `couimp` or `puism`.
- Define type of scaling model for electric arcs `modrec`. If scaling by a resetting plane is choosen then `idreca` defines the current density component and `crit_reca` the plane used for resetting of electromagnetic variables.



### 8.8.6 Post-processing output

The algebraic variables related to the electric module are provided by default provided that they are not explicitly contained in the `propce` array:

- gradient of real potential in  $Vm^{-1}$  ( $\nabla Pot_R = -\underline{E}$ )
- density of real current in  $Am^{-2}$  ( $\underline{j} = \sigma \underline{E}$ )

specifically for the Joule module effect with `ippmod(ieljou)=2` :

- gradient of imaginary potential in  $Vm^{-1}$
- density of real current in  $Am^{-2}$

specifically for the electric arcs module with `ippmod(ielarc)=2` :

- magnetic field in  $T$  ( $\underline{B} = \text{rot } \underline{A}$ )

The post-processing output will be created automatically (on all output volume meshes for which the automatic output of main variables is active).

## 8.9 Code\_Saturne-Code\_Saturne coupling

*Subroutine called once during the calculation initialisation.*

This user function `cs_user_saturne_coupling` (in `cs_user_coupling.c`) is used to couple *Code\_Saturne* with itself. It is used for turbo-machine applications for instance, the first *Code\_Saturne* managing the fluid around the rotor and the other the fluid around the stator. In the case of a coupling between two *Code\_Saturne* instances, first argument `saturne_name` of the function '`cs_sat_coupling_define`' is ignored. In case of multiple couplings, a coupling will be matched with available *Code\_Saturne* instances based on that argument, which should match the directory name for the given coupled domain.. The arguments of '`cs_sat_coupling_define`' are:

- `saturne_name`: the matching *Code\_Saturne* application name,
- `volume_sup_criteria`: the cell selection criteria for support,
- `boundary_sup_criteria`: the boundary face selection criteria for support (not functional),
- `volume_cpl_criteria`: the cell selection criteria for coupled cells,
- `boundary_cpl_criteria`: the boundary face selection criteria for coupled faces,
- `verbosity`: the verbosity level.

## 8.10 Fluid-Structure external coupling

*Subroutine called only once*

The subroutine `usaste` belongs to the module dedicated to external Fluid-Structure coupling with *Code\_Aster*. Here one defines the boundary faces coupled with *Code\_Aster* and the fluid forces components which are given to structural calculation. When using external coupling with *Code\_Aster*, structure numbers necessarily need to be negative; the references of coupled faces being i.e. -1, -2, etc. The subroutine performs the following operations:

- 'getfbr' is called to get a list of elements matching a geometrical criterion or reference number then a structure number (negative value) is associated to these elements.
- the value passed to `asddlf`, for user-chosen component, for every negative structure number, defines the movement imposed to the external structure.

## 8.11 ALE module

### 8.11.1 Initialisation of the options

This initialisation can be performed in the Graphical User Interface (GUI) or in the subroutines `usalin` and `usstr1`. First of all, in the GUI when the "Mobile mesh" is selected in the "Thermophysical models" heading, additional options are displayed. The user must choose a type of mesh viscosity and how to describe its spatial distribution, see Figure 58.

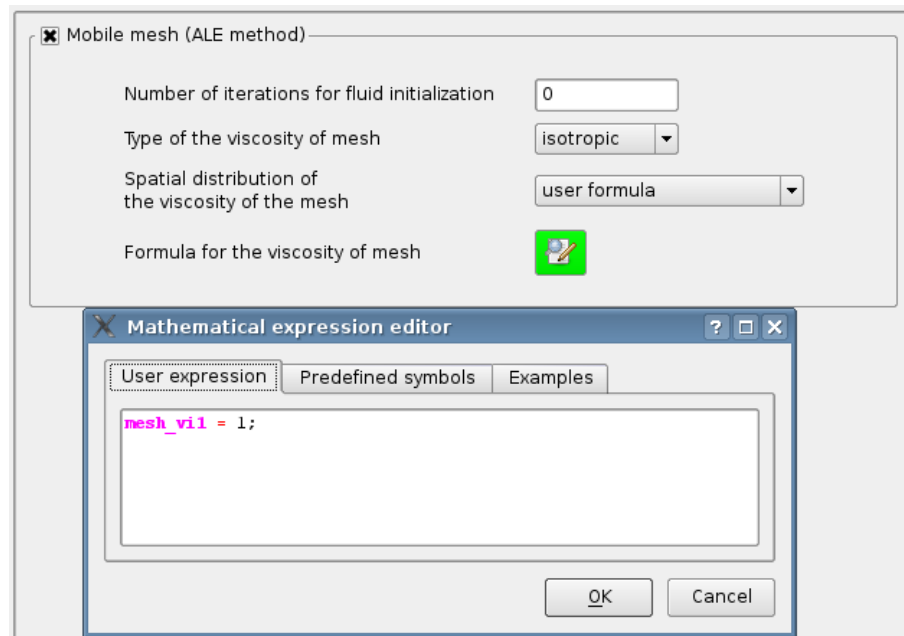


Figure 58: Thermophysical models - mobile mesh (ALE method)

The following paragraphs are relevant if the GUI is not used.

#### SUBROUTINE USALIN

*Subroutine called at the beginning.* This subroutine completes `cs_user_parameters.f90`.

`usalin` allows to set options for the ALE module, and in particular to active the ALE module (`iale=1`).

#### SUBROUTINE USSTR1

`usstr1` This subroutine reads in `cs_user_fluid_structure_interaction.f90`. It allows to specify for the structure module the following pieces of information:

- the index of the structure, (`idfstr(ifac)` where `ifac` is the index of the face). Then the total number of structures `nbstru` is automatically computed by the code. Be carefull, the value must belong to 1, ..., `nbstru`.

EDF R&D	<b><i>Code_Saturne</i> version 3.2.0 practical user's guide</b>	<i>Code_Saturne</i> documentation Page 132/202
---------	---	--

- the initial value of displacement, velocity and acceleration (`xstr0`, `xstreq` and `vstr0`).

Below is a list of the different variables that might be modified:

- `idfstr(ifac)`  
the index of the structure, (`idfstr(ifac)` where `ifac` is the index of the face), 0 if the face is not coupled to any structure.
- `xstr0(i,k)`  
initial position of a structure, where `i` is the dimension of space and `k` the index of the structure
- `xstreq(i,k)`  
position of balance of a structure, where `i` is the dimension of space and `k` the index of the structure
- `vstr0(i,k)`  
initial velocity of a structure, where `i` is the dimension of space and `k` the index of the structure

### 8.11.2 Boundary conditions of mesh velocity

The boundary conditions can be managed with the Graphical User Interface (GUI) or with the subroutine `usalcl` (called at each time step). In the GUI, when the item “Mobile mesh” is activated the item “Fluid structure interaction” appears under the heading “Boundary conditions”. Two types of Fluid-structure coupling are offered. The first one is internal, using a simplified structure model and the second is external with *Code\_Aster*, see Figure 59 and Figure 60.

#### SUBROUTINE USALCL

When the GUI is not used, the use of `usalcl` is mandatory to run a calculation using the ale module just as it is in `cs_user_parameters.f90`. The way of using it is the same as the way of using `cs_user_boundary_conditions` in the framework of standard calculations, that is to say a loop on the boundary faces marked out by their colour (or more generally by a property of their family), where the type of boundary condition of mesh velocity for each variable are defined.

The main numerical variables are described below.

`ialtyb(nfabor)` [ia]: In the ale module, the user defines the mesh velocity from the colour of the boundary faces, or more generally from their properties (colours, groups, ...), from the boundary conditions defined in `cs_user_boundary_conditions`, or even from their coordinates. To do so, the array `ialtyb(nfabor)` gives for each face `ifac` the mesh velocity boundary condition types marked out by the key words `ivimpo`, `igliss`, `ibfixe` or `ifresf`.

- If `ialtyb(ifac) = ivimpo`: imposed velocity.
  - In the case where all the nodes of a face have a imposed displacement, it is not necessary to fill the tables with boundary conditions mesh velocity for this face, they will be erased. In the other case, the value of the Dirichlet must be given in `rcodcl(ifac,ivar,1)` for every value of `ivar` (`iuma`, `ivma` and `iwma`). The other boxes of `rcodcl` and `icodcl` are completed automatically.
  - The tangential mesh velocity is taken like a tape speed under the boundary conditions of wall for the fluid, except if wall fluid velocity was specified by the user in the interface or `cs_user_boundary_conditions` (in which case it is this speed which is considered).
- if `ialtyb(ifac) = ibfixe`: fixed wall
  - the velocity is null.
- if `ialtyb(ifac) = igliss`: sliding wall

→ symmetry boundary condition on the mesh velocity vector, which means an homogeneous Neumann on the tangential mesh velocity and a zero Dirichlet on the normal mesh velocity.

- if `ialtyb(ifac) = ifrsf`: free-surface

→ an imposed mesh velocity so that the fluid mass flux is equal to the mesh displacement to mimic free-surface automatically. Note that the boundary condition on the fluid velocity must be set separately (homogeneous Neumann for instance).

### 8.11.3 Modification of the viscosity

The user subroutine `usvima` is used along the ALE (Arbitrary Lagrangian Eulerian Method) module, and fills mesh viscosity arrays. It is called at each time step. The user can modify mesh viscosity values to prevent cells and nodes from huge displacements in awkward areas, such as boundary layer for example. If `iortvm = 0`, the mesh viscosity modelling is considered as isotropic and therefore only the `viscmx` array must be filled. If `iortvm = 1`, mesh viscosity modelling is orthotropic therefore all arrays `viscmx`, `viscmz`, and `viscmz` need to be filled. Note that `viscmx`, `viscmz` and `viscmz` arrays are initialized at the first time step with the value 1.

### 8.11.4 Fluid - Structure internal coupling

In the subroutine `cs_user_fluid_structure_interaction` the user provides the parameters of two other subroutines. `usstr1` is called at the beginning of the calculation. It is used to define and initialise the internal structures where Fluid-Structure coupling occurs. For each boundary face `ifac`, `idfstr(ifac)` is the index of the structure the face belongs to (if `idfstr(ifac) = 0`, the face `ifac` doesn't belong to any structure). When using internal coupling, structure index necessarily must be strictly positive and smaller than the number of structures. The number of "internal" structures is automatically defined with the maximum value of the `idfstr` table, meaning that internal structure numbers must be defined sequentially with positive values, beginning with integer value '1'.

For each internal structure one can define:

- an initial velocity `vstr0`
- an initial displacement `xstr0` (i.e. `xstr0` is the value of the displacement `xstr` compared to the initial mesh at time  $t = 0$ )
- a displacement compared to equilibrium `xstreq` (i.e. `xstreq` is the initial displacement of the internal structure compared to its position at equilibrium; at each time step  $t$  and for a displacement `xstr(t)`, the associated internal structure will undergo a force  $-k * (t + XSTREQ)$  due to the spring).

`xstr0` and `vstr0` are initialised with the value 0. When starting a calculation using ALE, or re-starting a calculation with ALE, based on a first calculation without ALE, an initial iteration 0 is automatically performed in order to take initial arrays `xstr0`, `vstr0` and `xstreq` into account. In any other case, add the following expression `'italin=1'` in subroutine `usalin`, so that the code can deal with the arrays `xstr0`, `vstr0` and `xstreq`.

When `ihistr` is set to 1, the code writes in the output the history of the displacement, of the structural velocity, of the structural acceleration and of the fluid force. The value of structural history output step is the same as the one for standard variables `nthist`.

The second subroutine, `usstr2`, is called at each iteration. One defines in this subroutine structural parameters (considered as potentially time dependent): i.e., mass `m xmstru`, friction coefficients `c xcstru`, and stiffness `k xkstru`. `forstr` array gives fluid stresses acting on each internal structure. Moreover it is also possible to take external forces (gravity for example) into account.

- . xstr array indicates the displacement of the structure compared to its position in initial mesh,
- . xstr0 array gives the displacement of the structures in initial mesh compared to structural equilibrium,
- . vstr array stands for structural velocity.

xstr, xstr0 and vstr are DATA tables that can be used to define arrays Mass, Friction and Stiffness. Those are not to be modified.

The 3D structural equation that is solved is the following one:

$$\underline{\underline{m}}.\partial_{tt}\underline{x} + \underline{\underline{c}}.\partial_t\underline{x} + \underline{\underline{k}}.(\underline{x} + \underline{x}_0) = \underline{f}, \quad (6)$$

where  $x$  stands for the structural displacement compared to initial mesh position xstr,  $x_0$  represents the displacement of the structure in initial mesh compared to equilibrium. Note that  $\underline{\underline{m}}, \underline{\underline{c}},$  and  $\underline{\underline{k}}$  are 3x3 matrices. Equation (6) is solved using a Newmark HHT algorithm. Note that the time step used to solve this equation, dtstr, can be different from the one of fluid calculations. The user is free to define dtstr array. At the beginning of the calculation dtstr is initialised to the value of dtcel (fluid time step).

## 8.12 Management of the structure property

The use of `usstr2` is mandatory to run a calculation using the ALE module with a structure module. It is called at each time step.

For each structure, the system that will be solved is:

$$M.x'' + C.x'' + K.(x - x_0) = 0 \quad (7)$$

where

- $M$  is the mass stucture (`xmstru`).
- $C$  is the dumping coefficient of the stucture (`xcstru`).
- $K$  is the spring constant or force constant of the stucture (`xkstru`).
- $x_0$  is the initial position.

Below is a list of the different variables that might be modified:

- `xmstru(i,j,k)`  
the mass stucture of the structure, where `i,j` is the array of mass stucture and `k` the index of the structure.
- `xcstru(i,j,k)`  
dumping coefficient of the stucture, where `i,j` is the array of dumping coefficient and `k` the index of the structure.
- `xkstru(i,j,k)`  
spring constant of the stucture, where `i,j` is the array of spring constant and `k` the index of the structure.
- `forstr(i,k)`  
force vector of the stucture, where `i` is the force vector and `k` the index of the structure.

## 8.13 Management of the Atmospheric module

The selection can be done in the Graphical User Interface (GUI) under the heading “Thermophysical models” item “Calculation features” or in the routine `usppmo` as described in the section “Advanced modelling setup”.

One of the following atmospheric flow model can be selected:

- “constant density” : no thermal variable, to simulate neutral atmosphere,
- “dry atmosphere” : potential temperature is used as thermal variable to simulate dry stratified atmospheric flows,
- “humid atmosphere” : liquid potential temperature, total water content, and droplet number are used to simulate stratified atmospheric flows with water (liquid and vapor) together with phase changes. The model is described in (Bouzereau, 2004).

When one of the atmospheric option is selected, it activates an item under the same heading: “Atmospheric flows” where the path leading to a file containing meteorological data can be specified, see fig. 61. An example is given in the DATA/REFERENCE/meteo file.

The user can use the standard set of *Code\_Saturne* boundary conditions but is warned that even small inconsistencies can create very large buoyancy forces and spurious circulations.

Alternatively, the meteorological profile can be used to initialize the fields and to set up the inlet boundary conditions. Optionally, the inlet can be detected automatically according to the direction of the wind given in the meteorological file. This is often used for the lateral boundaries of the atmospheric domain.

**WARNING :** the definition of the potential temperature (and the liquid potential temperature) requires that the vertical component of the gravity is set at  $GZ=-9.81$  ( $GX=GY=0$ ) otherwise pressure and density are not correctly computed.

### 8.13.1 Initialisation of the variables

If a meteorological file is given, it is used by default for initializing the variables. If not the standard initialization is performed. The initialisation can be modified in the Graphical User Interface (GUI) or in the subroutine `cs_user_initialization_atmospheric` example. In addition if the atmospheric flow model chosen is “dry atmosphere”, the additional variable “PotTemp” appear below the other code variables. For the “humid atmosphere”, the additional variables are : “LiqPotTemp”, “TotWater” and “NbDrop”.

When the GUI is not used, `cs_user_initialization_atmospheric` example (called only during the calculation initialisation) allows to initialise or modify (in case of a restarted calculation) the calculation variables and the values of the time step. The example provided in the user file performs the initialisation of the variables from meteorological profiles using the interpolation routine `intprf`.

### 8.13.2 Management of the boundary conditions

Boundary conditions can be set with the GUI, including the automatic use of the meteorological profile for inlets. If not managed by the GUI, the boundary conditions must be managed by the `cs_user_boundary_conditions_atmospheric` subroutine which gives various examples . For each type of boundary condition, faces should be grouped as physical zones characterised by an arbitrary number `izone` chosen by the user. If a boundary condition is retrieved from a meteorological profile, the variable `iprofm(izone)` of the zone must be set to 1.

Internal coupling with a simplified structure model

External coupling with Code\_Aster


Internal coupling

Maximum number of sub-iterations for implicit coupling with internal structures

1

Relative precision for implicit coupling with internal structures

1e-05

Advanced options


Structures definition

Structure number	Label	Location

Initial position

X  m
Y  m
Z  m


Position of equilibrium


X  m
Y  m
Z  m


Initial velocity

V<sub>x</sub> m/s
V<sub>y</sub> m/s
V<sub>z</sub> m/s

Characteristics of the structure

Mass matrix


Damping matrix


Stiffness matrix


Force applied to the structure




Figure 59: Boundary conditions - internal coupling

Internal coupling with a simplified structure model    External coupling with Code\_Aster

External coupling

Fluid-structure post-treatment synchronisation ☐

Structures definition

Structure number	Label	Location
------------------	-------	----------

Blocage des DDL de force:

DDLX ☐

DDLX ☐

DDLZ ☐

Figure 60: Boundary conditions - external coupling

Atmospheric flows

☒ Read the file of meteorological data 

name of the data file:

Figure 61: Thermophysical models - atmospheric flows



## 9 Key word list

The key words are classified under headings. For each key word of the Kernel of *Code\_Saturne*, the following data are given:

Variable name	Type	Allowed values	[Default]	O/C	Level
	Description	Potential dependences			

- **Variable name:** Name of the variable containing the key word.
- **Type:** a (Array), i (Integer), r (Real number), c (Character string).
- **Allowed values:** list or range of allowed values.
- **Default:** value defined by the code before any user modification (every key word has one). In some cases, a non-allowed value is given (generally  $-999$  or  $-10^{12}$ ), to force the user to specify a value. If he does not do it, the code may:
  - automatically use a recommended value (for instance, automatic choice of the variables for which chronological records will be generated).
  - stop, if the key word is essential.
- **O/C:** Optional/Compulsory
  - O: optional key word, whose default value may be enough.
  - C: key word which must imperatively be specified.
- **Level:** L1, L2 or L3
  - L1 (level 1): the users will have to modify it in the framework of standard applications. The L1 key words are written in bold.
  - L2 (level 2): the users may have to modify it in the framework of advanced applications. The L2 key word are all optional.
  - L3 (level 3): the developers may have to modify it ; it keeps its default value in any other case. The L3 key word are all optional.
- **Description:** key word description, with its potential dependences.

The L1 key words can be modified through the Graphical Use Interface or in the `cs_user_parameters.f90` file. L2 and L3 key words can only be modified through the `cs_user_parameters.f90` file, even if they do not appear in the version proposed as example in the `SRC/REFERENCE/base` directory. It is however recommended not to modify the key words which do not belong to the L1 level.

The alphabetical key word list is displayed in the index, in the end of this report.

### NOTES

- The notation “d” refers to a double precision real. For instance, 1.8d-2 means 0.018.
- The notation “**grand**” (which can be used in the code) corresponds to  $10^{12}$ .

## 9.1 Input-output

### NOTES

- Two different files can have neither the same unit number nor the same name.

## 9.1.1 "Calculation" files

### GENERAL

#### VORTEX METHOD FOR LES

<code>impmvo</code>	i	strictly positive integer	[ <code>impmvo</code> ]	O	L3
unit of the upstream restart file for the vortex method useful if and only if <code>isuivo = 1</code> and <code>ivrtex=1</code>					
<code>impvvo</code>	i	strictly positive integer	[ <code>impvvo</code> ]	O	L3
unit of the downstream restart file for the vortex method useful if and only if <code>ivrtex=1</code>					
<code>impdvo</code>	i	strictly positive integer	[ <code>impdvo</code> ]	O	L3
unit of the <code>ficvor</code> data files for the vortex method. These files are text files. Their number and names are specified by the user in the <code>usvort</code> subroutine. (Although it corresponds to an "upstream" data file, <code>impdvo</code> is initialized to 20 because, in case of multiple vortex entries, it is opened at the same time as the <code>ficmvo</code> upstream restart file, which already uses unit 11) useful if and only if <code>ivrtex=1</code>					

#### THERMOCHEMISTRY

<code>impfpp</code>	i	strictly positive integer	[25]	O	L3
unit of the thermochemical data file useful in case of gas or pulverised coal combustion or electric arcs					
<code>ficfpp</code>	c	string of 32 characters	[ <code>dp_tch</code> ]	O	L3
name of the thermochemical data file. The launch script is designed to copy the user specified thermochemical data file in the temporary execution directory under the name <code>dp_tch</code> , for <i>Code_Saturne</i> to open it properly. Should the value of <code>ficfpp</code> be changed, the launch script would have to be adapted. useful in case of gas or pulverised coal combustion					
<code>impjnf</code>	i	strictly positive integer	[ <code>impfpp</code> ]	O	L3
unit of the JANAF data file useful in case of gas or pulverised coal combustion					
<code>ficjnf</code>	c	string of 5 characters	[JANAF]	O	L3
name of the JANAF data file. The launch script is designed to copy the user specified JANAF data file in the temporary execution directory under the name <code>JANAF</code> , for <i>Code_Saturne</i> to open it properly. Should the value of <code>ficjnf</code> be changed, the launch script would have to be adapted. useful in case of gas or pulverised coal combustion					

#### LAGRANGIAN

<b>impla1</b>	i	strictly positive integer	[50]	O	L3
		unit of a file specific to Lagrangian modelling useful in case of Lagrangian modelling			
<b>impla2</b>	i	strictly positive integer	[51]	O	L3
		unit of a file specific to Lagrangian modelling useful in case of Lagrangian modelling			
<b>impla3</b>	i	strictly positive integer	[52]	O	L3
		unit of a file specific to Lagrangian modelling useful in case of Lagrangian modelling			
<b>impla4</b>	i	strictly positive integer	[53]	O	L3
		unit of a file specific to Lagrangian modelling useful in case of Lagrangian mode ling			
<b>impla5</b>	ia	strictly positive integer	[54 to 68]	O	L3
		units of files specific Lagrangian modelling, 15-dimension array useful in case of Lagrangian modelling			

## 9.1.2 Post-processing for *EnSight* or other tools

### NOTES

- The format depends on the user choices, and most options are defined using the GUI or `cs_user_postprocess.c`.
- The post-processing files can be of the following formats: *EnSight Gold*, *MED* or *CGNS*. The use of the two latter formats depends on the installation of the corresponding external libraries.
- For each quantity (problem unknown, preselected numerical variable or preselected physical parameter), the user specifies if a post-processing output is wanted. The output frequency can be set.

<b>ichrvr</b>	ia	-999, 0 or 1	[-999]	O	L1
		for each quantity defined at the cell centres (physical or numerical variable), indicator of whether it should be post-processed or not = -999: not initialised. By default, the post-processed quantities are the unknowns (pressure, velocity, $k$ , $\varepsilon$ , $R_{ij}$ , $\omega$ , $\varphi$ , $\bar{f}$ , scalars), density, turbulent viscosity and the time step if is not uniform = 0: not post-processed = 1: post-processed useful if and only if the variable is defined at the cell centers: calculation variable, physical property (time step, density, viscosity, specific heat) or turbulent viscosity if $iturb \geq 10$			
<b>ipstdv</b>	i	integer $\geq 1$ : see below	[ipstyp*ipstcl*ipstft]	O	L1
		indicates the data to post-process on the boundary mesh (the boundary mesh must have been activated with <code>ichrbo=1</code> ). The value of <code>ipstdv</code> is the product of the following integers, depending on the variables that should be post-processed: ipstyp: $y^+$ at the boundary ipstcl: value of the variables at the boundary (using the boundary conditions but without reconstruction) ipstft: thermal flux at the boundary ( $W m^{-2}$ ), if a thermal scalar has been			

defined (`iscalt`)

For instance, with `ipstdv=ipstyp*ipstcl, y+` and the variables will be post-processed at the boundaries.

With `ipstdv=1`, none of these data are post-processed at the boundaries.

always useful if `ichrbo=1`

### 9.1.3 Chronological records of the variables on specific points

#### STANDARD USE THROUGH INTERFACE OR `CS_USER_PARAMETERS.F90`

For each quantity (problem unknown, preselected numerical variable or preselected physical parameter), the user indicates whether chronological records should be generated, the output period and the position of the probes. The code produces chronological records at the cell centers located closest to the geometric points defined by the user by means of their coordinates. For each quantity, the number of probes and their index-numbers must be specified (it is not mandatory to generate all the variables at all the probes).

<b>ncapt</b>	i	positive or null integer	[0]	O	L1
		total number of probes (limited to <code>ncaptm=100</code> )			
		always useful			

<b>xyzcap</b>	ra	real numbers	[0.0]	O	L1
		3D-coordinates of the probes			
		the coordinates are written: <code>xyzcap(i,j)</code> , with $i = 1, 2 \text{ or } 3$ and $j \leq \text{ncapt}$			
		useful if and only if <code>ncapt &gt; 0</code>			

<b>ihisvr</b>	ia	-999, -1 or positive or null integer	[-999]	O	L1
		number <code>ihisvr(n, 1)</code> and index-numbers <code>ihisvr(n, j&gt;1)</code> of the record probes to be used for each variable, <i>i.e.</i> calculation variable or physical property defined at the cell centers. With <code>ihisvr(n, 1)=-999</code> or <code>-1</code> , <code>ihisvr(n, j&gt;1)</code> is useless.			

- `ihisvr(n, 1)`: number of record probes to use for the variable N
  - = -999: by default: chronological records are generated on all the probes if N is one of the main variables (pressure, velocity, turbulence, scalars), the local time step or the turbulent viscosity. For the other quantities, no chronological record is generated.

- = -1: chronological records are produced on all the probes

- = 0: no chronological record on any probe

- > 0: chronological record on `ihisvr(n, 1)` probes to be specified with

`ihisvr(n, j>1)`

always useful, must be inferior or equal to `ncapt`

- `ihisvr(n, j>1)`: index-numbers of the probes used for the variable  $n$   
(with  $j \leq \text{ihisvr}(n,1)+1$ )

- = -999: by default: if `ihisvr(n, 1) ≠ -999`, the code stops. Otherwise, refer to the description of the case `ihisvr(n, 1)=-999`

useful if and only if `ihisvr(n, 1) > 0`

The condition `ihisvr(n, j) ≤ ncapt` must be respected.

For an easier use, it is recommended to simply specify `ihisvr(n,1)=-1` for all the interesting variables.

<b>emphis</b>	c	string of less than 80 characters	[./]	O	L3
		directory in which the potential chronological record files generated by the Kernel will be written (path related to the execution directory)			
		it is recommended to keep the default value and, if necessary, to modify the launch			

script to copy the files in the alternate destination directory  
 useful if and only if chronological record files are generated (*i.e.* there is **n** for which **ihisvr(n, 1)  $\neq$  0**)

<b>nthist</b>	i	-1 or strictly positive integer output period of the chronological record files = -1: no output > 0: period (every <b>nthist</b> time step) The default value is -1 if there is no chronological record file to generate (if there is no probe, <b>ncapt</b> = 0, or if <b>ihisvr(n, 1)=0</b> for all the variables) and 1 otherwise If chronological records are generated, it is usually wise to keep the default value <b>nthist=1</b> , in order to avoid missing any high frequency evolution (unless the total number of time steps is much too big) useful if and only if chronological record files are generated ( <i>i.e.</i> there are probes ( <b>ncapt</b> >0) there is <b>n</b> for which <b>ihisvr(n, 1) <math>\neq</math> 0</b> )	[1 or -1]	O	L1
<b>nthsav</b>	i	-1 or positive or null integer saving period the chronological record files (they are first stored in a temporary file and then saved every <b>nthsav</b> time step) = 0: by default (4 times during a calculation) = -1: saving at the end of the calculation > 0: period (every <b>nthsav</b> time step) During the calculation, the user can read the chronological record files in the execution directory when they have been saved, <i>i.e.</i> at the first time step, at the tenth time step and when the time step number is a multiple of <b>nthsav</b> (multiple of ( <b>ntmabs-ntpabs</b> )/4 if <b>nthsav=0</b> ) <i>Note: using the <b>control_file</b> file allows to update the value of <b>ntmabs</b>. Hence, if the calculation is at the time step <b>n</b>, the saving of the chronological record files can be forced by changing <b>ntmabs</b> to <b>ntpabs+4(n+1)</b> using <b>control_file</b>; after the files have been saved, <b>ntmabs</b> can be reset to its original value, still using <b>control_file</b>.</i> useful if and only if chronological record files are generated ( <i>i.e.</i> there are probes ( <b>ncapt</b> >0) there is <b>n</b> for which <b>ihisvr(n, 1) <math>\neq</math> 0</b> )	[0]	O	L3

#### NON-STANDARD USE THROUGH **USHIST**

(see p. 96)

<b>impush</b>	ia	strictly positive integer units of the user chronological record files useful if and only if the subroutine <b>ushist</b> is used	[33 to 32+nushmx=49]	O	L3
<b>ficush</b>	ca	strings of 13 characters names of the user chronological record files. In the case of a non-parallel calculation, the suffix applied the file name is a three digit number: <b>ush001</b> , <b>ush002</b> , <b>ush003</b> ... In the case of a parallel-running calculation, the processor index-number is added to the suffix. For instance, for a calculation running on two processors: <b>ush001.n_0001</b> , <b>ush002.n_0001</b> , <b>ush003.n_0001</b> ... and <b>ush001.n_0002</b> , <b>ush002.n_0002</b> , <b>ush003.n_0002</b> ... The opening, closing, format and location of these files must be managed by the user. useful if and only if the subroutine <b>ushist</b> is used	[ush* or ush*.n.*]	O	L2

### 9.1.4 Time averages

The code allows the calculation of time averages of the type  $\langle f_1 * f_2 \dots * f_n \rangle$ . The variables  $f_i$  (defined at the cell centres) which may be taken into account are the followings:

EDF R&D	<b>Code_Saturne version 3.2.0 practical user's guide</b>	Code_Saturne documentation Page 143/202
---------	--	---

- the solved calculation variables (velocity, pressure ...),
- the auxiliary variables from the array **propce** (density and physical properties when they are variable in space).

The averages are treated like auxiliary variables defined at the cell centers and stored in the **propce** array. The standard post-processing actions may therefore be activated, like the writing in the listing or the output of result files (EnSight, MED, ...). However, if the user wants to manipulate the averages in a more advanced way, it is recommended to refer first to the user examples for subroutine **cs\_user\_extra\_operations**. Indeed, the **propce** array does not contain the time averages directly, but only the cumulated value of the product  $f_1 * f_2 * \dots * f_n$  of the selected variables  $f_i$ . The division by the cumulated duration is done only before the writing of the results. See also page 38.

To calculate  $p$  time averages of the type  $\langle f_1 * f_2 * \dots * f_{n(imom)} \rangle$ , the user must:

- make sure that  $p \leq \text{nbmomx}$  (do not overstep the maximum number of averages),
- make sure that  $n(imom) \leq \text{ndgmox}$  for every average **imom** (do not overstep the maximum degree, *i.e* the maximum number of variables which may compose an average),
- define every average **imom** ( $1 \leq \text{imom} \leq p$ , without skipping any index-number) by marking out the  $n(imom)$  variables which form it by means of the array **idfmom(ii,imom)** (with  $1 \leq ii \leq n(imom)$ ),
- define for each average **imom** the time step number at which the calculation of the cumulated value must begin, by means of the array **ntdmom(imom)**.

The total number of averages ( $p = \text{nbmomt}$ ) is automatically determined by the code from the values of **idtmom**. The user must not specify it.

<b>idfmom</b>	ia	0, $\pm$ variable index-number	[0]	O	L1
Index-number of the variables composing a time average of the type $\langle f_1 * f_2 * \dots * f_n \rangle$ . For every time average <b>imom</b> to calculate:					
- if <b>idfmom(ii,imom)</b> is positive, it refers to the index-number of a solved variable (stored in the array <b>rtp</b> ), like for instance a velocity component ( <b>iu</b> , <b>iv</b> , <b>iw</b> ) or the pressure ( <b>ipr</b> )					
- if <b>idfmom(ii,imom)</b> is negative, it refers to the index-number of an auxiliary variable (stored in <b>propce</b> ), like for instance the density ( <b>idfmom(ii,imom)=-irom</b> )					
useful if and only if the user wants to calculate time averages					
<b>ntdmom</b>	ia	integer	[-1]	O	L1
For every average <b>imom</b> to calculate, absolute time step number at which the calculation should begin. The value -1 means "never". Every strictly negative value (in particular -1) will be considered an error and cause the calculation to stop (because the user is supposed to want to calculate the averages he has defined)					
useful if and only if the user wants to calculate time averages					
<b>imoold</b>	ia	-2, $1 \leq \text{integer} \leq \text{jbmomt}$	[-2]	O	L1
Correspondence table of the averages in the case of a calculation restart. In this case, for every average <b>imom</b> in the current calculation ( $1 \leq \text{imom} \leq \text{nbmomx}$ ), <b>imoold(imom)</b> gives the index-number of the corresponding average in the previous calculation (in which <b>jbmomt</b> averages were calculated).					
- if <b>imoold(imom) = -2</b> , the user lets the code automatically determine the correspondence. By default, the average <b>ii</b> in the current calculation will correspond to the average <b>ii</b> in the previous calculation, if it existed. Otherwise, <b>ii</b> will be a new average.					

- if `imooldd(imom) = -1`, the average is reset to zero.  
- if `imooldd(imom) = kk`, the average `imom` will correspond to the average `kk=imooldd(imom)` in the previous calculation.  
useful if and only if the user wants to calculate averages. Allows to add or suppress some averages, to reset them, to change their order, ...  
*Warning: if the calculation is not a restart, imooldd must not be specified (its value must remain -2)*

### 9.1.5 Others

<b>impusr</b>	ia	strictly positive integer	[70 to 69+nusrmx=79]	O	L3
		unit numbers for potential user specified files			
		useful if and only if the user needs files (therefore always useful, by security)			
<b>ficusr</b>	ca	string of 13 characters	[usrf* or usrf*.n.*]	O	L1
		name of the potential user specified files. In the case of a non-parallel calculation, the suffix applied the file name is a two digit number: from <code>usrf01</code> to <code>usrf10</code> . In the case of a parallel-running calculation, the four digit processor index-number is added to the suffix. For instance, for a calculation running on two processors: from <code>usrf01.n_0001</code> to <code>usrf10.n_0001</code> and from <code>usrf01.n_0002</code> to <code>usrf10.n_0002</code> . The opening, closing, format and location of these files must be managed by the user.			
		useful if and only if the user needs files (therefore always useful, by security)			
<b>ilisvr</b>	ia	-999, 1 or 0	[-999]	O	L1
		for every quantity (variable, physical or numerical property ...), indicator concerning the writing in the execution report file			
		= -999: automatically converted into 1 if the concerned quantity is one of the main variables (pressure, velocity, turbulence, scalar), the density, the time step if <code>idtvar</code> $\neq$ 0 or the turbulent viscosity. Otherwise converted into 0.			
		= 1: writing in the execution listing.			
		= 0: no writing.			
		always useful			
<b>iwarni</b>	ia	integer	[0]	O	L1
		<code>iwarni(ivar)</code> characterises the level of detail of the outputs for the variable <code>ivar</code> (from 1 to <code>nvar</code> ). The quantity of information increases with its value.			
		Impose the value 0 or 1 for a reasonable listing size. Impose the value 2 to get a maximum quantity of information, in case of problem during the execution.			
		always useful			
<b>nomvar</b>	ca	string of less than 80 characters	[“”]	O	L1
		name of the variables (unknowns, physical properties ...): used in the execution listing, in the post-processing files, etc.			
		“”: not initialised (the code chooses the manes by default)			
		It is recommended not to define variable names of more than 16 characters, to get a clear execution listing (some advanced writing levels take into account only the first 16 characters).			
		always useful			
<b>ntlist</b>	i	-1 or strictly positive integer	[1]	O	L1
		writing period in the execution report file			

= -1: no writing

> 0: period (every `ntlist` time step)

The value of `ntlist` must be adapted according to the number of iterations carried out in the calculation. Keeping `ntlist` to 1 will indeed provide a maximum volume of information, but if the number of time steps is too large, the execution report file might become too big and unusable (problems with disk space, memory problems while opening the file with a text editor, problems finding the desired information in the file, ...).

always useful

<b>ntsuit</b>	i	-1, 0 or positive or null integer	[0]	O	L3
---------------	---	-----------------------------------	-----	---	----

saving period of the restart files

= -2: no restart at all

= -1: only at the end of the calculation

= 0: by default (four times during the calculation)

> 0: period

always useful

## 9.2 Numerical options

### 9.2.1 Calculation management

<b>iecaux</b>	i	0 or 1	[1]	O	L2
---------------	---	--------	-----	---	----

indicates the writing (=1) or not (=0) of the auxiliary calculation restart file

always useful

<b>ileaux</b>	i	0 or 1	[1]	O	L2
---------------	---	--------	-----	---	----

indicates the reading (=1) or not (=0) of the auxiliary calculation restart file

useful only in the case of a calculation restart

<b>inpdt0</b>	i	0 or 1	[0]	O	L1
---------------	---	--------	-----	---	----

indicates the calculation mode: 1 for a zero time step control calculation, *i.e.* without solving the transport equations, and 0 for a standard calculation.

In case of a calculation using the control mode (`inpdt0=1`), when the calculation is not a restart, the equations are not solved, but the physical properties and the boundary conditions are calculated. When the calculation is a restart, the physical properties and the boundary conditions are those read from the restart file (note: in the case of a second-order time scheme, the mass flow is modified as if a normal time step was realised: the mass flow generated in an potential post-processing is therefore not the mass flow read from the restart file).

In the control mode (`inpdt0=1`), the variable `ntmabs` is not used.

In the standard mode (`inpdt0=0`), the code solves the equations at least once, even if `ntmabs=0`.

always useful

<b>isuite</b>	i	0 or 1	[0]	C	L1
---------------	---	--------	-----	---	----

indicator of a calculation restart (=1) or not (=0)

always useful. This value is set automatically by the code, depending on whether a restart directory is present, and should not be modified by the user

<b>ntcabs</b>	i	integer	[ntpabs]	O	L3
---------------	---	---------	----------	---	----

current time step number



always useful

**ntcabs** is initialised and updated automatically by the code, its value is not to be modified by the user

**ntmabs**      i      integer > **ntpabs**      [10]      C      L1  
number of the last time step after which the calculation stops. It is an absolute number: for the restart calculations, **ntmabs** takes into account the number of time steps of the previous calculations. For instance, after a first calculation of 3 time steps, a restart file of 2 time steps is realised by setting **ntmabs**=3+2=5  
always useful

**ntpabs**      i      integer      [0, read]      O      L3  
number of the last time step in the previous calculation. In the case of a restart calculation, **ntpabs** is read from the restart file. Otherwise it is initialised to 0  
always useful  
**ntpabs** is initialised automatically by the code, its value is not to be modified by the user

**tmarus**      r      -1 or strictly positive real      [-1]      O      L3  
margin in seconds on the remaining CPU time which is necessary to allow the calculation to stop automatically and write all the required results (for the machines having a queue manager)  
= -1: calculated automatically  
> 0: margin defined by the user  
always useful, but the default value should not be changed unless absolutely necessary.

**ttcabs**      r      positive or null real number      [tntpabs]      O      L3  
physical simulation time at the current time step. For the restart calculations, **ttcabs** takes into account the physical time of the previous calculations.  
If the time step is uniform (**idtvar**=0 or 1), **ttcabs** increases of **dt** (value of the time step) at each iteration. If the time step is non-uniform (**idtvar**=2), **ttcabs** increases of **dtref** at each time step.  
always useful  
**ttcabs** is initialised and updated automatically by the code, its value is not to be modified by the user

**ttpabs**      r      positive or null real number      [0, read]      O      L3  
simulation physical time at the last time step of the previous calculation. In the case of a restart calculation, **ttpabs** is read from the restart file. Otherwise it is initialised to 0.  
always useful  
**ttpabs** is initialised automatically by the code, its value is not to be modified by the user

## 9.2.2 Scalar unknowns

**iscold**      ia      -999,  $1 \leq \text{integer} \leq \text{jscal}$       [-999]      O      L1  
correspondence table of the scalars in the case of a calculation restart. For a calculation restart with **nscal** scalars, **iscold**(**iscal**) gives, for every scalar **iscal** of the current calculation ( $1 \leq \text{iscal} \leq \text{nscal}$ ), the index-number of the corresponding scalar in the previous calculation (in which **jscal** scalars were taken into account).

**iscold(iscal) = -999**: the code automatically determines the correspondence. By default, the following rules are applied:

- the user scalar **ii** of the current calculation is initialised by the the user scalar **ii** of the previous calculation, if this scalar existed already (otherwise, **ii** is a new scalar).

- the particular physics scalar **jj** is initialised by the particular physics scalar **jj** of the previous calculation if this scalar existed already (otherwise, **jj** is a new scalar).

**iscold(iscal) = kk**: the scalar **iscal** (user or particular physics scalar) is initialised by the scalar **kk=iscold(iscal)** of the previous calculation.

always useful. Allows to add or remove some scalars, to change the solving order, to change the physics, ...

**nscaus**      **i**       $0 \leq \text{integer} \leq \text{nsccmax}$       [0]      O      L1  
number of user scalars solutions of an advection equation  
always useful

**iscavr**      **ia**       $0, 1 \leq \text{integer} \leq \text{nscal}$       [0]      O      L1  
if the scalar **iscal** is the average of the square of the fluctuations of a scalar **kk**, then **iscavr(iscal)=kk**. Otherwise **iscavr(iscal)=0**. For **iscal** and **kk**, the user can only use index-numbers refering to user scalars ( $\leq \text{nscaus}$ ).  
always useful

**iscalt**      **ia**      -1 or integer  $> 0$       [-1]      O      L1  
**iscalt** is the index-number of the scalar representing the temperature or the enthalpy. If **iscalt=-1**, no scalar represents the temperature nor the enthalpy. When a specific physics module is activated (gas combustion, pulverised coal, electricity or compressible), the user must not modify **iscalt** (the choice is made automatically)<sup>31</sup>.  
useful if and only if **nscal**  $\geq 1$

**iscsth**      **ia**      -1, 0, 1, 2 or 3      [-10]      O      L1  
type of scalar  
= -10: not specified. By default, the code chooses **iscsth(iscal)=0** for the scalars apart from **iscalt**  
= -1: temperature in degrees Celsius (use only in case of radiation modelling)  
= 0: passive scalar  
= 1: temperature (in Kelvin if the radiation modelling is activated)  
= 2: enthalpy  
= 3: total energy (this value is automatically chosen by the code when using the compressible module, it must never be used otherwise and must never be specified by the user)  
useful if and only if **nscal**  $\geq 1$ . The distinction between **iscsth(iscal) = -1** or **1** (respectively degrees Celsius or Kelvin) is useful only in case of radiation modelling. For calculations without radiation modelling, use **iscsth(iscal)=1** for the temperature. When a particular physics module is activated (gas combustion, pulverised coal, electricity or compressible), the user must not modify **iscsth** (the choice is made automatically: the solved variable is the enthalpy or the total energy).  
It is also reminded that, in the case of a coupling with SYRTHES, the solved thermal variable should be the temperature (**iscsth(iscalt)=1** or **-1**). More precisely, everything is designed in the code to allow for the running of a calculation coupled with SYRTHES with the enthalpy as thermal variable (the correspondence and conversion

---

<sup>31</sup>in the case of the compressible module, **iscalt** does not correspond to the temperature nor enthalpy but to the total energy

is then specified by the user in the subroutine `usthht`). However this case has never been used in practice and has therefore not been tested. With the compressible model, it is possible to carry out calculations coupled with SYRTHES, although the thermal scalar represents the total energy and not the temperature.

<b>iclvfl</b>	ia	-1, 0, 1 or 2	[-1]	O	L3
for every scalar <code>iscal</code> representing the average of the square of the fluctuations of another scalar <code>ii=iscavr(iscal)</code> (noted $f$ ), indicator of the clipping method = -1: no clipping because the scalar does not represent the average of the square of the fluctuations of another scalar = 0: clipping to 0 for lower values = 1: clipping to 0 for lower values and to $(f - f_{min})(f_{max} - f)$ for higher values, where $f$ is the associated scalar, $f_{min}$ and $f_{max}$ its minimum and maximum values specified by the user ( <i>i.e.</i> <code>scamin(ii)</code> and <code>scamax(ii)</code> ) = 2: clipping to <code>max(0,scamin(iscal))</code> for lower values and to <code>scamax(iscal)</code> for higher values. <code>scamin</code> and <code>scamax</code> are limits specified by the user useful for the scalars <code>iscal</code> for which <code>iscavr(iscal)&gt;0</code> .					
<b>itbrrb</b>	i	0 or 1	[0]	O	L3
Reconstruction (=1) or not (=0) of the temperature, enthalpy or total energy value in the boundary cells. Useful in the case of coupling with SYRTHES and with radiation.					
<b>icpsyr</b>	ia	-999,0,1	[-999]	O	L3
For each scalar <code>iscal</code> , <code>icpsyr(iscal)</code> indicates if it is coupled with SYRTHES (=1) or not (=0). There can be only one coupled scalar per calculation. =-999: by default <ul style="list-style-type: none"> <li>• <code>icpsyr(iscal)=1</code> for the thermal scalar <code>iscal=(iscalt)</code> when a coupling with SYRTHES has been specified</li> <li>• <code>icpsyr(iscal)=0</code> otherwise</li> </ul> = 0: the scalar <code>iscal</code> is not coupled with SYRTHES = 1: the scalar <code>iscal</code> is coupled with SYRTHES useful in case of coupling with SYRTHES					

### 9.2.3 Definition of the equations

<b>istat</b>	ia	0 or 1	[1 or 0]	O	L2
for each unknown <code>ivar</code> to calculate, indicates if non-stationary terms are present ( <code>istat(ivar)=1</code> ) or not (0) in the matrices. By default, <code>istat</code> is set to 0 for the pressure (variable <code>ivar=ipr</code> ) or $\bar{f}$ in v2f modelling (variable <code>ivar=ifb</code> ) and set to 1 for the other unknowns. useful for all the unknowns					
<b>iconv</b>	ia	0 or 1	[1 or 0]	O	L2
for each unknown <code>ivar</code> to calculate, indicates if the convection is taken into account ( <code>iconv(ivar)=1</code> ) or not (0). By default, <code>iconv</code> is set to 0 for the pressure (variable <code>ivar=ipr</code> ) or $\bar{f}$ in v2f modelling (variable <code>ivar=ifb</code> ) and set to 1 for the other unknowns. useful for all the unknowns					
<b>idiff</b>	ia	0 or 1	[1]	O	L2
for each unknown <code>ivar</code> to calculate, indicates if the diffusion is taken into account					

(**idiff(ivar)**=1) or not (0)  
useful for all the unknowns

**idifft**      ia      0 or 1      [1]      O      L3  
for each unknown **ivar** to calculate, when diffusion is taken into account (**idiff(ivar)**=1), **idifft(ivar)** indicates if the turbulent diffusion is taken into account (**idifft(ivar)**=1) or not (0)  
useful for all the unknowns

**idircl**      ia      0 or 1      [1 or 0]      O      L3  
for each unknown **ivar** to calculate, indicates whether the diagonal of the matrix should be slightly shifted (**idircl(ivar)**=1) or not (0) if there is no Dirichlet boundary condition and if **istat**=0. Indeed, in such a case, the matrix for the general advection/diffusion equation is singular. A slight shift in the diagonal will make it invertible again.  
By default, **idircl** is set to 1 for all the unknowns, except  $\bar{f}$  in v2f modelling, since its equation contains another diagonal term that ensures the regularity of the matrix.  
useful for all the unknowns

**ivisse**      ia      0 or 1      [1]      O      L3  
indicates whether the source terms in transposed gradient and velocity divergence should be taken into account in the momentum equation. In the compressible module, these terms also account for the volume viscosity (cf. **viscv0** and **iviscv**):  

$$\partial_i [(\kappa - 2/3(\mu + \mu_t))\partial_k U_k] + \partial_j [(\mu + \mu_t)\partial_i U_j]$$
= 0: not taken into account  
= 1: taken into account  
always useful

## 9.2.4 Definition of the time advancement

**idtvar**      i      -1, 0, 1, 2      [0]      O      L1  
type of time step  
= 0: constant in time and spatially uniform  
= 1: variable in time and spatially uniform  
= 2: variable in time and in space  
= -1: steady-state algorithm  
If the numerical scheme is a second-order in time, only the option 0 is allowed.  
always useful

**idilat**      i      1, 2, 3, 4      [1]      O      L1  
Algorithm to take into account the density variation in time  
= 1: dilatable steady algorithm (default)  
= 2: dilatable unsteady algorithm  
= 3: low-Mach number algorithm  
= 4: non conservative algorithm for fire simulation  
always useful

**iptlro**      i      0 or 1      [0]      O      L2  
when density gradients and gravity are present, a local thermal time step can be calculated, based on the Brunt-Vaissala frequency. In numerical simulations, it is usually wise for the time step to be lower than this limit, otherwise numerical instabilities

may appear

**iptlro** indicates whether the time step should be limited to the local thermal time step (=1) or not (=0)

when **iptlro**=1, the listing shows the number of cells where the time step has been clipped due to the thermal criterium, as well as the maximum ratio between the time step and the maximum thermal time step. If **idtvar**=0, since the time step is fixed and cannot be clipped, this ratio can be larger than 1<sup>32</sup>. When **idtvar**>0, this ratio will be smaller than 1, except if the constraint **dtmin** has prevented the code from reaching a sufficiently low value for **dt**

useful when density gradients and gravity are present

<b>cdtvar</b>	r	strictly positive real number	[1]	O	L1
multiplicative factor applied to the time step for each scalar					
Hence, the time step used when solving the evolution equation for the variable is the time step used for the dynamic equations (velocity/pressure) multiplied by <b>cdtvar</b> . The size of the array <b>cdtvar</b> is <b>nvar</b> . For instance, the multiplicative coefficient applied to the scalar 2 is <b>cdtvar(isca(2))</b> . Yet, the value of <b>cdtvar</b> for the velocity components and the pressure is not used. Also, although it is possible to change the value of <b>cdtvar</b> for the turbulent variables, it is highly unrecommended					
useful if and only if <b>nscal</b> ≥ 1					
<b>coumax</b>	r	strictly positive real number	[1]	O	L1
target local or maximum Courant number in case of non-constant time step					
useful if <b>idtvar</b> ≠ 0					
<b>foumax</b>	r	strictly positive real number	[10]	O	L1
target local or maximum Fourier number in case of non-constant time step					
useful if <b>idtvar</b> ≠ 0					
<b>dtref</b>	r	strictly positive real number	[-grand*10]	C	L1
reference time step					
always useful.					
It is the time step value used in the case of a calculation run with a uniform and constant time step, <i>i.e.</i> <b>idtvar</b> =0 (restart calculation or not). It is the value used to initialise the time step in the case of an initial calculation run with a non-constant time step ( <b>idtvar</b> =1 or 2). It is also the value used to initialise the time step in the case of a restart calculation in which the type of time step has been changed (for instance, <b>idtvar</b> =1 in the new calculation and <b>idtvar</b> =0 or 2 in the previous calculation): see <b>cs.user_initialization</b>					
<b>dtmin</b>	r	positive or null real number	[0.1*dtref]	O	L2
lower limit for the calculated time step when non-constant time step is activated					
useful if <b>idtvar</b> ≠ 0					
<b>dtmax</b>	r	strictly positive real number	[1000*dtref]	O	L2
upper limit for the calculated time step when non-constant time step is activated					
useful if <b>idtvar</b> ≠ 0					
<b>varrdt</b>	r	strictly positive real number	[0.1]	O	L3
maximum allowed relative increase in the calculated time step value between two					

<sup>32</sup>it is then the user's choice to decide whether he should diminish DTREF or not

successive time steps (to ensure stability, any decrease in the time step is immediate and without limit)  
useful if `idtvar`  $\neq 0$

**relxst**      `r`       $0 < \text{real} \leq 1$       [0.9]      O      L2  
relaxation coefficient for the steady algorithm (`relaxp(iphas)=1`: no relaxation)  
useful if `idtvar=-1`

**relaxv**      `ra`       $0 \leq \text{real} \leq 1$       [0.7 or 1.]      O      L3  
for each variable `ivar`, relaxation coefficient of the variable. This relaxation parameter is only useful for the pressure with the unsteady algorithm (so as to improve the convergence in case of meshes of insufficient quality or and for some of the turbulent models (`iturb(iphas) = 20, 21, 50` or `60` and `ikecou(iphas)=0`; if `ikecou(iphas)=1`, `relaxv(ivar)` is not used, whatever its value may be). Default values are 0.7 for turbulent variables and 1. for pressure.  
It also stores the value of the relaxation coefficient when using the steady algorithm, deduced from the value of `relxst` (defaulting to `relaxv(ivar eq 1. - relxst)`)  
useful only for the pressure and for turbulent variables if and only if ( $k - \varepsilon$ ,  $v2f$  or  $k - \omega$  models without coupling) with the unsteady algorithm  
always useful with the steady algorithm

#### NON-CONSTANT TIME STEP

The calculation of the time step uses a reference time step `DTREF` (at the calculation beginning). Later, every time step, the time step value is calculated by taking into account the different existing limits, in the following order:

- `coumax`, `foumax`: the more restrictive limit between both is used (in the compressible module, the acoustic limitation is added),
- `varrdt`: progressive increase and immediate decrease in the time step,
- `iptlro`: limitation by the thermal time step,
- `dtmax` and `dtmin`: clipping of the time step to the maximum, then to the minimum limit.

### 9.2.5 Turbulence

**iturb**      `ia`      0, 10, 20, 21, 30, 31, 32, 40, 41, 42, 50, 51, 60, 70      O      L1  
indicator of the turbulence model `iturb`  
= -999: not initialised. This value is not allowed and must be modified by the user  
= 0: laminar  
= 10: mixing length (not validated)  
= 20:  $k - \varepsilon$   
= 21:  $k - \varepsilon$  with linear production (Laurence & Guimet)  
= 30:  $R_{ij} - \varepsilon$  "standard" LRR (Launder, Reece & Rodi)  
= 31:  $R_{ij} - \varepsilon$  SSG (Speziale, Sarkar & Gatski)  
= 32:  $R_{ij} - \varepsilon$  EBRSM (elliptic blending)  
= 40: LES (Smagorinsky model)  
= 41: LES (dynamic model)  
= 42: LES (WALE model)  
= 50:  $v2-f$ ,  $\varphi$ -model version  
= 51:  $v2-f$ , BL- $v2/k$  version  
= 60:  $k - \omega$ , SST version  
= 70: Spalart-Allmaras  
always useful

The  $k - \varepsilon$  (standard and linear production) and  $R_{ij} - \varepsilon$  (LRR and SSG) turbulence models implemented in *Code\_Saturne* are “High-Reynolds” models. It is therefore necessary to make sure that the thickness of the first cell neighboring the wall is larger than the thickness of the viscous sublayer (at the wall,  $y^+ > 2.5$  is required as a minimum, and preferably between 30 and 100)<sup>33</sup>. If the mesh does not respect this condition, the results may be biased (particularly if thermal processes are involved). Using scalable wall-functions (cf. key word **ideuch**) may help avoiding this problem.

The v2-f model is a “Low-Reynolds” model, it is therefore necessary to make sure that the thickness of the first cell neighboring the wall is smaller than the thickness of the viscous sublayer ( $y^+ < 1$ ).

The  $k - \omega$  SST model provides correct results whatever the thickness of the first cell. Yet, it requires the knowledge of the distance to the wall in every cell of the calculation domain. The user may refer to the key word **icdpar** for more details about the potential limitations.

The  $k - \varepsilon$  model with linear production allows to correct the known flaw of the standard  $k - \varepsilon$  model which overestimates the turbulence level in case of strong velocity gradients (stopping point).

With LES, the wall functions are usually not greatly adapted. It is generally more advisable (if possible) to refine the mesh towards the wall so that the first cell is in the viscous sub-layer, where the boundary conditions are simple natural no-slip conditions.

Concerning the LES model, the user may refer to the subroutine **ussmag** for complements about the dynamic model. Its usage and the interpretation of its results require particular attention. In addition, the user must pay further attention when using the dynamic model with the least squares method based on a partial extended neighbourhood (**imrga**=3). Indeed, the results may be degraded if the user does not implement his own way of averaging the dynamic constant in **ussmag** (*i.e.* if the user keeps the local average based on the extended neighbourhood).

**iturb**      ia      0, 10, 20 or 30      [0]      O      L1  
indicator of the turbulent flux model  $\overline{Y'u'}$  for any scalar  $Y$   
= 0: Simple Gradient Hypothesis (default value)  
= 10: Generalized Gradient Hypothesis  
= 20: Algebraic Flux Model  
= 30: Differential Flux Model (transport equations of the turbulent flux)  
GGDH, AFM and DFM are only available when a second order closure is used **iturb**=30, 31, 32.

**ideuch**      ia      0, 1 or 2      [0 or 1]      O      L2  
indicates the type of wall function is used for the velocity boundary conditions on a frictional wall.  
= 0: one-scale model  
= 1: two-scale model  
= 2: scalable wall function  
**ideuch** is initialised to 0 for **iturb**=0, 10, 40 or 41 (laminar, mixing length, LES).  
**ideuch** is initialised to 1 for **iturb**=20, 21, 30, 31 or 60 ( $k - \varepsilon$ ,  $R_{ij} - \varepsilon$  LRR,  $R_{ij} - \varepsilon$  SSG and  $k - \omega$  SST models).  
The v2f model (**iturb**=50) is not designed to use wall functions (the mesh must be “low Reynolds”).  
The value **ideuch**=1 is not compatible with **iturb**=0, 10, 40 or 41 (laminar, mixing length and LES).  
Concerning the  $k - \varepsilon$  and  $R_{ij} - \varepsilon$  models, the two-scales model is usually at least as satisfactory as the one-scale model.  
The scalable wall function allows to virtually “shift” the wall when necessary in order to be always in a logarithmic layer. It is used to make up for the problems related to the use of High-Reynolds models on very refined meshes.  
useful if **iturb** is different from 50

<sup>33</sup>While creating the mesh,  $y^+ = \frac{y u_*}{\nu}$  is generally unknown. It can be roughly estimated as  $\frac{y U}{10\nu}$ , where  $U$  is the characteristic velocity,  $\nu$  is the kinematic viscosity of the fluid and  $y$  is the mid-height of the first cell near the wall.

**ilogpo**      ia      0 or 1      [1]      O      L3  
type of wall function used for the velocity: power law (**ilogpo**=0, deprecated) or logarithmic law (**ilogpo**=1)  
always useful

**yp1uli**      ra      real number > 0      [1/xkappa, 10.88]      O      L3  
limit value of  $y^+$  for the viscous sublayer  
**yp1uli** depends on the chosen wall function: it is initialised to 10.88 for the scalable wall function (**ideuch**=2), otherwise it is initialised to  $1/\kappa \approx 2,38$   
In LES, **yp1uli** is taken by default to be 10.88  
always useful

#### $k - \varepsilon$ , $k - \varepsilon$ WITH LINEAR PRODUCTION, v2-f AND $k - \omega$ SST

**igrake**      ia      0 or 1      [1]      O      L1  
indicates if the terms related to gravity in the equations of  $k$  and  $\varepsilon$  or  $\omega$  are taken into account (**igrake**=1) or not (0)  
useful if and only if **iturb** = 20, 21, 50 or 60, (**gx**, **gy**, **gz**)  $\neq$  (0,0,0) and the density is not uniform

**igrhok**      ia      0 or 1      [0]      O      L2  
indicates if the term  $\frac{2}{3}\nabla\rho k$  is taken into account (**igrhok**=1) or not (0) in the velocity equation  
useful if and only if **iturb** = 20, 21, 50 or 60.  
This term may generate non-physical velocities at the wall. When it is not explicitly taken into account, it is implicitly included into the pressure.

**ikecou**      ia      0 or 1      [0 or 1]      O      L3  
indicates if the coupling of the source terms of  $k$  and  $\varepsilon$  or  $k$  and  $\omega$  is taken into account (**ikecou**=1) or not (0)  
if **ikecou**=0 in  $k - \varepsilon$  model, the term in  $\varepsilon$  in the equation of  $k$  is made implicit  
**ikecou** is initialised to 0 if **iturb** = 21 or 60, and to 1 if **iturb**= 20  
**ikecou**=1 is forbidden when using the v2f model (**iturb**=50)  
useful if and only if **iturb** = 20, 21 or 60 ( $k - \varepsilon$  and  $k - \omega$  models)

**iclkep**      ia      0 or 1      [0]      O      L3  
indicates the clipping method used for  $k$  and  $\varepsilon$ , for the  $k - \varepsilon$  and v2f models  
= 0: clipping in absolute value  
= 1: clipping from physical relations  
useful if and only if **iturb** = 20, 21 or 50 ( $k - \varepsilon$  and v2f models). The results obtained with the method corresponding to **iclkep**=1 showed in some cases a substantial sensitivity to the values of the length scale **almax**.  
The option **iclkep**=1 is therefore not recommended, and, if chosen, must be used cautiously.

**irccor**      i      0 or 1      [0]      O      L2  
Activation of rotation/curvature correction for an eddy viscosity turbulence models.  
= 0: activated  
= 1: not activated  
useful if and only if **iturb** = 20, 21, 50, 51, 60, 70 (eddy viscosity models)



**itycor**      i      1 or 2      [1 or 2]      O      L2  
Type of rotation/curvature correction for eddy viscosity turbulence models:  
= 1: Cazalhou correction (default when **irccor**=1 and **iturb**=20, 21 or 50,  
51)  
= 2: Spalart-Shur correction (default when **irccor**=1 and **iturb**=60 or 70)

### $R_{ij} - \varepsilon$ (LLR AND SSG)

**iclpnr**      ia      0 or 1      [0]      O      L3  
indicates if  $R_{ij}$  is made partially implicit (**iclpnr**=1) or not (0) in the wall boundary conditions.  
useful if and only if **iturb** = 30 or 31 ( $R_{ij} - \varepsilon$  model)

**iclsyr**      ia      0 or 1      [0]      O      L3  
indicates if  $R_{ij}$  is made partially implicit (**iclsyr**=1) or not (0) in the symmetry boundary conditions.  
useful if and only if **iturb** = 30 or 31 ( $R_{ij} - \varepsilon$  model)

**idifre**      ia      0 or 1      [1]      O      L3  
complete (**idifre**=1) or simplified (0) taking into account of the diagonals of the diffusion tensors of  $R_{ij}$  and  $\varepsilon$ , for the LLR model.  
useful if and only if **iturb** = 30 (LLR  $R_{ij} - \varepsilon$  model)

**igrari**      ia      0 or 1      [1]      O      L1  
indicates if the terms related to gravity are taken into account (**igrari**=1) or not (0) in the equations of  $R_{ij} - \varepsilon$ .  
useful if and only if **iturb** = 30 or 31 and (**gx**, **gy**, **gz**)  $\neq$  (0,0,0) ( $R_{ij} - \varepsilon$  model with gravity) and the density is not uniform

**irijec**      ia      0 or 1      [0]      O      L2  
indicates if the wall echo terms in  $R_{ij} - \varepsilon$  LLR model are taken into account (**irijec**=1) or not (0).  
useful if and only if **iturb** = 30 ( $R_{ij} - \varepsilon$  LLR).  
It is not recommended to take these terms into account: they have an influence only near the walls, their expression is hardly justifiable according to some authors and, in the configurations studied with *Code\_Saturne*, they did not bring any improvement in the results.  
In addition, their use induces an increase in the calculation time.  
The wall echo terms imply the calculation of the distance to the wall for every cell in the domain. See **icdpar** for potential restrictions due to this.

**irijnu**      ia      0 or 1      [0]      O      L3  
addition (**irijnu**=1) or not (0) of a turbulent viscosity in the matrix of the incremental system solved for the velocity in  $R_{ij} - \varepsilon$  models. The goal is to improve the stability of the calculation. The usefulness of **irijnu**=1 has however not been clearly demonstrated.  
Since the system is solved in incremental form, this extra turbulent viscosity does not change the final solution for steady flows. However, for unsteady flows, the parameter **nswrsm** should be increased.  
useful if and only if **iturb** = 30 or 31 ( $R_{ij} - \varepsilon$  model).

**irijrb**      ia      0 or 1      [0]      O      L3  
reconstruction (**irijrb**=1) or not (0) of the boundary conditions at the walls for  $R_{ij}$  and  $\varepsilon$ .  
useful if and only if **iturb** = 30 or 31 ( $R_{ij} - \varepsilon$  model)

### LES

**ivrtex**      i      0 or 1      [0]      O      L1  
activates (=1) or not (=0) the generation of synthetic turbulence at the different inlet boundaries with the LES model (generation of unsteady synthetic eddies)  
useful if **iturb**=40, 41 or 42  
this key word requires the completion of the routine **usvort**

**isuivo**      i      0 or 1      [suite]      O      L1  
for the vortex method, indicates whether the synthetic vortices at the inlet should be initialised (=0) or read from the restart file.  
useful if **iturb**=40, 41, 42 and **ivrtex**=1

**isuisy**      i      0 or 1      [suite]      O      L1  
Reading of the LES inflow module restart file.  
= 0: not activated  
= 1: activated  
If **isuisy**=1, synthetic fluctuations are not re-initialized in case of restart calculation.  
useful if **iturb**=40, 41 or 42

**idries**      ia      0 or 1      [0,1]      O      L2  
**idries** activates (1) or not (0) the van Driest wall-damping for the Smagorinsky constant (the Smagorinsky constant is multiplied by the damping function  $1 - e^{-y^+/\text{cdries}}$ , where  $y^+$  designates the non-dimensional distance to the nearest wall). The default value is 1 for the Smagorinsky model and 0 for the dynamic model.  
the van Driest wall-damping requires the knowledge of the distance to the nearest wall for each cell in the domain. Refer to key word **icdpar** for potential limitations  
useful if and only if **iturb** = 40 or 41

**cdries**      ra      real number > 0      [26]      O      L3  
**cdries** is the constant appearing in the van Driest damping function applied to the Smagorinsky constant:  $1 - e^{-y^+/\text{cdries}}$   
useful if and only if **iturb** = 40 or 41

**csmago**      ra      real number > 0      [0.065]      O      L2  
**csmago** is the Smagorinsky constant used in the Smagorinsky model for LES  
the sub-grid scale viscosity is calculated by  $\mu_{sg} = \rho C_{smago}^2 \bar{\Delta}^2 \sqrt{2\bar{S}_{ij}\bar{S}_{ij}}$  where  $\bar{\Delta}$  is the width of the filter and  $\bar{S}_{ij}$  the filtered strain rate  
useful if and only if **iturb** = 40

**smagmx**      ra      real number > 0      [10\*csmago]      O      L3  
**smagmx**<sup>2</sup> is the maximum allowed value for the variable  $C$  appearing in the LES dynamic model (the “square” comes from the fact that the variable of the dynamic model corresponds to the square of the constant of the Smagorinsky model). Any larger value

EDF R&D	<b>Code_Saturne version 3.2.0 practical user's guide</b>	Code_Saturne documentation Page 156/202
---------	--	---

yielded by the calculation procedure of the dynamic model will be clipped to **smagmx**<sup>2</sup> useful if and only if **iturb** = 41

<b>xlesfl</b>	ra	real number > 0	[2]	O	L3
<b>xlesfl</b> is a constant used to define, for each cell $\Omega_i$ , the width of the (implicit) filter: $\overline{\Delta} = xlesfl(ales *  \Omega_i )^{bles}$ useful if and only if <b>iturb</b> = 40 or 41					
<b>ales</b>	ra	real number > 0	[1]	O	L3
<b>ales</b> is a constant used to define, for each cell $\Omega_i$ , the width of the (implicit) filter: $\overline{\Delta} = xlesfl(ales *  \Omega_i )^{bles}$ useful if and only if <b>iturb</b> = 40 or 41					
<b>bles</b>	ra	real number > 0	[1/3]	O	L3
<b>bles</b> is a constant used to define, for each cell $\Omega_i$ , the width of the (implicit) filter: $\overline{\Delta} = xlesfl(ales *  \Omega_i )^{bles}$ useful if and only if <b>iturb</b> = 40 or 41					
<b>xlesfd</b>	ra	real number > 0	[1.5]	O	L3
<b>xlesfd</b> is the constant used to define, for each cell $\Omega_i$ , the width of the explicit filter used in the framework of the LES dynamic model: $\widetilde{\Delta} = xlesfd \overline{\Delta}$ useful if and only if <b>iturb</b> = 41					

## 9.2.6 Time scheme

By default, the standard time scheme is a first-order. A second-order scheme is activated automatically with LES modelling. On the other hand, when “specific physics” (gas combustion, pulverised coal, compressible module) are activated, the second-order scheme is not allowed.

In the current version, the second-order time scheme is not compatible with the estimators (**iescal**), the velocity-pressure coupling (**ipucou**), the modelling of hydrostatic pressure (**icalhy** and **iphydr**) and the time- or space-variable time step (**idtvar**).

Also, in the case of a rotation periodicity, a proper second-order is not ensured for the velocity, but calculations remain possible.

It is recommended to keep the default values of the variables listed below. Hence, in standard cases, the user does not need to specify these options.

<b>ischtp</b>	ia	1 or 2	[1 or 2]	O	L2
<b>ischtp</b> indicates the order of the activated time scheme (this indicator allows the code to automatically complete the other indicators related to the time scheme) = 1: first-order = 2: second-order when <b>ischtp</b> =2, the physical properties are by default not second-order. It is possible to modify this by means of the following indicators. due to specific coupling between certain variables, the source terms in the turbulence equations (except convection and diffusion) cannot be second order, except with the $R_{ij}$ models (cf. key word <b>isto2t</b> ) by default, <b>ischtp</b> is initialised to 2 with the LES model and 1 otherwise always useful					
<b>istmpf</b>	ia	0, 1 or 2	[0 or 1]	O	L3
<b>istmpf</b> specifies the time scheme activated for the mass flow. The chosen value for					

**istmpf** will automatically determine the value given to the variable **thetf1**

= 0: "explicit" first-order: the mass flow calculated at the previous time step ("n") is used in the convective terms of all the equations (momentum, turbulence and scalars)

= 1: "standard" first-order: the mass flow calculated at the previous time step ("n") is used in the convective terms of the momentum equation, and the updated mass flow (time "n+1") is used in the equations of turbulence and scalars

= 2: second-order: the mass flow used in the momentum equations is extrapolated at "n+**thetf1**" (=n+1/2) from the values at the two former time steps (Adams Bashforth); the mass flow used in the equations for turbulence and scalars is interpolated at time "n+**thetf1**" (=n+1/2) from the values at the former time step and at the newly calculated "n+1" time step.

by default, **istmpf**=2 is used in the case of a second-order time scheme (if **ischtp**=2) and **istmpf**=1 otherwise

always useful

**isno2t**      ia      0, 1 or 2      [0 or 1]      O      L3

**isno2t** specifies the time scheme activated for the source terms of the momentum equation, apart from convection and diffusion (for instance: head loss, transposed gradient, ...).

= 0: "standard" first-order: the terms which are linear functions of the solved variable are implicit and the others are explicit

= 1: second-order: the terms of the form  $S_i\phi$  which are linear functions of the solved variable  $\phi$  are expressed as second-order terms by interpolation (according to the formula  $(S_i\phi)^{n+\theta} = S_i^n[(1-\theta)\phi^n + \theta\phi^{n+1}]$ ,  $\theta$  being given by the value of **thetav** associated with the variable  $\phi$ ); the other terms  $S_e$  are expressed as second-order terms by extrapolation (according to the formula  $(S_e)^{n+\theta} = [(1+\theta)S_e^n - \theta S_e^{n-1}]$ ,  $\theta$  being given by the value of **thetsn**=0.5)

= 2: the linear terms  $S_i\phi$  are treated in the same way as when **isno2t**=1; the other terms  $S_e$  are extrapolated according to the same formula as when **isno2t**=1, but with  $\theta$ =**thetsn**=1

by default, **isno2t** is initialised to 1 (second-order) when the selected time scheme is second-order (**ischtp**=2), otherwise to 0.

always useful

**isto2t**      ia      0, 1 or 2      [0]      O      L3

**isto2t** specifies the time scheme activated for the source terms of the turbulence equations (related to  $k$ ,  $R_{ij}$ ,  $\varepsilon$ ,  $\omega$ ,  $\varphi$ ,  $\bar{f}$ ), apart from convection and diffusion.

= 0: "standard" first-order: the terms which are linear functions of the solved variable are implicit and the others are explicit

= 1: second-order: the terms of the form  $S_i\phi$  which are linear functions of the solved variable  $\phi$  are expressed as second-order terms by interpolation (according to the formula  $(S_i\phi)^{n+\theta} = S_i^n[(1-\theta)\phi^n + \theta\phi^{n+1}]$ ,  $\theta$  being given by the value of **thetav** associated with the variable  $\phi$ ); the other terms  $S_e$  are expressed as second-order terms by extrapolation (according to the formula  $(S_e)^{n+\theta} = [(1+\theta)S_e^n - \theta S_e^{n-1}]$ ,  $\theta$  being given by the value of **thetst**=0.5)

= 2: the linear terms  $S_i\phi$  are treated in the same way as when **isto2t**=1; the other terms  $S_e$  are extrapolated according to the same formula as when **isto2t**=1, but with  $\theta$ =**thetst**=1

due to certain specific couplings between the turbulence equations, **isto2t** is allowed the value 1 or 2 only for the  $R_{ij}$  models (**iturb**=30 or 31); hence, it is always initialised to 0.

always useful

<b>isso2t</b>	<p>ia      0, 1 or 2      [0 or 1]      O      L3</p> <p>for each scalar <b>iscal</b>, <b>isso2t(iscal)</b> specifies the time scheme activated for the source terms of the equation for the scalar, apart from convection and diffusion (for instance: variance production, user-specified terms, ...).</p> <p>= 0: "standard" first-order: the terms which are linear functions of the solved variable are implicit and the others are explicit</p> <p>= 1: second-order: the terms of the form <math>S_i\phi</math> which are linear functions of the solved variable <math>\phi</math> are expressed as second-order terms by interpolation (according to the formula <math>(S_i\phi)^{n+\theta} = S_i^n[(1-\theta)\phi^n + \theta\phi^{n+1}]</math>, <math>\theta</math> being given by the value of <b>thetav</b> associated with the variable <math>\phi</math>) ; the other terms <math>S_e</math> are expressed as second-order terms by extrapolation (according to the formula <math>(S_e)^{n+\theta} = [(1+\theta)S_e^n - \theta S_e^{n-1}]</math>, <math>\theta</math> being given by the value of <b>thetss(iscal)=0.5</b>)</p> <p>= 2: the linear terms <math>S_i\phi</math> are treated in the same way as when <b>isso2t=1</b>; the other terms <math>S_e</math> are extrapolated according to the same formula as when <b>isso2t=1</b>, but with <math>\theta=\mathbf{thetss(iscal)=1}</math></p> <p>by default, <b>isso2t(iscal)</b> is initialised to 1 (second-order) when the selected time scheme is second-order (<b>ischtp=2</b>), otherwise to 0.</p> <p>always useful</p>
<b>iroext</b>	<p>ia      0, 1 or 2      [0]      O      L3</p> <p><b>iroext</b> specifies the time scheme activated for the physical property <math>\phi</math> "density".</p> <p>= 0: "standard" first-order: the value calculated at the beginning of the current time step (from the variables known at the end of the previous time step) is used</p> <p>= 1: second-order: the physical property <math>\phi</math> is extrapolated according to the formula <math>\phi^{n+\theta} = [(1+\theta)\phi^n - \theta\phi^{n-1}]</math>, <math>\theta</math> being given by the value of <b>thetro=0.5</b></p> <p>= 2: first-order: the physical property <math>\phi</math> is extrapolated at <math>n+1</math> according to the same formula as when <b>iroext=1</b> but with <math>\theta=\mathbf{thetro=1}</math></p> <p>always useful</p>
<b>iviext</b>	<p>ia      0, 1 or 2      [0]      O      L3</p> <p><b>iviext</b> specifies the time scheme activated for the physical property <math>\phi</math> "total viscosity" (molecular+turbulent or sub-grid viscosities).</p> <p>= 0: "standard" first-order: the value calculated at the beginning of the current time step (from the variables known at the end of the previous time step) is used</p> <p>= 1: second-order: the physical property <math>\phi</math> is extrapolated according to the formula <math>\phi^{n+\theta} = [(1+\theta)\phi^n - \theta\phi^{n-1}]</math>, <math>\theta</math> being given by the value of <b>thetvi=0.5</b></p> <p>= 2: first-order: the physical property <math>\phi</math> is extrapolated at <math>n+1</math> according to the same formula as when <b>iviext=1</b>, but with <math>\theta=\mathbf{thetvi=1}</math></p> <p>always useful</p>
<b>icpext</b>	<p>ia      0, 1 or 2      [0]      O      L3</p> <p><b>icpext</b> specifies the time scheme activated for the physical property <math>\phi</math> "specific heat".</p> <p>= 0: "standard" first-order: the value calculated at the beginning of the current time step (from the variables known at the end of the previous time step) is used</p> <p>= 1: second-order: the physical property <math>\phi</math> is extrapolated according to the formula <math>\phi^{n+\theta} = [(1+\theta)\phi^n - \theta\phi^{n-1}]</math>, <math>\theta</math> being given by the value of <b>thetcp=0.5</b></p> <p>= 2: first-order: the physical property <math>\phi</math> is extrapolated at <math>n+1</math> according to the same formula as when <b>icpext=1</b>, but with <math>\theta=\mathbf{thetcp=1}</math></p> <p>always useful</p>
<b>ivsext</b>	<p>ia      0, 1 or 2      [0]      O      L3</p>

EDF R&D	<b>Code_Saturne version 3.2.0 practical user's guide</b>	Code_Saturne documentation Page 159/202
---------	--	---

for each scalar `iscal`, `ivsext(iscal)` specifies the time scheme activated for the physical property  $\phi$  “diffusivity”.

= 0: ”standard” first-order: the value calculated at the beginning of the current time step (from the variables known at the end of the previous time step) is used

= 1: second-order: the physical property  $\phi$  is extrapolated according to the formula  $\phi^{n+\theta} = [(1+\theta)\phi^n - \theta\phi^{n-1}]$ ,  $\theta$  being given by the value of `thetvs(iscal)`=0.5

= 2: first-order: the physical property  $\phi$  is extrapolated at  $n+1$  according to the same formula as when `ivsext`=1, but with  $\theta=\text{thetvs(iscal)}=1$   
always useful

**thetav**      ra       $0 \leq \text{real} \leq 1$       [1 or 0.5]      O      L3  
for each variable `ivar`, `thetav(ivar)` is the value of  $\theta$  used to express at the second-order the terms of convection, diffusion and the source terms which are linear functions of the solved variable (according to the formula  $\phi^{n+\theta} = (1-\theta)\phi^n + \theta\phi^{n+1}$ ). Generally, only the values 1 and 0.5 are used. The user is not allowed to modify this variable.  
= 1: first-order  
= 0.5: second-order  
Concerning the pressure, the value of `thetav` is always 1. Concerning the other variables, the value `thetav`=0.5 is used when the second-order time scheme is activated by `ischtp`=2 (standard value for LES calculations), otherwise `thetav` is set to 1.  
always useful

**thetfl**      ra       $0 \leq \text{real} \leq 1$       [0 or 0.5]      O      L3  
`thetfl` is the value of  $\theta$  used to interpolate the convective fluxes of the variables when a second-order time scheme has been activated for the mass flow (see `istmpf`)  
generally, only the value 0.5 is used. The user is not allowed to modify this variable.  
= 0.0: “explicit” first-order (corresponds to `istmpf`=0 or 1)  
= 0.5: second-order (corresponds to `istmpf`=2). The mass flux will be interpolated according to the formula  $Q^{n+\theta} = \frac{1}{2-\theta}Q^{n+1} + \frac{1-\theta}{2-\theta}Q^{n+1-\theta}$ .  
always useful

**thetsn**      ra       $0 \leq \text{real} \leq 1$       [0, 0.5 or 1]      O      L3  
`thetsn` is the value of  $\theta$  used to extrapolate the non linear explicit source terms  $S_e$  of the momentum equation, when the source term extrapolation has been activated (see `isno2t`), following the formula  
 $(S_e)^{n+\theta} = (1+\theta)S_e^n - \theta S_e^{n-1}$   
the value of  $\theta=\text{thetsn}$  is deduced from the value chosen for `isno2t`. Generally, only the value 0.5 is used. The user is not allowed to modify this variable.  
= 0: first-order (unused, corresponds to `isno2t`=0)  
= 0.5: second-order (used when `isno2t`=1)  
= 1: first-order (used when `isno2t`=2)  
always useful

**thetst**      ra       $0 \leq \text{real} \leq 1$       [0, 0.5 or 1]      O      L3  
`thetst` is the value of  $\theta$  used to extrapolate the non linear explicit source terms  $S_e$  of the turbulence equations, when the source term extrapolation has been activated (see `isto2t`), following the formula  
 $(S_e)^{n+\theta} = (1+\theta)S_e^n - \theta S_e^{n-1}$   
the value of  $\theta=\text{thetsn}$  is deduced from the value chosen for `isto2t`. Generally, only the value 0.5 is used. The user is not allowed to modify this variable.  
= 0: first-order (unused, corresponds to `isto2t`=0)  
= 0.5: second-order (used when `isto2t`=1)

= 1: first-order (used when `isto2t=2`)  
always useful

**thetss**      ra       $0 \leq \text{real} \leq 1$       [0, 0.5 or 1]      O      L3  
for each scalar `iscal`, `thetss(iscal)` is the value of  $\theta$  used to extrapolate the non linear explicit source terms  $S_e$  of the scalar equation, when the source term extrapolation has been activated (see `isso2t`), following the formula  
 $(S_e)^{n+\theta} = (1 + \theta)S_e^n - \theta S_e^{n-1}$   
the value of  $\theta=\text{thetss(iscal)}$  is deduced from the value chosen for `isso2t(iscal)`. Generally, only the value 0.5 is used. The user is not allowed to modify this variable.  
= 0: first-order (unused, corresponds to `isso2t(iscal)=0`)  
= 0.5: second-order (used when `isso2t(iscal)=1`)  
= 1: first-order (used when `isso2t(iscal)=2`)  
useful if `nscal>1`

**thetro**      ra       $0 \leq \text{real} \leq 1$       [0, 0.5 or 1]      O      L3  
**thetro** is the value of  $\theta$  used to extrapolate the physical property  $\phi$  “density” when the extrapolation has been activated (see `iroext`), according to the formula  $\phi^{n+\theta} = (1 + \theta)\phi^n - \theta\phi^{n-1}$   
the value of  $\theta=\text{thetro}$  is deduced from the value chosen for `iroext`. Generally, only the value 0.5 is used. The user is not allowed to modify this variable.  
= 0: first-order (unused, corresponds to `iroext=0`)  
= 0.5: second-order (corresponds to `iroext=1`)  
= 1: first-order (corresponds to `iroext=2`)  
always useful

**thetvi**      ra       $0 \leq \text{real} \leq 1$       [0, 0.5 or 1]      O      L3  
**thetvi** is the value of  $\theta$  used to extrapolate the physical property  $\phi$  “total viscosity” when the extrapolation has been activated (see `iviext`), according to the formula  $\phi^{n+\theta} = (1 + \theta)\phi^n - \theta\phi^{n-1}$   
the value of  $\theta=\text{thetvi}$  is deduced from the value chosen for `iviext`. Generally, only the value 0.5 is used. The user is not allowed to modify this variable.  
= 0: first-order (unused, corresponds to `iviext=0`)  
= 0.5: second-order (corresponds to `iviext=1`)  
= 1: first-order (corresponds to `iviext=2`)  
always useful

**thetcp**      ra       $0 \leq \text{real} \leq 1$       [0, 0.5 or 1]      O      L3  
**thetcp** is the value of  $\theta$  used to extrapolate the physical property  $\phi$  “specific heat” when the extrapolation has been activated (see `icpext`), according to the formula  $\phi^{n+\theta} = (1 + \theta)\phi^n - \theta\phi^{n-1}$   
the value of  $\theta=\text{thetcp}$  is deduced from the value chosen for `icpext`. Generally, only the value 0.5 is used. The user is not allowed to modify this variable.  
= 0: first-order (unused, corresponds to `icpext=0`)  
= 0.5: second-order (corresponds to `icpext=1`)  
= 1: first-order (corresponds to `icpext=2`)  
always useful

**thetvs**      ra       $0 \leq \text{real} \leq 1$       [0, 0.5 or 1]      O      L3  
for each scalar `iscal`, `thetvs(iscal)` is the value of  $\theta$  used to extrapolate the physical property  $\phi$  “diffusivity” when the extrapolation has been activated (see `ivsext`), according to the formula  $\phi^{n+\theta} = (1 + \theta)\phi^n - \theta\phi^{n-1}$   
the value of  $\theta=\text{thetvs(iscal)}$  is deduced from the value chosen for `ivsext(iscal)`.



Generally, only the value 0.5 is used. The user is not allowed to modify this variable.  
= 0: first-order (unused, corresponds to `ivsext(iscal)=0`)  
= 0.5: second-order (corresponds to `ivsext(iscal)=1`)  
= 1: first-order (corresponds to `ivsext(iscal)=2`)  
useful if `nscal>1`

## 9.2.7 Gradient reconstruction

<b>imrgra</b>	i	0, 1, 2, 3, 4, 5, or 6	[0]	O	L2
indicates the type of gradient reconstruction (one method for all the variables) = 0: iterative reconstruction of the non-orthogonalities = 1: least squares method based on the first neighbour cells (cells which share a face with the treated cell) = 2: least squares method based on the extended neighborhood (cells which share a node with the treated cell) = 3: least squares method based on a partial extended neighbourhood (all first neighbours plus the extended neighbourhood cells that are connected to a face where the non-orthogonality angle is larger than parameter <b>anomax</b> ) = 4: iterative reconstruction with initialisation using the least squares method (first neighbours) = 5: iterative reconstruction with initialisation using the least squares method based on an extended neighbourhood = 6: iterative reconstruction with initialisation using the least squares method based on a partial extended neighbourhood if <b>imrgra</b> fails due to probable mesh quality problems, it is usually effective to use <b>imrgra=3</b> . Moreover, <b>imrgra=3</b> is usually faster than <b>imrgra=0</b> (but with less feedback on its use). It should be noted that <b>imrgra=1, 2 or 3</b> automatically triggers a gradient limitation procedure. See <b>imligr</b> . useful if and only if there is <b>n</b> so that <b>nswrgr(n) &gt; 1</b> . Also, pressure gradients (or other gradients deriving from a potential) always use an iterative reconstruction. To force a non-iterative reconstruction for those gradients, a negative value of this keyword may be used, in which case the method matching the absolute value of the keyword will be used.					
<b>nswrgr</b>	ia	positive integer	[100]	O	L3
for each unknown <b>ivar</b> , <b>nswrgr(ivar) ≤ 1</b> indicates that the gradients are not reconstructed if <b>imrgra = 0 or 4</b> , <b>nswrgr(ivar)</b> is the number of iterations for the gradient reconstruction if <b>imrgra = 1, 2 or 3</b> , <b>nswrgr(ivar) &gt; 1</b> indicates that the gradients are reconstructed (but the method is not iterative, so any value larger than 1 for <b>nswrgr</b> yields the same result) useful for all the unknowns					
<b>epsrgr</b>	ra	real number > 0	[10 <sup>-5</sup> ]	O	L3
for each unknown <b>ivar</b> , relative precision for the iterative gradient reconstruction: <b>epsrgr(ivar)</b> useful for all the unknowns when <b>imrgra = 0 or 4</b>					
<b>imligr</b>	ia	-1, 0 or 1	[-1 or 1]	O	L3
for each unknown <b>ivar</b> , indicates the type of gradient limitation: <b>imligr(ivar)</b>					



=-1: no limitation  
 = 0: based on the neighbors  
 = 1: superior order

for all the unknowns, **imligr** is initialised to -1 if **imrgra**=0 or 4 and to 1 if **imrgra** = 1, 2 or 3  
useful for all the unknowns

**climgr**      ra      real number > 0      [1.5]      O      L3  
 for each unknown **ivar**, factor of gradient limitation: **climgr(ivar)** (high value means little limitation)  
 useful for all the unknowns **ivar** for which **imligr(ivar) ≠ -1**

**extrag**      ra      0, 0.5 or 1      [0]      O      L3  
 for the variable “pressure” **ivar=ipr**, extrapolation coefficient of the gradients at the boundaries. It affects only the Neumann conditions. The only possible values of **extrag(ipr)** are:  
     = 0: homogeneous Neumann calculated at first-order  
     = 0.5: improved homogeneous Neumann, calculated at second-order in the case of an orthogonal mesh and at first-order otherwise  
     = 1: gradient extrapolation (gradient at the boundary face equal to the gradient in the neighbour cell), calculated at second-order in the case of an orthogonal mesh and at first-order otherwise  
**extrag** often allows to correct the non-physical velocities that appear on horizontal walls when density is variable and there is gravity. It is strongly advised to keep **extrag**=0 for the variables apart from pressure. See also **iphydr**.  
 In practice, only the values 0 and 1 are allowed. The value 0.5 is not allowed by default (but the lock can be overridden if necessary, contact the development team).  
 always useful

**anamax**      r       $0 \leq \text{real} \leq \pi/2$       [ $\pi/4$ ]      O      L3  
 limit non-orthogonality angle used to restrict the extended neighborhood for the gradient calculation with **imrgra**=3.  
**anamax**=0 will yield the same result as **imrgra**=2 (full extended neighbourhood).  
**anamax**= $\pi/2$  will yield the same result as **imrgra**=1 (first neighbours only)<sup>34</sup>  
 useful if and only if **imrgra**=3

## 9.2.8 Solution of the linear systems

**iresol**      ia      -1, **ipol**\*1000+j      [-1]      O      L3  
 for each unknown **ivar**, **iresol(ivar)** determines the method used for the solution of the linear system  
     = -1: automatically managed by the code (conjugate gradient for the pressure **ivar=ipr** or any variable which is not convected, Jacobi for the others. Diagonal preconditioning with conjugate gradient).  
     = **ipol**\*1000+j with j= 0: conjugate gradient  
     = **ipol**\*1000+j with j= 200: single-reduction conjugate gradient  
         j= 1: Jacobi  
         j= 2: stabilised bi-conjugate gradient (BI-CGSTAB)  
         j= 3: GMRES  
**ipol** is the degree of the Neumann polynomial used for the preconditioning<sup>35</sup>.

<sup>34</sup>except for pathological cases where the non-orthogonality angle of a face would be larger than  $\pi/2$

<sup>35</sup> $D$  being the diagonal part of  $A$  and  $X$  its extra-diagonal part, it can be written  $A = D(Id + D^{-1}X)$ . Therefore  $A^{-1} = (Id + D^{-1}X)^{-1}D^{-1}$ . A series development of  $Id + D^{-1}X$  can then be used which yields, symbolically,

`ipol = -1` may be used to deactivate preconditioning.

`ipol` is necessarily 0 with the Jacobi algorithm.

Concerning the computational time, the performance depends on the case. If a preconditioning method different from the diagonal preconditioning is to be used, it seems to be better to restrict to a first-order preconditioning (`ipol=1`). This preconditioning may slightly increase performance in some cases but may decrease it in others.  
always useful

<b>nitmax</b>	ia	integer > 0	[10000]	O	L3	for each unknown <b>ivar</b> , maximum number of iterations for the solution of the linear systems: <b>nitmax(ivar)</b> when the algebraic multi-grid option is activated for the variable <b>ivar</b> ( <b>imgr(ivar)=1</b> ), <b>nitmax(ivar)</b> is the maximum number of iterations for the solution on the coarsest mesh always useful
<b>epsilo</b>	ra	real number > 0	[10 <sup>-8</sup> ,10 <sup>-5</sup> ]	O	L3	for each unknown <b>ivar</b> , relative precision for the solution of the linear system. The default value is <b>epsilo(ivar)=10<sup>-8</sup></b> . This value is set low on purpose. When there are enough iterations on the reconstruction of the right-hand side of the equation, the value may be increased (by default, in case of second-order in time, with <b>nswrsm = 5</b> or 10, <b>epsilo</b> is increased to 10 <sup>-5</sup> ). always useful
<b>imgr</b>	ia	0 or 1	[0]	O	L3	for each unknown <b>ivar</b> , indicates the use ( <b>imgr(ivar)=1</b> ) or not (=0) of the algebraic multi-grid method for the solution of the linear systems <b>imgr(ivar)</b> can be set independently for every variable always useful. Generally, its use is designed for the variable “pressure” in case of meshes with strongly stretched cells. It is recommended not to modify <b>imgr</b>
<b>ncegrm</b>	i	integer > 0	[30]	O	L3	for the multi-grid method, maximum number of cells on the coarsest grid useful if and only if <b>imgr(ivar) = 1</b> for at least one variable <b>ivar</b>
<b>ncymax</b>	ia	integer > 0	[100]	O	L3	for each unknown <b>ivar</b> , <b>ncymax(ivar)</b> is the maximum number of cycles when using the multi-grid method. useful if and only if <b>imgr(ivar) = 1</b>
<b>ngrmax</b>	i	1 ≤ integer ≤ <b>ngrmmx</b>	[ <b>ngrmmx</b> ]	O	L3	when using the multi-grid method, maximum number of grid levels useful if and only if <b>imgr(ivar) = 1</b> for at least one variable <b>ivar</b>
<b>ncymax</b>	ia	integer > 0	[10]	O	L3	for each unknown <b>ivar</b> , <b>ncymax(ivar)</b> is the maximum number of multi-grid cycles. useful if and only if <b>imgr(ivar) = 1</b>
<b>nitmgf</b>	ia	integer > 0	[10]	O	L3	for each unknown <b>ivar</b> , <b>nitmgf(ivar)</b> is the maximum number of iterations on all

---


$$Id + \sum_{I=1}^{IPOL} (-D^{-1}X)^I.$$

grids except for the coarsest when the multi-grid method is used; the resolution on the coarsest grid uses `nitmax`.  
useful if and only if `imgr(ivar) = 1`

<b>rlxp1</b>	<code>r</code>	$0 \leq \text{real} \leq 1$	[0.95]	O	L3
--------------	----------------	-----------------------------	--------	---	----

relaxation parameter for the multi-grid.

#### WARNING

The algebraic multi-grid method has only been tested for the “pressure” variable (`imgr(ipr)=1`).

### 9.2.9 Convective scheme

<b>blencv</b>	<code>ra</code>	$0 \leq \text{real} \leq 1$	[0 or 1]	O	L1
---------------	-----------------	-----------------------------	----------	---	----

for each unknown `ivar` to calculate, `blencv(ivar)` indicates the proportion of second-order convective scheme (0 corresponds to an “upwind” first-order scheme) ; in case of LES calculation, a second-order scheme is recommended and activated by default (`blencv=1`)  
useful for all the unknowns `ivar` for which `iconv(ivar) = 1`

<b>ischcv</b>	<code>ia</code>	0 or 1	[1]	O	L2
---------------	-----------------	--------	-----	---	----

for each unknown `ivar` to calculate, `ischcv(ivar)` indicates the type of second-order convective scheme  
= 0: Second Order Linear Upwind  
= 1: Centered  
useful for all the unknowns `ivar` which are convected (`iconv(ivar)=1`) and for which a second-order scheme is used (`blencv(ivar) > 0`)

<b>isstpc</b>	<code>ia</code>	0 or 1	[0]	O	L2
---------------	-----------------	--------	-----	---	----

for each unknown `ivar` to calculate, `isstpc(ivar)` indicates whether a “slope test” should be used to switch from a second-order to an “upwind” convective scheme under certain conditions, to ensure stability.  
= 0: “slope test” activated for the considered unknown  
= 1: “slope test” deactivated for the considered unknown  
useful for all the unknowns `ivar` which are convected (`iconv(ivar)=1`) and for which a second-order scheme is used (`blencv(ivar) > 0`).  
the use of the “slope test” stabilises the calculation but may bring the order in space to decrease quickly.

### 9.2.10 Pressure-continuity step

<b>iprco</b>	<code>i</code>	0 or 1	[1]	O	L3
--------------	----------------	--------	-----	---	----

indicates if the pressure-continuity step is taken into account (1) or not (0)  
always useful

<b>arak</b>	<code>ra</code>	$0 < \text{real} \leq 1$	[1]	O	L3
-------------	-----------------	--------------------------	-----	---	----

**arak** is the Arakawa coefficient before the Rhie& Chow filter  
always useful

<b>irevmc</b>	<code>ia</code>	0, 1 or 2	[0]	O	L3
---------------	-----------------	-----------	-----	---	----

method used to update the velocity after the pressure correction:

EDF R&D	<b><i>Code_Saturne</i> version 3.2.0 practical user's guide</b>	<i>Code_Saturne</i> documentation Page 165/ <a href="#">202</a>
---------	---	---

- standard gradient of pressure increment (**irevmc**=0)
- least squares on the pressure increment (**irevmc**=1)
- “rt0” *i.e.* least squares on the updated mass flux (**irevmc**=2)

**irevmc**=1 or 2 are only available when using the segregated algorithm for velocity components coupling (**ivelco** = 0).

the method **irevmc**=2 is generally not recommended.

always useful

**iphydr**      i      0 or 1      [0]      O      L2

method for taking into account the balance between the pressure gradient and the source terms (gravity and head losses): by extension it will be referenced as “taking into account of the hydrostatic pressure”

= 0: standard algorithm

= 1: improved algorithm

always useful

When the density effects are important, the choice of **iphydr**=1 allows to improve the interpolation of the pressure and correct the non-physical velocities which may appear in highly stratified areas or near horizontal walls (thus avoiding the use of **extrag** if the non-physical velocities are due only to gravity effects).

The improved algorithm also allows to eradicate the velocity oscillations which tend to appear at the frontiers of areas with high head losses.

In the case of a stratified flow, the calculation cost is higher when the improved algorithm is used (about 30% depending on the case) because the hydrostatic pressure must be recalculated at the outlet boundary conditions: see **icalhy**.

On meshes of insufficient quality, in order to improve the convergence, it may be useful to increase the number of iterations for the reconstruction of the pressure right-hand member, *i.e.* **nswrsm(ipr)**.

If head losses are present just along an outlet boundary, it is necessary to specify **icalhy**=0 in order to deactivate the recalculation of the hydrostatic pressure at the boundary, which may otherwise cause instabilities.

**icalhy**      i      0 or 1      [0 or 1]      O      L3

activates the calculation of hydrostatic pressure boundary conditions at outlet boundaries

= 0: no calculation of the hydrostatic pressure at the outlet boundary

= 1: calculation of the hydrostatic pressure at the outlet boundary

always useful

This option is automatically specified depending on the choice of **iphydr** and the value of gravity (**icalhy**=1 if **iphydr**=1 and gravity is different from 0; otherwise **icalhy**=0). The activation of this option generates an additional calculation cost (about 30% depending on the case).

If head losses are present just along an outlet boundary, it is necessary to specify **icalhy**=0 in order to deactivate the recalculation of the hydrostatic pressure at the boundary, which may otherwise cause instabilities

**epsdp**      r      real number > 0      [10<sup>-12</sup>]      O      L3

Parameter of diagonal pressure strengthening

**iporos**      i      0 or 1      [0]      O      L2

indicates if the porosity formulation is taken into account

= 0: standard algorithm (without porosity)

= 1: porosity taken into account

useful when head losses are taken into account

### 9.2.11 Error estimators for Navier-Stokes

There are currently `nestmx=4` types of local estimators provided at every time step, with two possible definitions for each<sup>36</sup>. These scalars indicate the areas (cells) in which some error types may be important. They are stored in the array `propce` containing the properties at the cells (see `iestim`). For each estimator, the code writes the minimum and maximum values in the listing and generates post-processing outputs along with the other variables.

The additional memory cost is about one real number per cell and per estimator. The additional calculation cost is variable. For instance, on a simple test case, the total estimator `iestot` generates an additional cost of 15 to 20 % on the CPU time<sup>37</sup>; the cost of the three others may be neglected. If the user wants to avoid the calculation of the estimators during the computation, it is possible to run a calculation without estimators first, and then activate them on a restart of one or two time steps.

It is recommended to use the estimators only for visual and qualitative analysis. Also, their use is compatible neither with a second-order time scheme nor with a calculation with a frozen velocity field.

**iest = iespre: prediction** (default name: EsPre). After the velocity prediction step (yielding  $\underline{u}^*$ ), the estimator  $\eta_{i,k}^{pred}(\underline{u}^*)$ , local variable calculated at every cell  $\Omega_i$ , is created from  $\underline{\mathcal{R}}^{pred}(\underline{u}^*)$ , which represents the residual of the equation solved during this step:

$$\begin{aligned}\underline{\mathcal{R}}^{pred}(\underline{u}^*) &= \rho^n \frac{\underline{u}^* - \underline{u}^n}{\Delta t} + \rho^n \underline{u}^n \cdot \underline{\text{grad}}(\underline{u}^*) - \text{div} \left( (\mu + \mu_t)^n \underline{\text{grad}}(\underline{u}^*) \right) + \nabla(P^n) \\ &- \text{rest of the right-hand member } (\underline{u}^n, P^n, \text{other variables}^n)\end{aligned}$$

By definition:

$$\eta_{i,k}^{pred}(\underline{u}^*) = |\Omega_i|^{(k-2)/2} \|\underline{\mathcal{R}}^{pred}(\underline{u}^*)\|_{\mathcal{L}^2(\Omega_i)}$$

- The first family,  $k = 1$ , suppresses the volume  $|\Omega_i|$  which intrinsically appears with the norm  $\mathcal{L}^2(\Omega_i)$ .

- The second family,  $k = 2$ , exactly represents the norm  $\mathcal{L}^2(\Omega_i)$ . The size of the cell therefore appears in its calculation and induces a weighting effect.

$\eta_{i,k}^{pred}(\underline{u}^*)$  is ideally equal to zero when the reconstruction methods are perfect and the associated system is solved exactly.

**iest = iesder: drift** (default name: EsDer). The estimator  $\eta_{i,k}^{der}(\underline{u}^{n+1})$  is based on the following quantity (intrinsic to the code):

$$\begin{aligned}\eta_{i,k}^{der}(\underline{u}^{n+1}) &= |\Omega_i|^{(k-2)/2} \|\text{div}(\text{corrected mass flow after the pressure step}) - \Gamma\|_{\mathcal{L}^2(\Omega_i)} \\ &= |\Omega_i|^{(1-k)/2} |\text{div}(\text{corrected mass flow after the pressure step}) - \Gamma|\end{aligned}\quad (8)$$

Ideally, it is equal to zero when the Poisson equation related to the pressure is solved exactly.

**iest = iescor: correction** (default name: EsCor). The estimator  $\eta_{i,k}^{corr}(\underline{u}^{n+1})$  comes directly from the mass flow calculated with the updated velocity field:

$$\eta_{i,k}^{corr}(\underline{u}^{n+1}) = |\Omega_i|^{\delta_{2,k}} |\text{div}(\rho^n \underline{u}^{n+1}) - \Gamma|$$

The velocities  $\underline{u}^{n+1}$  are taken at the cell centers, the divergence is calculated after projection on the faces.

$\delta_{2,k}$  represents the Kronecker symbol.

- The first family,  $k = 1$ , is the absolute raw value of the divergence of the mass flow minus the mass source term.

- The second family,  $k = 2$ , represents a physical property and allows to evaluate the difference in  $\text{kg.s}^{-1}$ .

Ideally, it is equal to zero when the Poisson equation is solved exactly and the projection from the mass flux at the faces to the velocity at the cell centers is made in a set of functions with null divergence.

<sup>36</sup>choice made by the user

<sup>37</sup>indeed, all the first-order in space differential terms have to be recalculated at the time  $t^{n+1}$

**iest = iestot: total** (default name: EsTot). The estimator  $\eta_{i,k}^{tot}(\underline{u}^{n+1})$ , local variable calculated at every cell  $\Omega_i$ , is based on the quantity  $\underline{\mathcal{R}}^{tot}(\underline{u}^{n+1})$ , which represents the residual of the equation using the updated values of  $\underline{u}$  and  $P$ :

$$\begin{aligned} \underline{\mathcal{R}}^{tot}(\underline{u}^{n+1}) &= \rho^n \frac{\underline{u}^{n+1} - \underline{u}^n}{\Delta t} + \rho^n \underline{u}^{n+1} \cdot \underline{\underline{grad}}(\underline{u}^{n+1}) - \text{div} \left( (\mu + \mu_t)^n \underline{\underline{grad}}(\underline{u}^{n+1}) \right) + \nabla(P^{n+1}) \\ &- \text{rest of the right-hand member } (\underline{u}^{n+1}, P^{n+1}, \text{other variables}^n) \end{aligned}$$

By definition:

$$\eta_{i,k}^{tot}(\underline{u}^{n+1}) = |\Omega_i|^{(k-2)/2} \|\underline{\mathcal{R}}^{tot}(\underline{u}^{n+1})\|_{\mathbb{L}^2(\Omega_i)}$$

The mass flux in the convective term is recalculated from  $\underline{u}^{n+1}$  expressed at the cell centres (and not taken from the updated mass flow at the faces).

As for the prediction estimator:

- The first family,  $k = 1$ , suppresses the volume  $|\Omega_i|$  which intrinsically appears with the norm  $\mathbb{L}^2(\Omega_i)$ .
- The second family,  $k = 2$ , exactly represents the norm  $\mathbb{L}^2(\Omega_i)$ . The size of the cell therefore appears in its calculation and induces a weighting effect.

The estimators are evaluated depending on the values of **iescal**.

<b>iescal</b>	ia	0, 1 or 2	[0]	O	L1
---------------	----	-----------	-----	---	----

**iescal**(**iest**) indicates the calculation mode for the error estimator **iest** (**iespre**, **iesder**, **iescor** or **iestot**), for the Navier-Stokes equation:  
**iescal** = 0: estimator not calculated,  
**iescal** = 1: the estimator  $\eta_{i,1}^*$  is calculated, without contribution of the volume,  
**iescal** = 2: the estimator  $\eta_{i,2}^*$  is calculated, with contribution of the volume ("norm  $L^2$ "), except for **iescor**, for which  $|\Omega_i| \eta_{i,1}^{corr}$  is calculated.

The name of the estimators appearing in the listing and the post-processing is made up of the default name (given before), followed by the value of **iescal**. For instance, EsPre2 is the estimator **iespre** calculated with **iescal**=2.  
always useful

## 9.2.12 Calculation of the distance to the wall

<b>icdpar</b>	i	-1, 1, -2 or 2	[-1]	O	L2
---------------	---	----------------	------	---	----

specifies the method used to calculate the distance to the wall  $y$  and the non-dimensional distance  $y^+$  for all the cells of the calculation domain (when necessary):  
= 1: standard algorithm (based on a Poisson equation for  $y$  and convection equation for  $y^+$ ), with reading of the distance to the wall from the restart file if possible  
=-1: standard algorithm (based on a Poisson equation for  $y$  and convection equation for  $y^+$ ), with systematic recalculation of the distance to the wall in case of calculation restart  
= 2: former algorithm (based on geometrical considerations), with reading of the distance to the wall from the restart file if possible  
=-2: former algorithm (based on geometrical considerations) with systematic recalculation of the distance to the wall in case of calculation restart  
In case of restart calculation, if the position of the walls haven't changed, reading the distance to the wall from the restart file can save a fair amount of CPU time.  
Useful in  $R_{ij} - \varepsilon$  model with wall echo (**iturb**=30 and **irijec**=1), in LES with van Driest damping (**iturb**=40 and **idries**=1) and in  $k - \omega$  SST (**iturb**=60).

By default, `icdpar` is initialised to -1, in case there has been a change in the definition of the boundary conditions between two computations (change in the number or the positions of the walls). Yet, with the  $k - \omega$  SST model, the distance to the wall is needed to calculate the turbulent viscosity, which is done before the calculation of the distance to the wall. Hence, when this model is used (and only in that case), `icdpar` is set to 1 by default, to ensure total continuity of the calculation at restart.

**As a consequence, with the  $k - \omega$  SST model, if the number and positions of the walls are changed at a calculation restart, it is mandatory for the user to set `icdpar` explicitly to -1**, otherwise the distance to the wall used will not correspond to the actual position of the walls.

The former algorithm is not compatible with parallelism nor periodicity. Also, whatever the value chosen for `icdpar`, the calculation of the distance to the wall is made at the most once for all at the beginning of the calculation. It is therefore not compatible with moving walls. Please contact the development team if you need to override this limitation.

The following options are related to `icdpar`=1 or -1. The options of level 2 are described first. Some options are used only in the case of the calculation of the non-dimensional distance to the wall  $y^+$  (LES model with van Driest damping). Most of these key words are simple copies of the key words for the numerical options of the general equations, with a potentially specific value in the case of the calculation of the distance to the wall.

<code>iwarny</code>	i integer	[0]	O	L2	specifies the level of the output writing concerning the calculation of the distance to the wall with <code>icdpar</code> =1 or -1. The higher the value, the more detailed the outputs useful when <code>icdpar</code> =1 or -1
<code>ntcmxy</code>	i positive integer	[1000]	O	L2	number of pseudo-time iterations for the calculation of the non-dimensional distance to the wall $y^+$ useful when <code>icdpar</code> =1 or -1 for the calculation of $y^+$
<code>nitmay</code>	i integer > 0	[10000]	O	L3	maximum number of iterations for the solution of the linear systems useful when <code>icdpar</code> =1 or -1
<code>nswrsy</code>	i positive integer	[1]	O	L3	number of iterations for the reconstruction of the right-hand members: corresponds to <code>nswrsm</code> useful when <code>icdpar</code> =1 or -1
<code>nswrgy</code>	i positive integer	[100]	O	L3	number of iterations for the gradient reconstruction: corresponds to <code>nswrgr</code> useful when <code>icdpar</code> =1 or -1
<code>imligy</code>	i -1, 0 or 1	[-1 or 1]	O	L3	type of gradient limitation: corresponds to <code>imligr</code> useful when <code>icdpar</code> =1 or -1
<code>ircfly</code>	i 0 or 1	[1]	O	L3	indicates the reconstruction of the convective and diffusive fluxes at the faces: corresponds to <code>ircflu</code> useful when <code>icdpar</code> =1 or -1

EDF R&D	<i>Code_Saturne</i> version 3.2.0 practical user's guide			<i>Code_Saturne</i> documentation Page 169/202
<b>ischcy</b>	i	0 or 1	[1]	O L3
type of second-order convective scheme: corresponds to <b>ischcv</b> useful when <b>icdpar</b> =1 or -1 for the calculation of $y^+$				
<b>isstpy</b>	i	0 or 1	[0]	O L3
indicates if a “slope test” should be used for a second-order convective scheme: corresponds to <b>isstpc</b> useful when <b>icdpar</b> =1 or -1 for the calculation of $y^+$				
<b>imgrpy</b>	i	0 or 1	[0]	O L3
indicates whether the algebraic multi-grid method should be used ( <b>imgr</b> ( <b>ivar</b> )=1) or not (0): corresponds to <b>imgr</b> useful when <b>icdpar</b> =1 or -1				
<b>blency</b>	r	$0 \leq \text{real} \leq 1$	[0]	O L3
proportion of second-order convective scheme: corresponds to <b>blencv</b> useful when <b>icdpar</b> =1 or -1 for the calculation of $y^+$				
<b>epsily</b>	r	real number $> 0$	$[10^{-8}]$	O L3
relative precision for the solution of the linear systems: corresponds to <b>epsilo</b> useful when <b>icdpar</b> =1 or -1				
<b>epsrgy</b>	r	real number $> 0$	$[10^{-5}]$	O L3
relative precision for the iterative gradient reconstruction: corresponds to <b>epsrgr</b> useful when <b>icdpar</b> =1 or -1				
<b>epsrsy</b>	r	real number $> 0$	$[10^{-5}]$	O L3
relative precision for the right hand side reconstruction: corresponds to <b>epsrsm</b> useful when <b>icdpar</b> =1 or -1				
<b>climgy</b>	r	real number $> 0$	[1.5]	O L3
limitation factor of the gradients: corresponds to <b>climgr</b> useful when <b>icdpar</b> =1 or -1				
<b>extray</b>	r	0, 0.5 or 1	[0]	O L3
extrapolation coefficient of the gradients at the boundaries: corresponds to <b>extrag</b> useful when <b>icdpar</b> =1 or -1				
<b>counxy</b>	r	strictly positive real number	[5000]	O L3
Target Courant number for the calculation of the non-dimensional distance to the wall useful when <b>icdpar</b> =1 or -1 for the calculation of $y^+$				
<b>epscvy</b>	r	strictly positive real number	$[10^{-8}]$	O L3
relative precision for the convergence of the pseudo-transient regime for the calculation of the non-dimensional distance to the wall useful when <b>icdpar</b> =1 or -1 for the calculation of $y^+$				
<b>yplmxy</b>	r	real number	[200]	O L3
value of the non-dimensional distance to the wall above which the calculation of the distance is not necessary (for the damping) useful when <b>icdpar</b> =1 or -1 for the calculation of $y^+$				



### 9.2.13 Others

<b>iccvfg</b>	i	0 or 1	[0]	O	L1	<p>indicates whether the dynamic field should be frozen (1) or not (0)  in such a case, the values of velocity, pressure and the variables related to the potential turbulence model (<math>k</math>, <math>R_{ij}</math>, <math>\varepsilon</math>, <math>\varphi</math>, <math>\bar{f}</math>, <math>\omega</math>, turbulent viscosity) are kept constant over time and only the equations for the scalars are solved  also, if <b>iccvfg</b>=1, the physical properties modified in <b>usphyv</b> will keep being updated. Beware of non-consistencies if these properties would normally affect the dynamic field (modification of density for instance)  useful if and only if <b>nscal</b> &gt; 0 and the calculation is a restart</p>
<b>ivelco</b>	i	0 or 1	[1]	O	L1	<p>indicates the algorithm for the velocity components coupling  = 0: segregated (deprecated)  = 1: coupled (default)  always useful</p>
<b>ipucou</b>	i	0 or 1	[0]	O	L1	<p>indicates the algorithm for velocity/pressure coupling  = 0: standard algorithm  = 1: reinforced coupling in case calculation with long time steps  always useful (it is seldom advised, but it can prove very useful, for instance, in case of flows with weak convection effects and highly variable viscosity)</p>
<b>nterup</b>	i	real number > 0	[1]	O	L2	<p>number of iterations on the pressure-velocity coupling on Navier-Stokes (for the PISO algorithm). useful for unsteady algorithm</p>
<b>epsup</b>	r	real number > 0	[10 <sup>-5</sup> ]	O	L2	<p>relative precision for the convergence test of the iterative process on pressure-velocity coupling (PISO). useful for unsteady algorithm</p>
<b>isuit1</b>	i	0 or 1	[0]	O	L1	<p>for the 1D wall thermal module, activation (1) or not(0) of the reading of the mesh and of the wall temperature from the <b>ficmt1</b> restart file  useful if <b>nfpt1d</b>&gt;0.</p>
<b>imvisf</b>	i	0 or 1	[0]	O	L3	<p>indicates the interpolation method used to project variables from the cell centers to the faces  = 0: linear  = 1: harmonic  always useful</p>
<b>ircflu</b>	ia	0 or 1	[1]	O	L2	<p>for each unknown <b>ivar</b>, <b>ircflu(ivar)</b> indicates whether the convective and diffusive fluxes at the faces should be reconstructed:  = 0: no reconstruction  = 1: reconstruction  deactivating the reconstruction of the fluxes can have a stabilising effect on the calculation. It is sometimes useful with the <math>k-\varepsilon</math> model, if the mesh is strongly non-orthogonal</p>

in the near-wall region, where the gradients of  $k$  and  $\varepsilon$  are strong. In such a case, setting `ircflu(ik)=0` and `ircflu(iep)=0` will probably help (switching to a first order convective scheme, `blencv=0`, for  $k$  and  $\varepsilon$  might also help in that case) always useful

<b>nswrsm</b>	ia	positive integer	[1, 2, 5 or 10]	O	L3
for each unknown <b>ivar</b> , <b>nswrsm(ivar)</b> indicates the number of iterations for the reconstruction of the right-hand members of the equations with a first-order scheme in time (standard case), the default values are 2 for pressure and 1 for the other variables. With a second-order scheme in time ( <b>ischtp=2</b> ) or LES, the default values are 5 for pressure and 10 for the other variables. useful for all the unknowns					
<b>epsrsm</b>	ra	real number > 0	[10 <sup>-8</sup> ,10 <sup>-5</sup> ]	O	L3
for each unknown <b>ivar</b> , relative precision on the reconstruction of the right hand-side. The default value is <b>epsrsm(ivar)=10<sup>-8</sup></b> . This value is set low on purpose. When there are enough iterations on the reconstruction of the right-hand side of the equation, the value may be increased (by default, in case of second-order in time, with <b>nswrsm</b> = 5 or 10, <b>epsrsm</b> is increased to 10 <sup>-5</sup> ). always useful					

## 9.3 Numerical, physical and modelling parameters

### 9.3.1 Numeric Parameters

These parameters correspond to numeric reference values in the code. They can be used but shall not be modified (they are defined as **parameter**).

<b>zero</b>	r	0	[0]	O	L3
Parameter containing the value 0					
<b>epzero</b>	r	10 <sup>-12</sup>	[10 <sup>-12</sup> ]	O	L3
“Small” real parameter, used for the comparisons of real numbers (absolute value of the difference lower than <b>epzero</b> )					
<b>pi</b>	r	3.141592653589793	[3.141592653589793]	O	L3
Parameter containing an approximate value of $\pi$					
<b>grand</b>	r	10 <sup>12</sup>	[10 <sup>12</sup> ]	O	L3
“Large” real parameter, generally used by default as a non physical value for the initialisations of variables which have to be modified by the user					
<b>rinfin</b>	r	10 <sup>30</sup>	[10 <sup>30</sup> ]	O	L3
Real parameter used to represent “infinity”					

### 9.3.2 Physical parameters

These parameters correspond to physical reference values in the code. They can be used but shall not be modified (they are defined as **parameter**).

EDF R&D	<b>Code_Saturne version 3.2.0 practical user's guide</b>		Code_Saturne documentation Page 172/202	
---------	--	--	---	--

<b>tkelvi</b>	r	273.15	[273.15]	O	L3
Temperature in Kelvin corresponding to 0 degrees Celsius.					
<b>tkelvn</b>	r	-273.15	[-273.15]	O	L3
Temperature in degrees Celsius corresponding to 0 Kelvin.					
<b>rr</b>	r	8.31434	[8.31434]	O	L3
Perfect gas constant in $J/mol/K$					
<b>treft</b>	r	25 + tkelvi	[25 + tkelvi]	O	L3
Reference temperature for the specific physics, in $K$					
<b>prefth</b>	r	101325	[101325]	O	L3
Reference pressure for the specific physics, in $Pa$					
<b>volmol</b>	r	22.41.10 <sup>-3</sup>	[22.41.10 <sup>-3</sup> ]	O	L3
Molar volume under normal pressure and temperature conditions (1 atmosphere, 0°C) in $m^{-3}$					
<b>stephn</b>	r	5.6703.10 <sup>-8</sup>	[5.6703.10 <sup>-8</sup> ]	O	L3
Stephan constant for the radiative module $\sigma$ in $W.m^{-2}.K^{-4}$					
<b>permvi</b>	r	1.2566.10 <sup>-6</sup>	[1.2566.10 <sup>-6</sup> ]	O	L3
Vacuum magnetic permeability $\mu_0$ ( $=4\pi.10^{-7}$ ) in $kg.m.A^{-2}.s^{-2}$					
<b>epszer</b>	r	8.854.10 <sup>-12</sup>	[8.854.10 <sup>-12</sup> ]	O	L3
Vacuum permittivity $\epsilon_0$ in $F.m^{-1}$					

### 9.3.3 Physical variables

<b>gx,gy,gz</b>	r	3 real numbers	[0,0,0]	O	L1
gravity components always useful					
<b>irovar</b>	ia	0 or 1	[-1]	C	L1
<b>irovar</b> =0 indicates that the density is constant. Its value is the reference density <b>ro0</b> . <b>irovar</b> =1 indicates that the density is variable: its variation law must be given in the user subroutine <b>usphyv</b> negative value: not initialised always useful					
<b>ivivar</b>	ia	0 or 1	[-1]	C	L1
<b>ivivar</b> =0 indicates that the molecular dynamic viscosity is constant. Its value is the reference molecular dynamic viscosity <b>visc10</b> . <b>ivivar</b> =1 indicates that the molecular dynamic viscosity is variable: its variation law must be given in the user subroutine <b>usphyv</b> negative value: not initialised always useful					

<b>ro0</b>	<p>ra real number <math>\geq 0</math> <span style="float: right;">[-grand*10]</span> <span style="float: right;">C L1</span></p> <p>ro0 is the reference density negative value: not initialised its value is not used in gas or coal combustion modelling (it will be calculated following the perfect gas law, with <math>P0</math> and <math>T0</math>). With the compressible module, it is also not used by the code, but it may be (and often is) referenced by the user in user subroutines; it is therefore better to specify its value. always useful otherwise, even if a law defining the density is given by the user subroutine <b>usphyv</b> or <b>uselph</b> indeed, except with the compressible module, <i>Code_Saturne</i> does not use the total pressure <math>P</math> when solving the Navier-Stokes equation, but a reduced pressure <math>P^* = P - \rho_0 g \cdot (\underline{x} - \underline{x}_0) + P_0^* - P_0</math> where <math>\underline{x}_0</math> is a reference point (see <b>xyzp0</b>) and <math>P_0^*</math> and <math>P_0</math> are reference values (see <b>pred0</b> and <b>p0</b>). Hence, the term <math>-\underline{\nabla}P + \rho g</math> in the equation is treated as <math>-\underline{\nabla}P^* + (\rho - \rho_0)g</math>. The closer <b>ro0</b> is to the value of <math>\rho</math>, the more <math>P^*</math> will tend to represent only the dynamic part of the pressure and the faster and more precise its solution will be. Whatever the value of <b>ro0</b>, both <math>P</math> and <math>P^*</math> appear in the listing and the post-processing outputs. with the compressible module, the calculation is made directly on the total pressure</p>
<b>viscl0</b>	<p>ra real number <math>&gt; 0</math> <span style="float: right;">[-grand*10]</span> <span style="float: right;">C L1</span></p> <p>viscl0 is the reference molecular dynamic viscosity negative value: not initialised always useful, it is the used value unless the user specifies the viscosity in the subroutine <b>usphyv</b></p>
<b>srrom</b>	<p>r <math>0 \leq \text{réel} &lt; 1</math> <span style="float: right;">[-grand or 0]</span> <span style="float: right;">c or O L1</span></p> <p>With gas combustion, pulverised coal or the electric module, <b>srrom</b> is the sub-relaxation coefficient for the density, following the formula: <math>\rho^{n+1} = \text{srrom} \rho^n + (1 - \text{srrom}) \rho^{n+1}</math> hence, with a zero value, there is no sub-relaxation. With combustion and pulverised coal, <b>srrom</b> is initialised to <b>-grand</b> and the user must specify a proper value through the Interface or the initialisation subroutines (<b>usd3p1</b>, <b>usebu1</b>, <b>uslwc1</b>, <b>uscpi1</b> or <b>uscpl1</b>). With the electric module, <b>srrom</b> is initialised in to 0 and may be modified by the user in <b>usel11</b>. With gas combustion, pulverised coal or electric arcs, <b>ssrom</b> is automatically used after the second time-step. With Joule effect, the user decides whether or not it will be used in <b>uselph</b> from the coding law giving the density.  always useful with gas combustion, pulverized coal or the electric module.</p>
<b>p0</b>	<p>ra real number <span style="float: right;">[1.013e - 5]</span> <span style="float: right;">O L1</span></p> <p>p0 is the reference pressure for the total pressure except with the compressible module, the total pressure <math>P</math> is evaluated from the reduced pressure <math>P^*</math> so that <math>P</math> is equal to p0 at the reference position <math>\underline{x}_0</math> (given by <b>xyzp0</b>) with the compressible module, the total pressure is solved directly always useful</p>
<b>pred0</b>	<p>ra real number <span style="float: right;">[0]</span> <span style="float: right;">O L3</span></p> <p>pred0 is the reference value for the reduced pressure <math>P^*</math> (see <b>ro0</b>) it is especially used to initialise the reduced pressure and as a reference value for the outlet boundary conditions for an optimised precision in the resolution of <math>P^*</math>, it is wiser to keep <b>pred0</b> to 0 with the compressible module, the “pressure” variable appearing in the equations</p>

EDF R&D	<b>Code_Saturne version 3.2.0 practical user's guide</b>	Code_Saturne documentation Page 174/202
---------	--	---

directly represents the total pressure. It is therefore initialised to **p0** and not **pred0** (see **ro0**)

always useful, except with the compressible module

<b>xyzp0</b>	<p>ra      3 real numbers      [0,0,0]      O      L1</p> <p><b>xyzp0(ii)</b> is the <b>ii</b> coordinate (<math>1 \leq ii \leq 3</math>) of the reference point <math>\underline{x}_0</math> for the total pressure when there are no Dirichlet conditions for the pressure (closed domain), <b>xyzp0</b> does not need to be specified (unless the total pressure has a clear physical meaning in the configuration treated)</p> <p>when Dirichlet conditions on the pressure are specified but only through standart outlet conditions (as it is in most configurations), <b>xyzp0</b> does not need to be specified by the user, since it will be set to the coordinates of the reference outlet face (<i>i.e.</i> the code will automatically select a reference outlet boundary face and set <b>xyzp0</b> so that <math>P</math> equals <b>p0</b> at this face). Nonetheless, if <b>xyzp0</b> is specified by the user, the calculation will remain correct</p> <p>when direct Dirichlet conditions are specified by the user (specific value set on specific boundary faces), it is better to specify the corresponding reference point (<i>i.e.</i> specify where the total pressure is <b>p0</b>). This way, the boundary conditions for the reduced pressure will be close to <b>pred0</b>, ensuring an optimal precision in the resolution. If <b>xyzp0</b> is not specified, the reduced pressure will be shifted, but the calculations will remain correct.</p> <p>with the compressible module, the “pressure” variable appearing in the equations directly represents the total pressure. <b>xyzp0</b> is therefore not used.</p> <p>always useful, except with the compressible module</p>
<b>t0</b>	<p>ra      real number      [0]      O      L1</p> <p><b>t0</b> is the reference temperature</p> <p>useful for the specific physics gas or coal combustion (initialisation of the density), for the electricity modules to initialise the domain temperature and for the compressible module (initialisations). It must be given in Kelvin.</p>
<b>cp0</b>	<p>ra      real number &gt; 0      [-grand*10]      O      L1</p> <p><b>cp0</b> is the reference specific heat</p> <p>useful if there is <math>1 \leq n \leq nscaus^{38}</math> so that <b>iscsth(n)</b>=1 (there is a scalar “temperature”), unless the user specifies the specific heat in the user subroutine <b>usphyv</b><sup>39</sup> (<b>icp</b> &gt; 0)</p> <p>with the compressible module or coal combustion, <b>cp0</b> is also needed even when there is no user scalar</p>
<b>icp</b>	<p>ia      0 or 1      [0]      O      L1</p> <p>indicates if the specific heat <math>C_p</math> is variable (<b>icp</b>=1) or not (0)</p> <p>When gas or coal combustion is activated, <b>icp</b> is automatically set to 0 (constant <math>C_p</math>). With the electric module, it is automatically set to 1. The user is not allowed to modify these default choices.</p> <p>When <b>icp</b>=1 is specified, the code automatically modifies this value to make <b>icp</b> designate the effective index-number of the property “specific heat”. For each cell <b>iel</b>, the value of <math>C_p</math> is then specified by the user in the appropriate subroutine (<b>usphyv</b> for the standard physics) and stored in the array <b>propce(iel,ipproc(icp))</b> (see p. 83 for specific conditions of use)</p> <p>useful if there is <math>1 \leq N \leq nscal</math> so that <b>iscsth(n)</b>=1 (there is a scalar “temperature”) or with the compressible module for non perfect gases</p>

<sup>38</sup>none of the scalars from the specific physics is a temperature

<sup>39</sup>when using the Graphical Interface, **cp0** is also used to calculate the diffusivity of the thermal scalars, based on their conductivity; it is therefore needed, unless the diffusivity is also specified in **usphyv**

<b>visls0</b>	<p>ra      real number &gt; 0      [-grand*10]      C      L1</p> <p><b>visls0(j)</b>: reference molecular diffusivity related to the scalar J (<math>kg.m^{-1}.s^{-1}</math>) negative value: not initialised useful if <math>1 \leq J \leq nscal</math>, unless the user specifies the molecular diffusivity in the appropriate user subroutine (<b>usphyv</b> for the standard physics) (<b>ivisls(iscal) &gt; 0</b>) <i>Warning: visls0 corresponds to the diffusivity. For the temperature, it is therefore defined as <math>\lambda/C_p</math> where <math>\lambda</math> and <math>C_p</math> are the conductivity and specific heat. When using the Graphical Interface, <math>\lambda</math> and <math>C_p</math> are specified separately, and visls0 is calculated automatically</i> <i>With the compressible module, visls0 (given in <b>uscfx2</b>) is directly the thermal conductivity <math>W.m^{-1}.K^{-1}</math></i> <i>With gas or coal combustion, the molecular diffusivity of the enthalpy (<math>kg.m^{-1}.s^{-1}</math>) must be specified by the user in the variable <b>diftl0</b> (<b>usebu1</b>, <b>usd3p1</b>, <b>uslwc1</b>, <b>uscpi1</b>, <b>uscpl1</b>)</i> <i>With the electric module, for the Joule effect, the diffusivity is specified by the user in <b>uselph</b> (even if it is constant). For the electric arcs, it is calculated from the thermochemical data file</i></p>
<b>ivisls</b>	<p>ia      positive or zero integer      [0]      O      L1</p> <p>indicates if the viscosity related to the scalar <b>iscal</b> is variable (<b>ivisls(iscal)=1</b>) or not (0). The user must specify <b>ivisls</b> only for the user scalars (<b>iscal <math>\leq</math> nscaus</b>). When <b>ivisls(iscal)=1</b> is specified, the code automatically modifies this value to make <b>ivisls(iscal)</b> designate the effective index-number of the property “diffusivity of the scalar <b>iscal</b>”. For each cell <b>iel</b>, the value is then specified by the user in the appropriate subroutine (<b>usphyv</b> for the standard physics) and stored in the array <b>propce(iel,ipproc(ivisls))</b> (see p. 83 for specific conditions of use) useful if <math>1 \leq n \leq nscal</math></p>
<b>diftl0</b>	<p>r      real number &gt; 0      [-grand]      C      L1</p> <p>molecular diffusivity for the enthalpy (<math>kg.m^{-1}.s^{-1}</math>) for gas or coal combustion (the code then automatically sets <b>visls0</b> to <b>diftl0</b> for the scalar representing the enthalpy) always useful for gas or coal combustion</p>
<b>scamin</b>	<p>ra      real number      [grand]      O      L1</p> <p><b>scamin(iscal)</b> is the lower limit value for the scalar <b>iscal</b>. At each time step, in every cell where the calculated value for <b>rtp(iel,isca(iscal))</b> is lower than <b>scamin(iscal)</b>, <b>rtp(iel,isca(iscal))</b> will be reset to <b>scamin(iscal)</b> there is no limitation if <b>scamin(iscal) &gt; scamax(iscal)</b> <b>scamin</b> shall not be specified for non-user scalars (specific physics) or for scalar variances useful if and only if <math>1 \leq iscal \leq nscaus</math></p>
<b>scamax</b>	<p>ra      real number      [-grand]      O      L1</p> <p><b>scamax(iscal)</b> is the higher limit value for the scalar <b>iscal</b>. At each time step, in every cell where the calculated value for <b>rtp(iel,isca(iscal))</b> is higher than <b>scamax(iscal)</b>, <b>rtp(iel,isca(iscal))</b> will be reset to <b>scamax(iscal)</b> there is no limitation if <b>scamin(iscal) &gt; scamax(iscal)</b> <b>scamax</b> shall not be specified for non-user scalars (specific physics) or for scalar variances useful if and only if <math>1 \leq iscal \leq nscaus</math></p>
<b>sigmas</b>	<p>ra      real number &gt; 0      [1]      O      L2</p>

**sigmas(iscal)**: turbulent Prandtl (or Schmidt) number for the scalar **iscal**  
useful if and only if  $1 \leq \text{iscal} \leq \text{nscals}$

**rvarfl**      ra      real number  $> 0$       [0.8]      O      L2  
when **iscavr(iscal)** $>0$ , **rvarfl(iscal)** is the coefficient  $R_f$  in the dissipation term  $-\frac{\rho}{R_f} \frac{\varepsilon}{k}$  of the equation concerning the scalar **iscal**, which represents the root mean square of the fluctuations of the scalar **iscavr(iscal)**  
useful if and only if there is  $1 \leq \text{iscal} \leq \text{nscals}$  such as **iscavr(iscal)** $>0$

### 9.3.4 modelling parameters

**xlomlg**      ra      real number  $> 0$       [-grand\*10]      O      L1  
**xlomlg** is the mixing length  
useful if and only if **iturb**= 10 (mixing length)

**almax**      ra      -grand, real number  $> 0$       [-grand\*10]      O      L2  
**almax** is a characteristic macroscopic length of the domain, used for the initialisation of the turbulence and the potential clipping (with **iclkep**=1)  
negative value: not initialised (the code then uses the cubic root of the domain volume)  
useful if and only if **turb**= 20, 21, 30, 31, 50 or 60 (RANS models)

**uref**      ra      real number  $> 0$       [-grand\*10]      C      L1  
**uref** is the characteristic flow velocity, used for the initialisation of the turbulence  
negative value: not initialised  
useful if and only if **iturb**= 20, 21, 30, 31, 50 or 60 (RANS model) and the turbulence is not initialised somewhere else (restart file or subroutine **cs\_user\_initialization**)

#### BASIC CONSTANTS OF THE $k - \varepsilon$ AND THE OTHER RANS MODELS

**xkappa**      r      real number  $> 0$       [0.42]      O      L3  
Kármán constant  
useful if and only if **iturb** $\geq 10$  (mixing length,  $k - \varepsilon$ ,  $R_{ij} - \varepsilon$ , LES, v2f or  $k - \omega$ )

**cstlog**      r      real number  $> 0$       [5.2]      O      L3  
constant of the logarithmic wall function  
useful if and only if **iturb** $\geq 10$  (mixing length,  $k - \varepsilon$ ,  $R_{ij} - \varepsilon$ , LES, v2f or  $k - \omega$ )

**cmu**      r      real number  $> 0$       [0.09]      O      L3  
constant  $C_\mu$  for all the RANS turbulence models except for the v2f model (see **cv2fmu** for the value of  $C_\mu$  in case of v2f modelling)  
useful if and only if **iturb**= 20, 21, 30, 31 or 60 ( $k - \varepsilon$ ,  $R_{ij} - \varepsilon$  or  $k - \omega$ )

**ce1**      r      real number  $> 0$       [1.44]      O      L3  
constant  $C_{\varepsilon 1}$  for all the RANS turbulence models except for the v2f and the  $k - \omega$  models  
useful if and only if **iturb**= 20, 21, 30 or 31 ( $k - \varepsilon$  or  $R_{ij} - \varepsilon$ )

**ce2**      r      real number  $> 0$       [1.92]      O      L3  
constant  $C_{\varepsilon 2}$  for the  $k - \varepsilon$  and  $R_{ij} - \varepsilon$  LRR models  
useful if and only if **iturb**= 20, 21 or 30 ( $k - \varepsilon$  or  $R_{ij} - \varepsilon$  LRR)

EDF R&D	<b>Code_Saturne version 3.2.0 practical user's guide</b>	Code_Saturne documentation Page 177/202
---------	--	---

**ce4**      r      real number  $> 0$       [1.2]      O      L3  
constant  $C_{\varepsilon 4}$  for the interfacial term (Lagrangian module) in case of two-way coupling  
useful in case of Lagrangian modelling, in  $k - \varepsilon$  and  $R_{ij} - \varepsilon$  with two-way coupling

**sigmak**      r      real number  $> 0$       [1.0]      O      L3  
Prandtl number for  $k$  with  $k - \varepsilon$  and v2f models  
useful if and only if **iturb**=20, 21 or 50 ( $k - \varepsilon$  or v2f)

**sigmae**      r      real number  $> 0$       [1.3]      O      L3  
Prandtl number for  $\varepsilon$   
useful if and only if **iturb**= 20, 21, 30, 31 or 50 ( $k - \varepsilon$ ,  $R_{ij} - \varepsilon$  or v2f)

#### CONSTANTS SPECIFIC TO THE $R_{ij} - \varepsilon$ LRR MODEL (**iturb**=30)

**crij1**      r      real number  $> 0$       [1.8]      O      L3  
constant  $C_1$  for the  $R_{ij} - \varepsilon$  LRR model  
useful if and only if **iturb**=30 ( $R_{ij} - \varepsilon$  LRR)

**crij2**      r      real number  $> 0$       [0.6]      O      L3  
constant  $C_2$  for the  $R_{ij} - \varepsilon$  LRR model  
useful if and only if **iturb**=30 ( $R_{ij} - \varepsilon$  LRR)

**crij3**      r      real number  $> 0$       [0.55]      O      L3  
constant  $C_3$  for the  $R_{ij} - \varepsilon$  LRR model  
useful if and only if **iturb**=30 ( $R_{ij} - \varepsilon$  LRR)

**csrij**      r      real number  $> 0$       [0.22]      O      L3  
constant  $C_s$  for the  $R_{ij} - \varepsilon$  LRR model  
useful if and only if **iturb**=30 ( $R_{ij} - \varepsilon$  LRR)

**crijp1**      r      real number  $> 0$       [0.5]      O      L3  
constant  $C'_1$  for the  $R_{ij} - \varepsilon$  LRR model, corresponding to the wall echo terms  
useful if and only if **iturb**=30 and **irijec**=1 ( $R_{ij} - \varepsilon$  LRR)

**crijp2**      r      real number  $> 0$       [0.3]      O      L3  
constant  $C'_2$  for the  $R_{ij} - \varepsilon$  LRR model, corresponding to the wall echo terms  
useful if and only if **iturb**=30 and **irijec**=1 ( $R_{ij} - \varepsilon$  LRR)

#### CONSTANTS SPECIFIC TO THE $R_{ij} - \varepsilon$ SSG MODEL

**cssgs1**      r      real number  $> 0$       [1.7]      O      L3  
constant  $C_{s1}$  for the  $R_{ij} - \varepsilon$  SSG model  
useful if and only if **iturb**=31 ( $R_{ij} - \varepsilon$  SSG)

**cssgs2**      r      real number  $> 0$       [-1.05]      O      L3  
constant  $C_{s2}$  for the  $R_{ij} - \varepsilon$  SSG model  
useful if and only if **iturb**=31 ( $R_{ij} - \varepsilon$  SSG)



EDF R&D	<b><i>Code_Saturne</i> version 3.2.0 practical user's guide</b>	<i>Code_Saturne</i> documentation Page 178/ <a href="#">202</a>
---------	---	---

<b>cssgr1</b>	r      real number $> 0$ constant $C_{r1}$ for the $R_{ij} - \varepsilon$ SSG model useful if and only if <b>iturb</b> =31 ( $R_{ij} - \varepsilon$ SSG)	[0.9]	O	L3
<b>cssgr2</b>	r      real number $> 0$ constant $C_{r2}$ for the $R_{ij} - \varepsilon$ SSG model useful if and only if <b>iturb</b> =31 ( $R_{ij} - \varepsilon$ SSG)	[0.8]	O	L3
<b>cssgr3</b>	r      real number $> 0$ constant $C_{r3}$ for the $R_{ij} - \varepsilon$ SSG model useful if and only if <b>iturb</b> =31 ( $R_{ij} - \varepsilon$ SSG)	[0.65]	O	L3
<b>cssgr4</b>	r      real number $> 0$ constant $C_{r4}$ for the $R_{ij} - \varepsilon$ SSG model useful if and only if <b>iturb</b> =31 ( $R_{ij} - \varepsilon$ SSG)	[0.625]	O	L3
<b>cssgr5</b>	r      real number $> 0$ constant $C_{r1}$ for the $R_{ij} - \varepsilon$ SSG model useful if and only if <b>iturb</b> =31 ( $R_{ij} - \varepsilon$ SSG)	[0.2]	O	L3
<b>cssge2</b>	r      real number $> 0$ constant $C_{\varepsilon 2}$ for the $R_{ij} - \varepsilon$ SSG model useful if and only if <b>iturb</b> =31 ( $R_{ij} - \varepsilon$ SSG)	[1.83]	O	L3

#### CONSTANTS SPECIFIC TO THE v2f $\varphi$ -MODEL

<b>cv2fa1</b>	r      real number $> 0$ constant $a_1$ for the v2f $\varphi$ -model useful if and only if <b>iturb</b> =50 (v2f $\varphi$ -model)	[0.05]	O	L3
<b>cv2fe2</b>	r      real number $> 0$ constant $C_{\varepsilon 2}$ for the v2f $\varphi$ -model useful if and only if <b>iturb</b> =50 (v2f $\varphi$ -model)	[1.85]	O	L3
<b>cv2fmu</b>	r      real number $> 0$ constant $C_\mu$ for the v2f $\varphi$ -model useful if and only if <b>iturb</b> =50 (v2f $\varphi$ -model)	[0.22]	O	L3
<b>cv2fc1</b>	r      real number $> 0$ constant $C_1$ for the v2f $\varphi$ -model useful if and only if <b>iturb</b> =50 (v2f $\varphi$ -model)	[1.4]	O	L3
<b>cv2fc2</b>	r      real number $> 0$ constant $C_2$ for the v2f $\varphi$ -model useful if and only if <b>iturb</b> =50 (v2f $\varphi$ -model)	[0.3]	O	L3
<b>cv2fct</b>	r      real number $> 0$ constant $C_T$ for the v2f $\varphi$ -model useful if and only if <b>iturb</b> =50 (v2f $\varphi$ -model)	[6]	O	L3

EDF R&D	<b><i>Code_Saturne</i> version 3.2.0 practical user's guide</b>	<i>Code_Saturne</i> documentation Page 179/202
---------	---	--

cv2fc1	r	real number $> 0$ constant $C_L$ for the v2f $\varphi$ -model useful if and only if <code>iturb=50</code> (v2f $\varphi$ -model)	[0.25]	O	L3
--------	---	--	--------	---	----

cv2fet	r	real number $> 0$ constant $C_\eta$ for the v2f $\varphi$ -model useful if and only if <code>iturb=50</code> (v2f $\varphi$ -model)	[110]	O	L3
--------	---	---	-------	---	----

#### CONSTANTS SPECIFIC TO THE $k - \omega$ SST MODEL

ckwsk1	r	real number $> 0$ constant $\sigma_{k1}$ for the $k - \omega$ SST model useful if and only if <code>iturb=60</code> ( $k - \omega$ SST)	[1/0.85]	O	L3
--------	---	---	----------	---	----

ckwsk2	r	real number $> 0$ constant $\sigma_{k2}$ for the $k - \omega$ SST model useful if and only if <code>iturb=60</code> ( $k - \omega$ SST)	[2]	O	L3
--------	---	---	-----	---	----

ckwsw1	r	real number $> 0$ constant $\sigma_{\omega 1}$ for the $k - \omega$ SST model useful if and only if <code>iturb=60</code> ( $k - \omega$ SST)	[2]	O	L3
--------	---	---	-----	---	----

ckwsw2	r	real number $> 0$ constant $\sigma_{\omega 2}$ for the $k - \omega$ SST model useful if and only if <code>iturb=60</code> ( $k - \omega$ SST)	[1/0.856]	O	L3
--------	---	---	-----------	---	----

ckwb1	r	real number $> 0$ constant $\beta_1$ for the $k - \omega$ SST model useful if and only if <code>iturb=60</code> ( $k - \omega$ SST)	[0.075]	O	L3
-------	---	---	---------	---	----

ckwb2	r	real number $> 0$ constant $\beta_2$ for the $k - \omega$ SST model useful if and only if <code>iturb=60</code> ( $k - \omega$ SST)	[0.0828]	O	L3
-------	---	---	----------	---	----

ckwgm1	r	real number $> 0$ constant $\gamma_1$ for the $k - \omega$ SST model useful if and only if <code>iturb=60</code> ( $k - \omega$ SST) <i>Warning: <math>\gamma_1</math> is calculated before the call to <code>usipsu</code>. Hence, if <math>\beta_1</math>, <math>C_\mu</math>, <math>\kappa</math> or <math>\sigma_{\omega 1}</math> is modified in <code>usipsu</code>, CKWGM1 must also be modified in accordance</i>	$[\frac{\beta_1}{C_\mu} - \frac{\kappa^2}{\sqrt{C_\mu \sigma_{\omega 1}}}]$	O	L3
--------	---	--	---	---	----

ckwgm2	r	real number $> 0$ constant $\gamma_2$ for the $k - \omega$ SST model useful if and only if <code>iturb=60</code> ( $k - \omega$ SST) <i>Warning: <math>\gamma_2</math> is calculated before the call to <code>usipsu</code>. Hence, if <math>\beta_2</math>, <math>C_\mu</math>, <math>\kappa</math> or <math>\sigma_{\omega 2}</math> is modified in <code>usipsu</code>, CKWGM2 must also be modified in accordance</i>	$[\frac{\beta_2}{C_\mu} - \frac{\kappa^2}{\sqrt{C_\mu \sigma_{\omega 2}}}]$	O	L3
--------	---	--	---	---	----

ckwa1	r	real number $> 0$ constant $a_1$ for the $k - \omega$ SST model useful if and only if <code>iturb=60</code> ( $k - \omega$ SST)	[0.31]	O	L3
-------	---	---	--------	---	----

EDF R&D	<b>Code_Saturne version 3.2.0 practical user's guide</b>	Code_Saturne documentation Page 180/202
---------	--	---

<b>ckwc1</b>	r	real number $> 0$ constant $c_1$ for the $k - \omega$ SST model useful if and only if <b>iturb</b> =60 ( $k - \omega$ SST)	[10]	O	L3
--------------	---	--	------	---	----

## 9.4 ALE

<b>iale</b>	i	0 or 1 activates (=1) or not (=0), activate the ALE module	[C]	O	L1
<b>nalinf</b>	i	0 or positive integer The number of sub-iterations of initialization of the fluid	[0]	C	L2
<b>nbstru</b>	i	0 or positive integer number of structures, automatically computed	[0]	C	L1
<b>alpnmk</b>	r	real <i>alpha</i> newmark's method	[0]	C	L3
<b>betnmk</b>	r	real <i>beta</i> newmark's method	[-grand]	C	L3
<b>gamnmk</b>	r	real <i>gamma</i> newmark's method	[-grand]	C	L3
<b>nalimx</b>	i	positive integer maximum number of implicitation iterations of the structure displacement	[15]	C	L2
<b>epalim</b>	r	positive real Relative precision of implicitation of the structure displacement	[1.10 <sup>-5</sup> ]	C	L2
<b>iflxmw</b>	i	0 or 1 method to compute interior mass flux due to ALE mesh velocity = 1: based on cell center mesh velocity = 0: based on nodes displacement	[1]	O	L2

## 9.5 Thermal radiative transfers: global settings

All the following key words may be modified in the user subroutines **usray\*** (or, for some of them, by through the thermochemical data files). It is however not recommended to modify those which do not belong to level L1.

<b>iirayo</b>	ia	0, 1, 2 <b>iirayo</b> activates ( $> 0$ ) or deactivates (=0) the radiation module The different values correspond to the following modellings: = 1: discrete ordinates (standard option for radiation in semi-transparent media) = 2: "P-1" model <i>Warning: the P-1 model allows faster computations, but it may only be applied to media with uniform large optical thickness, such as some cases of pulverised coal combustion</i>	[0]	O	L1
---------------	----	--	-----	---	----

<b>imodak</b>	i	0 or 1	[0]	O	L3	when gas or coal combustion is activated, <b>imodak</b> indicates whether the absorption coefficient shall be calculated “automatically” (=1) or read from the data file (=0) useful if the radiation module is activated
<b>isuidr</b>	i	0 or 1	[suite]	C	L1	indicates whether the radiation variables should be initialised (=0) or read from a restart file (=1) useful if and only if the radiation module is activated (in this case, a restart file <i>rayamo</i> must be available)
<b>nfreqr</b>	i	strictly positive integer	[1]	O	L1	period of the radiation module the radiation module is called every <b>nfreqr</b> time steps (more precisely, every time <b>ntcabs</b> is a multiple of <b>nfreqr</b> ). Also, in order to have proper initialisation of the variables, whatever the value of <b>nfreqr</b> , the radiation module is called at the first time step of a calculation (restart or not) useful if and only if the radiation module is activated
<b>ndirec</b>	i	32 or 128	[32]	O	L1	number of directions for the angular discretisation of the radiation propagation with the DOM model ( <b>iirayo</b> =1) no other possible value, because of the way the directions are calculated the calculation with 32 directions may break the symmetry of physically axi-symmetric cases (but the cost in CPU time is much lower than with 128 directions) useful if and only if the radiation module is activated with the DOM method
<b>xnp1mx</b>	r	real number	[10]	O	L3	with the P-1 model ( <b>iirayo</b> =2), <b>xnp1mx</b> is the percentage of cells of the calculation domain for which it is acceptable that the optical thickness is lower than unity <sup>40</sup> , although it is not to be desired useful if and only if the radiation module is activated with the P-1 method
<b>idiver</b>	i	0, 1 or 2	[2]	C	L1	indicates the method used to calculate the radiative source term: = 0: semi-analytic calculation (compulsory with transparent media) = 1: conservative calculation = 2: semi-analytic calculation corrected in order to be globally conservative useful if and only if the radiation module is activated <i>Note: if the medium is transparent, the choice has no effect on the calculation</i>
<b>iimpar</b>	i	0, 1 or 2	[1]	O	L1	choice of the display level in the listing concerning the calculation of the wall temperatures: = 0: no display = 1: standard = 2: complete useful if and only if the radiation module is activated
<b>iimlum</b>	i	0, 1 or 2	[1]	O	L1	choice of the display level in the listing concerning the solution of the radiative transfer

<sup>40</sup>more precisely, where  $KL$  is lower than 1, where  $K$  is the absorption coefficient of the medium and  $L$  is a characteristic length of the domain

equation:

- = 0: no display
- = 1: standard
- = 2: complete

useful if and only if the radiation module is activated

**nbrvaf**      ca      string of less than 80 characters      [name]      O      L1  
name associated for the post-processing to each of the following variables, defined at the boundary faces (*see* [6] for more details concerning their definitions):

- nbrvaf(itparp)**: wall temperature at the boundary faces ( $K$ )
- nbrvaf(iqincp)**: radiative incident flux density ( $W/m^2$ )
- nbrvaf(ixlamp)**: thermal conductivity of the boundary faces ( $W/m/K$ )
- nbrvaf(iepap)**: wall thickness ( $m$ )
- nbrvaf(iepsp)**: wall emissivity
- nbrvaf(ifnetp)**: net radiative flux density ( $W/m^2$ )
- nbrvaf(ifcomp)**: convective flux density ( $W/m^2$ )
- nbrvaf(ihcomp)**: convective exchange coefficient ( $W/m^2/K$ )

The default values are:

- nbrvaf(itparp)** = Wall\_temp
- nbrvaf(iqincp)** = Incident\_flux
- nbrvaf(ixlamp)** = Th\_conductivity
- nbrvaf(iepap)** = Thickness
- nbrvaf(iepsp)** = Emissivity
- nbrvaf(ifnetp)** = Net\_flux
- nbrvaf(ifcomp)** = Convective\_flux
- nbrvaf(ihcomp)** = Convective\_exch\_coef

useful if and only if the radiation module is activated

**irayvf**      ia      -1 or 1      [-1]      O      L1  
activates (=1) or deactivates (= -1) the post-processing for each of the following variables defined at the boundary faces:

- irayvf(itparp)**: wall temperature at the boundary faces ( $K$ )
- irayvf(iqincp)**: radiative incident flux density ( $W/m^2$ )
- irayvf(ixlamp)**: thermal conductivity of the boundary faces ( $W/m/K$ )
- irayvf(iepap)**: wall thickness ( $m$ )
- irayvf(iepsp)**: wall emissivity
- irayvf(ifnetp)**: net radiative flux density ( $W/m^2$ )
- irayvf(ifcomp)**: convective flux density ( $W/m^2$ )
- irayvf(ihcomp)**: convective exchange coefficient ( $W/m^2/K$ )

useful if and only if the radiation module is activated

**tmin**      r      real number positif      [0]      O      L3  
minimum allowed value for the wall temperatures in Kelvin  
useful if and only if the radiation module is activated

**tmax**      r      real number positif      [grand + 273.15]      O      L3  
maximum allowed value for the wall temperatures in Kelvin  
useful if and only if the radiation module is activated

## 9.6 Electric module (Joule effect and electric arcs): specificities

The electric module is composed of a Joule effect module (**ippmod(ieljou)**) and an electric arcs module (**ippmod(ielarc)**).

The Joule effect module is designed to take into account the Joule effect (for instance in glass furnaces) with real or complex potential in the enthalpy equation. The Laplace forces are not taken into account in the impulse momentum equation. Specific boundary conditions can be applied to account for the coupled effect of transformers (offset) in glass furnaces.

The electric arcs module is designed to take into account the Joule effect (only with real potential) in the enthalpy equation. The Laplace forces are taken into account in the impulse momentum equation.

The key words used in the global settings are quite few. They are found in the subroutine **usel11** (see the description of this user subroutine §8.8.5).

<b>ielcor</b>	i	0, 1	[0]	O	L1	when <b>ielcor</b> =1, the boundary conditions for the potential will be tuned at each time step in order to reach a user-specified target dissipated power <b>puisim</b> (Joule effect) or a user-specified target current intensity <b>couimp</b> (electric arcs) the boundary condition tuning is controlled by subroutines <b>elreca</b> or <b>uselrc</b> always useful
<b>couimp</b>	r	real number $\geq 0$	[0]	O	L1	with the electric arcs module, <b>couimp</b> is the target current intensity ( <i>A</i> ) for the calculations with boundary condition tuning for the potential the target intensity will be reached if the boundary conditions are expressed using the variable <b>dpot</b> or if the initial boundary conditions are multiplied by the variable <b>coejou</b> useful with the electric arcs module if <b>ielcor</b> =1
<b>puisim</b>	r	real number $\geq 0$	[0]	O	L1	with the Joule effect module, <b>puisim</b> is the target dissipated power ( <i>W</i> ) for the calculations with boundary condition tuning for the potential the target power will be reached if the boundary conditions are expressed using the variable <b>dpot</b> or if the initial boundary conditions are multiplied by the variable <b>coejou</b> useful with the Joule effect module if <b>ielcor</b> =1
<b>dpot</b>	r	real number $\geq 0$	[0]	O	L1	<b>dpot</b> is the potential difference ( <i>V</i> ) which generates the current (and the Joule effect) for the calculations with boundary conditions tuning for the potential. This value is initialised set by the user ( <b>usel11</b> ). It is then automatically tuned depending on the value of dissipated power (Joule effect module) or the intensity of current (electric arcs module). In order for the correct power or intensity to be reached, the boundary conditions for the potential must be expressed with <b>dpot</b> ( <b>uselc1</b> ). The tuning can be controlled in <b>uselrc</b> useful if <b>ielcor</b> =1
<b>modrec</b>	i	0, 1, 2	[1]	O	L1	when <b>ielcor</b> =0, we use user function <b>uselrc</b> for boundary condition tuning when <b>ielcor</b> =1, we use standard model for boundary condition tuning when <b>ielcor</b> =2, we use plane scaling model for boundary condition tuning. In this case, we need define plane and current density component used to
<b>idreca</b>	i	1, 2, 3	[3]	O	L1	define current density component used to calculate current in plane useful if <b>modrec</b> =2
<b>crit_reca</b>	r	real number $\geq 0$	[0]	O	L1	



always useful

The volume viscosity  $\kappa$  is defined by the formula expressing the stress:

$$\underline{\underline{\sigma}} = -P \underline{\underline{Id}} + \mu(\nabla \underline{u} + {}^t \underline{\underline{grad}} \underline{u}) + (\kappa - \frac{2}{3}\mu) \text{div}(\underline{u}) \underline{\underline{Id}} \quad (9)$$

<b>viscv0</b>	r	real number $\geq 0$	[0]	O	L1
viscv0 is the reference volume viscosity (noted $\kappa$ in the equation expressing $\underline{\underline{\sigma}}$ in the paragraph dedicated to <b>iviscv</b> )					
always useful, it is the used value, unless the user specifies the volume viscosity in the user subroutine <b>uscfpv</b>					
<b>igrdpp</b>	i	0 or 1	[1]	O	L3
indicates whether the pressure should be updated (=1) or not (=0) after the solution of the acoustic equation					
always useful					

## 9.8 Lagrangian multiphase flows

Most of these key words may be modified in the user subroutines **uslag1**, **uslag2**, **uslaen**, **uslast** and **uslaed**. It is however strongly recommended not to modify those belonging to the level L3.

First of all, it should be noted that the Lagrangian module is compliant with almost all the RANS turbulence models and with laminar flows. However, the standard particle turbulent dispersion model does not take full advantage of the second-order  $R_{ij} - \varepsilon$  models. The same isotropic model is used as in the  $k - \varepsilon$  models, with  $k$  calculated from the trace of  $R_{ij}$ . Also, two-way coupling is not compatible with the  $k - \omega$  SST model.

### 9.8.1 Global settings

<b>iilagr</b>	I	0, 1, 2, 3	[0]	C	L1
activates (>0) or deactivates (=0) the Lagrangian module					
the different values correspond to the following modellings:					
= 1 Lagrangian two-phase flow in one-way coupling (no influence of the particles on the continuous phase)					
= 2 Lagrangian two-phase flow with two-way coupling (influence of the particles on the dynamics of the continuous phase). Dynamics, temperature and mass may be coupled independently.					
= 3 Lagrangian two-phase flow on frozen continuous phase. This option can only be used in case of a calculation restart. All the Eulerian fields are frozen (including the scalar fields). This option automatically implies <b>iccvfg</b> = 1					
always useful					
<b>isuila</b>	i	0, 1	[0]	C	L1
activation (=1) or not (=0) of a Lagrangian calculation restart. The calculation restart file read when this option is activated ( <b>ficaml</b> ) only contains the data related to the particles (see also <b>isuist</b> )					
the global calculation must also be a restart calculation					
always useful					
<b>isuist</b>	i	0, 1	[0]	C	L1
during a Lagrangian calculation restart, indicates whether the particle statistics (volume and boundary) and two-way coupling terms are to be read from a restart file (=1)					



or reinitialised (=0). The file to be read is **ficmls**  
useful if **isuila** = 1

<b>nbpmax</b>	i	positive or null integer	[1000]	C	L1	maximum number of particles allowed simultaneously in the calculation domain. It must be reminded that the required memory evolves accordingly
<b>nbpart</b>	i	positive or null integer	[0]	O	L3	number of particles treated during one Lagrangian time step <b>nbpart</b> must always be lower than <b>nbpmax</b> always useful, but initialised and updated without intervention of the user
<b>nvls</b>	i	integer between 0 and 10	[0]	O	L2	number of additional variables related to the particles the additional variables can be accessed in the arrays <b>ettp</b> and <b>ettpa</b> by means of the pointer <b>jvls</b> : <b>ettp(nbpt,jvls(ii))</b> and <b>ettpa(nbpt,jvls(ii))</b> ( <b>nbpt</b> is the index-number of the treated particle, and <b>ii</b> an integer between 1 and <b>nvls</b> )
<b>isttio</b>	i	0, 1	[0]	C	L1	indicates the steady (=1) or unsteady (=0) state of the continuous phase flow in particular, <b>isttio</b> = 1 is needed in order to: calculate stationary statistics in the volume or at the boundaries (starting respectively from the Lagrangian iterations <b>nstist</b> and <b>nstbor</b> ) calculate time-averaged two-way coupling source terms (from the Lagrangian iteration <b>nstits</b> ) useful if <b>iilagr</b> =1 or <b>iilagr</b> =2 (if <b>iilagr</b> =3, then <b>isttio</b> =1 automatically)
<b>injcon</b>	i	0, 1	[0]	O	L1	activates (=1) or not (=0) the continuous injection of particles this option allows to inject particles continuously during the duration of the Lagrangian time step <b>dtp</b> rather than only once at the beginning of the Lagrangian iteration. It helps avoiding the fractioning of the particles cloud close to the injection areas
<b>iroule</b>	i	0, 1	[0]	O	L1	activates (=1) or not (=0) of the particle cloning/fusion technique (option also called "Russian roulette") when <b>iroule</b> = 1, the importance function must be specified <i>via</i> the array <b>croule</b> in the user subroutine <b>uslaru</b>
<b>ttclag</b>	r	positive real number	[0]	O	L3	physical time of the Lagrangian simulation always useful
<b>iplas</b>	i	integer > 0	[1]	O	L3	absolute iteration number (including the restarts) in the Lagrangian module ( <i>i.e.</i> Lagrangian time step number) always useful

## 9.8.2 Specific physics models associated with the particles

<b>iphyla</b>	i	0, 1, 2	[0]	C	L1	<p>activates (<math>&gt;0</math>) or deactivates (<math>=0</math>) the physical models associated to the particles:</p> <p>= 1: allows to associate with the particles evolution equations on their temperature (in degrees Celsius), their diameter and their mass</p> <p>= 2: the particles are pulverised coal particles. Evolution equations on temperature (in degree Celsius), mass of reactive coal, mass of char and diameter of the shrinking core are associated with the particles. This option is available only if the continuous phase represents a pulverised coal flame</p> <p>always useful</p>
<b>idpvar</b>	i	0, 1	[0]	O	L1	<p>activation (<math>=1</math>) or not (<math>=0</math>) of an evolution equation on the particle diameter</p> <p>useful if <b>iphyla</b> = 1</p>
<b>itpvar</b>	i	0, 1	[0]	O	L1	<p>activation (<math>=1</math>) or not (<math>=0</math>) of an evolution equation on the particle temperature (in degrees Celsius)</p> <p>useful if <b>iphyla</b> = 1 and if there is a thermal scalar associated with the continuous phase</p>
<b>impvar</b>	i	0, 1	[0]	O	L1	<p>activation (<math>=1</math>) or not (<math>=0</math>) of an evolution equation on the particle mass</p> <p>useful if <b>si iphyla</b> = 1</p>
<b>tpart</b>	r	real number $> \mathbf{tkelvn}$	[700]	O	L1	<p>initialisation temperature (in degree Celsius) for the particles already present in the calculation domain when an evolution equation on the particle temperature is activated during a calculation (<b>iphyla</b> = 1 and <b>itpvar</b> = 1)</p> <p>useful if <b>isuila</b> = 1 and <b>itpvar</b> = 0 in the previous calculation</p>
<b>cppart</b>	r	positive real number	[5200]	O	L1	<p>initialisation value for the specific heat (<math>J.kg^{-1}.K^{-1}</math>) of the particles already present in the calculation domain when an evolution equation on the particle temperature is activated during a calculation (<b>iphyla</b> = 1 and <b>itpvar</b> = 1)</p> <p>useful if <b>isuila</b> = 1 and <b>itpvar</b> = 0 in the previous calculation</p>
<b>iencra</b>	i	0, 1	[0]	O	L1	<p>activates (<math>=1</math>) or not (<math>=0</math>) the option of coal particle fouling. It then is necessary to specify the domain boundaries on which fouling may take place.</p> <p>useful if <b>iphyla</b> = 2</p>
<b>tprenc</b>	r	real number $> \mathbf{tkelvn}$	[600]	O	L1	<p>limit temperature (in degree Celsius) below which the coal particles do not cause any fouling (if the fouling model is activated)</p> <p>useful if <b>iphyla</b> = 2 and <b>iencra</b> = 1</p>
<b>visref</b>	r	positive real number	[10000]	O	L1	<p>ash critical viscosity in <math>kg.m^{-1}.s^{-1}</math>, in the fouling model <sup>41</sup></p> <p>useful if <b>iphyla</b> = 2 and <b>iencra</b> = 1</p>

<sup>41</sup>J.D. Watt et T. Fereday (*J.Inst.Fuel*, Vol.42-p99)

### 9.8.3 Options for two-way coupling

<b>nstits</b>	i	strictly positive integer	[1]	O	L1
		number of absolute Lagrangian iterations (including the restarts) after which a time-average of the two-way coupling source terms is calculated			
		indeed, if the flow is steady ( <b>isttio</b> =1), the average quantities that appear in the two-way coupling source terms can be calculated over different time steps, in order to get a better precision			
		if the number of absolute Lagrangian iterations is strictly inferior to <b>nstits</b> , the code considers that the flow has not yet reached its steady state (transition period) and the averages appearing in the source terms are reinitialised at each time step, as it is the case for unsteady flows ( <b>isttio</b> =0)			
		useful if <b>iilag</b> = 2 and <b>isttio</b> = 1			
<b>ltsdyn</b>	i	0, 1	[0]	O	L1
		activation (=1) or not (=0) of the two-way coupling on the dynamics of the continuous phase			
		useful if <b>iilag</b> = 2 and <b>iccvfg</b> = 0			
<b>ltsmas</b>	i	0, 1	[0]	O	L1
		activation (=1) or not (=0) of the two-way coupling on the mass			
		useful if <b>iilag</b> = 2, <b>iphyla</b> = 1 and <b>impvar</b> = 1			
<b>ltsthe</b>	i	0, 1	[0]	O	L1
		if <b>iphyla</b> = 1 and <b>itpvar</b> = 1, <b>ltsthe</b> activates (=1) or not (=0) the two-way coupling on temperature			
		if <b>iphyla</b> = 2, <b>ltsthe</b> activates (=1) or not (=0) the two-way coupling on the eulerian variables related to pulverised coal combustion			
		useful if <b>iilag</b> = 2			

### 9.8.4 Numerical modelling

<b>nordre</b>	i	1, 2	[2]	O	L2
		order of integration for the stochastic differential equations			
		= 1 integration using a first-order scheme			
		= 2 integration using a second-order scheme			
		always useful			
<b>ilapoi</b>	i	0, 1	[0]	O	L3
		activation (=1) or not (=0) of the solution of a Poisson's equation for the correction of the particle instantaneous velocities (in order to obtain a null divergence)			
		this option is not validated and reserved to the development team. Do not change the default value			
<b>idistu</b>	i	0, 1	[1]	O	L3
		activation (=1) or not (=0) of the particle turbulent dispersion			
		the turbulent dispersion is compatible only with the RANS turbulent models ( $k - \varepsilon$ , $R_{ij} - \varepsilon$ , v2f or $k - \omega$ )			
		( <b>iturb</b> =20, 21, 30, 31, 50 or 60)			
		always useful			

<b>idiff1</b>	i	0, 1	[0]	O	L3
<b>idiff1</b> =1 suppresses the crossing trajectory effect, making turbulent dispersion for the particles identical to the turbulent diffusion of fluid particles useful if <b>idistu</b> =1					
<b>modcpl</b>	i	positive integer	[0]	O	L1
activates (>0) or not (=0) the complete turbulent dispersion model when <b>modcpl</b> is strictly positive, its value is interpreted as the absolute Lagrangian time step number (including restarts) after which the complete model is applied since the complete model uses volume statistics, <b>modcpl</b> must either be 0 or be larger than <b>idstnt</b> useful if <b>istala</b> = 1					
<b>idirla</b>	i	1, 2, 3	[1]	O	L1
<i>x, y</i> or <i>z</i> direction of the complete model it corresponds to the main directions of the flow useful if <b>modcpl</b> > 0					

### 9.8.5 Volume statistics

<b>istala</b>	i	0, 1	[0]	C	L1
activation (=1) or not (=0) of the calculation of the volume statistics related to the dispersed phase if <b>istala</b> = 1, the calculation of the statistics is activated starting from the absolute iteration (including the restarts) <b>idstnt</b> by default, the statistics are not stationary (reset to zero at every Lagrangian iteration). But if <b>isttio</b> =1, since the flow is steady, the statistics will be averaged over the different time steps the statistics represent the significant results on the particle cloud always useful					
<b>seuil</b>	r	positive real number	[0]	O	L1
every cell of the calculation domain contains a certain quantity of particles, representing a certain statistical weight (sum of the statistical weights of all the particles present in the cell). <b>seuil</b> is the limit statistical weight value, below which the contribution of the cell in term of statistical weight is not taken into account in the volume statistics (for the complete turbulent dispersion model, in the Poisson's equation used to correct the mean velocities or in the listing and post-processing outputs) useful if <b>istala</b> = 1					
<b>idstnt</b>	i	strictly positive integer	[1]	C	L1
absolute Lagrangian iteration number (including the restarts) after which the calculation of the volume statistics is activated useful if <b>istala</b> = 1					
<b>nstist</b>	i	integer $\geq$ <b>idstnt</b>	[ <b>idstnt</b> ]	O	L1
absolute Lagrangian iteration number (including the restarts) after which the volume statistics are cumulated over time (they are then said to be stationary) if the absolute Lagrangian iteration number is lower than <b>nstist</b> , or if the flow is unsteady ( <b>isttio</b> =0), the statistics are reset to zero at every Lagrangian iteration (the volume statistics are then said to be non-stationary) useful if <b>istala</b> =1 and <b>isttio</b> =1					

<b>nomlag</b>	ca	string of less than 50 characters	[VarLagXXXX]	O	L1
		name of the volumetric statistics, displayed in the listing and the post-processing files. The default value is given above, with “XXXX” representing a four digit number (for instance 0001, 0011 ...) useful if <b>istala</b> = 1 <i>Warning: this name is also used to reference information in the restart file (<b>isuist</b> = 1). If the name of a variable is changed between two calculations, it will not be possible to read its value from the restart file</i>			
<b>nlvsts</b>	i	$0 \leq \text{integer} \leq \text{nussta}=20$	[0]	O	L1
		number of additional user volume statistics the additional statistics (or their cumulated value in the stationary case) can be accessed in the array <b>statis</b> by means of the pointer <b>ilvu</b> : <b>statis(iel,ilvu(ii))</b> ( <b>iel</b> is the cell index-number and <b>ii</b> an integer between 1 and <b>nlvsts</b> ) useful if <b>istala</b> = 1			
<b>npst</b>	i	positive integer	[0]	O	L3
		number of iterations during which stationary volume statistics have been cumulated useful if <b>istala</b> =1, <b>isttio</b> =1 and if <b>nstist</b> is inferior or equal to the current Lagrangian iteration <b>npst</b> is initialised and updated automatically by the code, its value is not to be modified by the user			
<b>npstt</b>	i	positive integer	[0]	O	L3
		number of iterations during which volume statistics have been calculated (the potential iterations during which non-stationary statistics have been calculated are counted in <b>npstt</b> ) useful if <b>istala</b> =1 <b>npstt</b> is initialised and updated automatically by the code, its value is not to be modified by the user			
<b>tstat</b>	r	positive real number	[dtp]	O	L3
		if the volume statistics are calculated in a stationary way, <b>tstat</b> represents the physical time during which the statistics have been cumulated if the volume statistics are calculated in a non-stationary way, then <b>tstat=dtp</b> (it is the Lagrangian time step, because the statistics are reset to zero at every iteration) useful if <b>istala</b> =1 <b>tstat</b> is initialised and updated automatically by the code, its value is not to be modified by the user			

## 9.8.6 Display of particles and trajectories

The definition of the particle visualization output meshes itself is done with that of all other postprocessing mesh zones, either using the GUI, or the `cs_user_postprocess_meshes` function of `cs_user_postprocess.c`.

The following options determine which variables are output on those particle or particle trajectory segment output meshes for which the “automatic output” is activated.

<b>ivisv1</b>	i	0, 1	[0]	O	L1
		associates (=1) or not (=0) the variable “velocity of the locally undisturbed fluid flow field” with the output of particles or trajectories always useful			

<b>ivisv2</b>	i      0, 1      [0]      O      L1 associates (=1) or not (=0) the variable “particle velocity” with the output of particles or trajectories always useful
<b>ivistp</b>	i      0, 1      [0]      O      L1 associates (=1) or not (=0) the variable “residence time” with the output of particles or trajectories always useful
<b>ivisdm</b>	i      0, 1      [0]      O      L1 associates (=1) or not (=0) the variable “particle diameter” with the output of particles or trajectories always useful
<b>iviste</b>	i      0, 1      [0]      O      L1 associates (=1) or not (=0) the variable “particle temperature” with the output of particles or trajectories always useful
<b>ivismp</b>	i      0, 1      [0]      O      L1 associates (=1) or not (=0) the variable “particle mass” with the output of particles or trajectories always useful
<b>ivisdk</b>	i      0, 1      [0]      O      L1 associates (=1) or not (=0) the variable “shrinking core diameter of the coal particles” with the output of particles or trajectories useful only if <b>iphyla</b> = 2
<b>ivisch</b>	i      0, 1      [0]      O      L1 associates (=1) or not (=0) the variable “mass of reactive coal of the coal particles” with the output of particles or trajectories useful only if <b>iphyla</b> = 2
<b>ivisck</b>	i      0, 1      [0]      O      L1 associates (=1) or not (=0) the variable “mass of coal of the coal particles” with the output of particles or trajectories useful only if <b>iphyla</b> = 2

### 9.8.7 Display of the particle/boundary interactions and the statistics at the boundaries

<b>iensi3</b>	i      0, 1      [0]      C      L1 activation (=1) or not (=0) of the recording of the particle/boundary interactions in <b>parbor</b> , and of the calculation of the statistics at the corresponding boundaries, for post-processing ( <i>EnSight6</i> format) By default, the statistics are non-stationary (reset to zero at every Lagrangian iteration). They may be stationary if <b>isttio</b> =1 ( <i>i.e.</i> calculation of a cumulated value over time, and then calculation of an average over time or over the number of interactions)
---------------	--

with the boundary)  
always useful

<b>nstbor</b>	i	strictly positive integer	[1]	O	L1	number of absolute Lagrangian iterations (including the restarts) after which the statistics at the boundaries are considered stationary and are averaged (over time or over the number of interactions) If the number of absolute Lagrangian iterations is lower than <b>nstbor</b> , or if <b>isttio</b> =0, the statistics are reset to zero at every Lagrangian iteration (non-stationary statistics) useful if <b>iensi3</b> =1 and <b>isttio</b> =1
<b>seuilf</b>	r	positive real number	[0]	O	L1	every boundary face of the mesh undergoes a certain number of interactions with particles, expressed in term of statistical weight (sum of the statistical weights of all the particles which have interacted with the boundary face). <b>seuilf</b> is the limit statistical weight value, below which the contribution of the face is not taken into account in the statistics at the boundaries for post-processing useful if <b>iensi3</b> =1
<b>inbrbd</b>	i	0, 1	[1]	O	L1	activation (=1) or not (=0) of the recording of the number of particle/boundary interactions, and of the calculation of the associated boundary statistics. <b>inbrd</b> = 1 is a compulsory condition to use the particulate average <b>imoybr</b> = 2 useful if <b>iensi3</b> =1
<b>iflmbd</b>	i	0, 1	[0]	O	L1	activation (=1) or not (=0) of the recording of the particulate mass flow related to the particle/boundary interactions, and of the calculation of the associated boundary statistics <b>inbrd</b> = 1 is a compulsory condition to use <b>iflmbd</b> =1 useful if <b>iensi3</b> =1 and <b>inbrbd</b> =1
<b>iangbd</b>	i	0, 1	[0]	O	L1	activation (=1) or not (=0) of the recording of the angle between a particle trajectory and a boundary face involved in a particle/boundary interaction, and of the calculation of the associated boundary statistics useful if <b>iensi3</b> =1
<b>ivitbd</b>	i	0, 1	[0]	O	L1	activation (=1) or not (=0) of the recording of the velocity of a particle involved in a particle/boundary interaction, and of the calculation of the associated boundary statistics useful if <b>iensi3</b> =1
<b>iencbd</b>	i	0, 1	[0]	O	L1	activation (=1) or not (=0) of the recording of the mass of coal particles stuck to the wall due to fouling, on the boundary faces of the <b>iencr1</b> interaction type useful if <b>iensi3</b> =1, <b>iphyla</b> =2, <b>iencra</b> =1, and if there is at least one boundary face of the <b>iencr1</b> interaction type
<b>nusbor</b>	i	positive integer	[0]	O	L1	number additional user data to record for the calculation of additional boundary statistics in <b>parbor</b> useful if <b>iensi3</b> =1

EDF R&D	<b>Code_Saturne version 3.2.0 practical user's guide</b>	Code_Saturne documentation Page 193/202
---------	--	---

<b>nombrd</b>	ca	string of less than 50 characters [see <b>uslag1</b> ]	O	L1	name of the boundary statistics, displayed in the listing and the post-processing files useful if <b>iensi3=1</b> <i>Warning: this name is also used to reference information in the restart file (<b>isuist</b> =1). If the name of a variable is changed between two calculations, it will not be possible to read its value from the restart file</i>
<b>imoybr</b>	ia	0, 1, 2 [0 , 1 or 2]	O	L1	the recordings in <b>parbor</b> at every particle/boundary interaction are cumulated values (possibly reset to zero at every iteration in the non-stationary case). They must therefore be divided by a quantity to get boundary statistics. The user can choose between two average types: = 0: no average is applied to the recorded cumulated values = 1: a time-average is calculated. The cumulated value is divided by the physical duration in the case of stationary averages ( <b>isttio=1</b> ). The cumulated value is divided by the value of the last time step in the case of non-stationary averages ( <b>isttio=0</b> ), and also in the case of stationary averages while the absolute Lagrangian iteration number is inferior to <b>nstbor</b> = 2: a particulate average is calculated. The cumulated value is divided by the number of particle/boundary interactions (in terms of statistical weight) recorded in <b>parbor(nfabor,inbr)</b> . This average can only be calculated when <b>inbrbd=1</b> . The average is calculated if the number of interactions (in statistical weight) of the considered boundary face is strictly higher than <b>seuilf</b> , otherwise the average at the face is set to zero only the cumulated value is recorded in the restart file useful if <b>iensi3=1</b>
<b>npstf</b>	i	positive integer [0]	O	L3	number of iterations during which stationary boundary statistics have been cumulated useful if <b>iensi3=1</b> , <b>isttio=1</b> and <b>nstbor</b> inferior or equal to the current Lagrangian iteration <b>npstf</b> is initialised and updated automatically by the code, its value is not to be modified by the user
<b>npstft</b>	i	positive integer [0]	O	L3	number of iterations during which boundary statistics have been calculated (the potential iterations during which non-stationary statistics have been calculated are counted in <b>npstft</b> ) useful if <b>iensi3=1</b> <b>npstft</b> is initialised and updated automatically by the code, its value is not to be modified by the user
<b>tstatp</b>	r	positive real number [dtp]	O	L3	if the recording of the boundary statistics is stationary, <b>tstatp</b> contains the cumulated physical duration of the recording of the boundary statistics if the recording of the boundary statistics is non-stationary, then <b>tstat=dtp</b> (it is the Lagrangian time step, because the statistics are reset to zero at every time step) useful if <b>iensi3=1</b>



## 10 Bibliography

- [1] F. ARCHAMBEAU, N. MÉCHITOUA, M. SAKIZ,  
*Code\_Saturne: a Finite Volume Code for the Computation of Turbulent Incompressible Flows*,  
Industrial Applications, International Journal on Finite Volumes, Vol. 1, 2004.
- [2] F. ARCHAMBEAU, *et al.*,  
*Note de validation de Code\_Saturne version 1.1.0*,  
EDF Report HI-83/04/003/A, 2004 (in french).
- [3] S. BENHAMADOUCHE,  
*Modélisation de sous-maille pour la LES - Validation avec la Turbulence Homogène Isotrope (THI)*  
*dans une version de développement de Code\_Saturne*,  
EDF Report HI-83/01/033/A, 2001 (in french).
- [4] M. BOUCKER, F. ARCHAMBEAU, N. MÉCHITOUA,  
*Quelques éléments concernant la structure informatique du Solveur Commun - Version 1.0\_init0*,  
Compte-rendu express EDF I81-00-8, 2000 (in french).
- [5] M. BOUCKER, J.D. MATTÉI,  
*Proposition de modification des conditions aux limites de paroi turbulente pour le Solveur Commun*  
*dans le cadre du modèle  $k - \varepsilon$  standard*,  
EDF Report HI-81/00/019/A, 2000 (in french).
- [6] A. DOUCE, N. MÉCHITOUA,  
*Mise en œuvre dans Code\_Saturne des physiques particulières. Tome3 : Transfert thermique radiatif*  
*en milieu gris semi-transparent*,  
EDF Report HI-81/02/019/A, 2002 (in french).
- [7] A. DOUCE,  
*Physiques particulières dans Code\_Saturne 1.1, Tome 5 : modélisation stochastique lagrangienne*  
*d'écoulements turbulents diphasiques polydispersés*,  
EDF Report, HI-81/04/03/A, 2005 (in french).
- [8] A. ESCAICH, P. PLION, *Mise en œuvre dans Code\_Saturne des modélisations physiques particulières.*  
*Tome 1 : Combustion en phase gaz*,  
EDF Report, HI-81/02/03/A, 2002 (in french).
- [9] A. ESCAICH, *Mise en œuvre dans Code\_Saturne des modélisations physiques particulières. Tome 2 :*  
*Combustion du charbon pulvérisé*,  
EDF Report, HI-81/02/09/A, 2002 (in french).
- [10] N. MÉCHITOUA, F. ARCHAMBEAU,  
*Prototype de solveur volumes finis co-localisé sur maillage non-structuré pour les équations de*  
*Navier-Stokes 3D incompressibles et dilatables avec turbulence et scalaire passif*,  
EDF Report HE-41/98/010/B, 1998 (in french).
- [11] Code\_Saturne DOCUMENTATION,  
*Code\_Saturne 3.2.0 Theory and Programmer's guide*,  
on line with the release of Code\_Saturne 3.2.0 (`code_saturne info --guide theory`).
- [12] M. SAKIZ, VALIDATION TEAM,  
*Validation de Code\_Saturne version 1.2 : note de synthèse*,  
EDF Report H-I83-2006-00818-FR, 2006 (in french).
- [13] M. TAGORTI., S. DAL-SECCO, A. DOUCE, N. MÉCHITOUA,  
*Physiques particulières dans Code\_Saturne, tome 4 : le modèle P-1 pour la modélisation des trans-*  
*ferts thermiques radiatifs en milieu gris semi-transparent*,  
EDF Report HI-81/03/017/A, 2003 (in french).

EDF R&D	<b><i>Code_Saturne</i> version 3.2.0 practical user's guide</b>	<i>Code_Saturne</i> documentation Page 195/ <a href="#">202</a>
---------	---	---

- [14] *Code\_Saturne* DOCUMENTATION,  
*Code\_Saturne version 3.2.0 tutorial*, on line with the release of *Code\_Saturne* 3.2.0 (`code_saturne info --guide tutorial`).

## Index of the main variables and keywords

### – Symbols –

iortvm	37	ce1	176
isvhb	38	ce2	176
isvtb	39	ce4	177
ttclag	186	cebu	110
coejou	49	ckabsg	101
dpot	49	ckupdc	40, 88
dtcmom	38	ckwa1	179
icdpar	49	ckwbt1	179
icdtmo	38	ckwbt2	179
icodcl	72	ckwc1	180
idtmom	38	ckwgm1	179
iep	35	ckwgm2	179
ifb	35	ckwsk1	179
ik	35	ckwsk2	179
iomg	35	ckwsw1	179
iphi	35	ckwsw2	179
ipr	34	climgr	162
ir11	35	climgy	169
ir12	35	cmu	176
ir13	35	coefa	39
ir22	35	coefau	39
ir23	35	coefb	39
ir33	35	coefbu	39
iscapp	35	cofafu	39
iscavr	35	cofbfu	39
isca	35	compog	100
itypfb	72	couimp	128, 183
iu	34	coumax	150
iv	34	coumxy	169
iw	35	cp0	174
rcodcl	72	cpgd1	46
		cpgd2	46
		cpght	46
		cppart	187

### – A –

ales	84, 156
almax	176
alpnmk	180
anomax	162
arak	164
atgaze	100
auxl	48

### – B –

betnmk	180
blencv	164
blency	169
bles	84, 156

### – C –

cdgfac	33
cdgfbo	33
cdries	155
cdtvar	150

ce1	176
ce2	176
ce4	177
cebu	110
ckabsg	101
ckupdc	40, 88
ckwa1	179
ckwbt1	179
ckwbt2	179
ckwc1	180
ckwgm1	179
ckwgm2	179
ckwsk1	179
ckwsk2	179
ckwsw1	179
ckwsw2	179
climgr	162
climgy	169
cmu	176
coefa	39
coefau	39
coefb	39
coefbu	39
cofafu	39
cofbfu	39
compog	100
couimp	128, 183
coumax	150
coumxy	169
cp0	174
cpgd1	46
cpgd2	46
cpght	46
cppart	187
crij1	177
crij2	177
crij3	177
crijp1	177
crijp2	177
crit_reca	183
croule	48, 186
csmago	84, 155
csrij	177
cssge2	178
cssgr1	178
cssgr2	178
cssgr3	178
cssgr4	178
cssgr5	178
cssgs1	177
cssgs2	177
cstlog	176

cv2fa1	178
cv2fc1	178
cv2fc2	178
cv2fcl	179
cv2fct	178
cv2fe2	178
cv2fet	179
cv2fmu	176, 178

– D –

diftl0	110, 175
diipb	33
dijpf	33
dispar	40
distb	34
distch	107
divukw	41
dofij	34
dpot	183
dt	41
dtmax	150
dtmin	150
dtp	186, 190, 193
dtpt1d	117
dtref	150

– E –

ehgazg	101
emphis	141
epalim	180
eppt1d	40, 116
epscvy	169
epsdp	165
epsilo	163
epsily	169
epsrgr	161
epsrgy	169
epsrsm	171
epsrsy	169
epsup	170
epszer	172
epzero	171
ettp	45, 186
ettpa	46, 186
extrag	162
extray	169

– F –

ficfpp	139
ficjnf	139
ficush	142
ficusr	144
field	34
fment	106
foumax	150
fs(1)	101

– G –

gamnmk	180
gradpr	46
gradvf	46
grand	171
gx,gy,gz	172

– H –

hbord	38
hept1d	117

– I –

iale	180
ialtyb	132
iangbd	192
ibfixe	132
icalhy	165
icapt	96
iccoal	98
iccvfg	170
icdpar	152, 167
icdtmo	38
icelbr	34
icepdc	40
icepdp	88
icetsm	40, 89
icfgrp	184
icfuel	98
ichrvr	107, 140
ickabs	108
iclkep	153
iclprr	154
iclrtp	39
iclsyr	154
iclt1d	117
iclvfl	148
iclvor	78
icmome	37
icocel	45
icod3p	97
icoebu	98
icoef	39
icoeff	39
icoefr	39
icolwc	98
icompf	99
iconv	148
icour	37
icp	36, 174
icpa	37
icpext	158
icpl3c	98
icpsyr	148
idebty	77
idepo1	121

idepo2 .....	121	ifm .....	103, 108
ideuch .....	152	ifmcel .....	41
idfnom .....	143	ifmfbr .....	39
idiam2 .....	109	ifour .....	37
idiff .....	148	ifp2m .....	103, 109
idiffl .....	189	ifp3m .....	104
idifft .....	149	ifpt1d .....	40
idifre .....	154	ifresf .....	132
idilat .....	149	ifrlag .....	120
idircl .....	149	igfuel .....	100
idirla .....	189	igliss .....	132
idist .....	34	igmdch .....	109
idistu .....	188	igmdv1 .....	109
idiver .....	181	igmdv2 .....	109
idpvar .....	187	igmhet .....	109
idreca .....	183	igoxy .....	100
idries .....	155	igrake .....	153
idstnt .....	189	igrari .....	154
idtvar .....	149	igrdpp .....	185
iecaux .....	48, 145	igrhok .....	153
ielarc .....	98, 182	ih2 .....	104, 109
ielcor .....	183	ihisvr .....	141
ieljou .....	98, 182	ihm .....	103, 108, 109
iencbd .....	192	iilagr .....	185
iencra .....	187	iimlum .....	181
iencrl .....	121	iimpar .....	181
iensi3 .....	191	iindef .....	72
ientat .....	106	iirayo .....	180
ientcp .....	106	ikecou .....	153
ientfu .....	106	ilapoi .....	188
ientgb .....	106	ileaux .....	48, 145
ientgf .....	106	ilisvr .....	107, 144
ientox .....	106	ilogpo .....	153
ientre .....	72, 106	ilvu .....	190
ientrl .....	121	imgr .....	163
ieos .....	184	imgropy .....	169
iescal .....	167	imligr .....	161
iescor .....	37, 166	imligy .....	168
iesder .....	37, 166	immel .....	109
iespre .....	37, 166	imodak .....	181
iestim .....	37, 166	imoold .....	143
iestot .....	37, 167	imoybr .....	193
if1m .....	104, 109	impdvo .....	139
if2m .....	104, 109	impfpp .....	139
if3m .....	104, 109	impjnf .....	139
if3p2m .....	109	impla1 .....	140
if4p2m .....	104	impla2 .....	140
if4pm .....	109	impla3 .....	140
ifabor .....	33	impla4 .....	140
ifacel .....	33	impla5 .....	140
ifapat .....	40	impmvto .....	139
ifinty .....	77	impush .....	142
iflmbd .....	192	impusr .....	144
ifluaa .....	37	impvar .....	187
iflxmw .....	180	impvvo .....	139

imrga	161	isno2t	157
imvisf	170	isolib	72
inbrbd	192	isortl	121
indep	46	isso2t	158
indjon	99	isstpc	164
injcon	186	isstpy	169
inp	104, 109	istala	189
inpd0	145	istat	148
iparoi	72	istmpf	156
iparug	72	isto2t	157
iphydr	165	isttio	186
iphyla	187	isuila	185
iplas	186	isuidr	181
ipnfac	33	isuist	185
ipnfbr	33	isuisy	155
iporos	165	isuit1	170
ippmod	97	isuite	145
ipppro	107	isuivo	155
ipprob	36	isymet	72
ipproc	36, 107	isympa	39
iprco	164	it3m	109
iprfml	39	it4m	109
iprtot	37	itbrrb	148
ipstcl	140	itemp	108
ipstdv	140	itemp1	109
ipstft	140	itemp2	109
ipstyp	140	itepa	46
iptlro	149	itpvar	187
ipucou	170	itrifb	39, 77
iqimp	106	itsnsa	37
irayvf	182	itssca	37
irccor	153	itstua	37
ircflu	170	iturb	151
ircfly	168	iturt	152
irebol	121	itycel	45
irepvo	77	itycor	154
iresol	162	itypfb	39
irevmc	164	itypsm	40, 89
irijec	154	iu	106
irijnu	154	iushlb	121
irjrb	155	iushag	122
iroext	158	iushcl	120
irrom	36	iushvis	123
irrom2	109	iv	106
irroma	36	ivelco	170
iroule	186	iviext	158
irovar	172	ivimpo	132
iscalt	35, 147	ivisch	191
iscavr	147	ivisck	191
ischcv	164	iviscl	36
ischcy	169	ivisct	36
iscthp	156	iviscv	184
iscold	146	ivisdh	191
iscsth	147	ivisdh	191
ismago	37	ivisla	36

ivisls	37, 175	nbmomt	32
ivisma	37	nbmomx	32
ivismp	191	nbpart	186
ivissa	37	nbpmax	45, 186
ivisse	149	nbrvaf	182
ivista	36	nbstru	180
iviste	191	ncapt	141
ivistp	191	ncegrm	163
ivisv1	190	ncel	31
ivisv2	191	ncelbr	31
ivitbd	192	ncelet	31
ivivar	172	ncepdc	40, 88
ivrtex	48, 155	ncepdp	88
ivsext	158	ncesmp	89
iw	106	ncetsm	40, 89
iwarni	144	ncharb	122
iwarny	168	ncharm	98, 122
ix2	109	nclacp	32
ixch	104, 109	nclagm	121
ixck	104, 109	nclcpm	32
ixkabe	101	nclpch	98
iygfm	103, 108	ncpcmx	98
iy(1)	108	ncymax	163
iy(2)	108	ndgmox	32
iy(3)	108	ndim	31
iy1(1)	109	ndirec	181
iy1(2)	109	ndlagm	122
iy1(3)	109	ndlaim	122
iy1(4)	109	nestmx	32, 166
iy1(5)	109	nfabor	31
iy1(6)	109	nfac	31
iy1(7)	110	nflagm	120
izone	106	nfml	31
– J –		nfpt1d	40, 117
jvls	186	nfreqr	181
– K –		nfrlag	120
kabse	100	ngaze	100
– L –		ngazg	100, 101
lndfac	31	ngrmax	163
lndfbr	31	nitmax	163
lndnod	45	nitmay	168
ltsdyn	188	nitmgf	163
ltsmas	188	nivep	45
ltsthe	188	mment	77
– M –		mmod	31
modcpl	189	nodfac	31, 33
modrec	183	nodfbr	31, 33
– N –		nombrd	193
nalimx	180	nomcoe	100
nalinf	180	nomlag	190
nato	100	nomvar	144
		nordre	188
		npo	100, 101
		nppt1d	40, 116
		nprfml	31

nproce .....	32	puisim .....	183
nprofb .....	32	puismp .....	128
npromx .....	32		
npst .....	190	– Q –	
npstf .....	193	qimp .....	106
npstft .....	193	qimpat .....	106
npstt .....	190	qimpcp .....	107
nrgaz .....	100		
nscal .....	31	– R –	
nscamx .....	31	ra .....	41
nscapp .....	32	rcodcl .....	106
nscaus .....	32, 147	rcpt1d .....	117
nstbor .....	192	relaxv .....	151
nstist .....	189	relxst .....	151
nstits .....	188	rgpt1d .....	116
nswrgr .....	161	rinfm .....	171
nswrgy .....	168	rlxp1 .....	164
nswrsm .....	171	ro0 .....	173
nswrsy .....	168	rr .....	172
ntcabs .....	145	rtp .....	34, 103
ntcmxy .....	168	rtpa .....	34
ntdmom .....	143	ruslag .....	122
ntersl .....	45	rvarfl .....	176
nterup .....	170		
nthist .....	142	– S –	
nthsav .....	142	s2kw .....	41
ntlist .....	144	scamax .....	175
ntmabs .....	146	scamin .....	175
ntpabs .....	146	seuil .....	189
ntsuit .....	145	seuifl .....	192
ntypmx .....	74	sigmae .....	177
nusbor .....	45, 192	sigmak .....	177
nushmx .....	32	sigmas .....	175
nvar .....	31	smacel .....	40, 89
nvep .....	45	smagmx .....	155
nvgaus .....	45	srrom .....	110, 173
nvisbr .....	45	statis .....	47, 190
nvisls .....	32	stephn .....	172
nvls .....	45, 186	stoeg .....	100
nvlst .....	45	surf .....	33
nvlst .....	45, 190	surfan .....	34
nvort .....	77	surfbn .....	34
nvp .....	45	surfbo .....	33
– P –		– T –	
p0 .....	173	t0 .....	174
parbor .....	47, 191, 192	tbord .....	39
permvi .....	172	tepa .....	46
pi .....	171	tept1d .....	117
pond .....	34	th .....	101
porosi .....	40	thetav .....	159
pred0 .....	173	thetcp .....	160
prefth .....	172	thetfl .....	159
propce .....	34	thetro .....	160
propfb .....	34	thetsn .....	159
		thetss .....	160



thetst .....	159	xyzcen .....	33
thetvi .....	160	xyznod .....	33
thetvs .....	160	xyzp0 .....	174
timpat .....	106		
timpcp .....	107	– Y –	
tinmue .....	106	yplmxy .....	169
tinoxy .....	106	yplpar .....	40
tkelvi .....	172	ypluli .....	153
tkelvn .....	172		
tkent .....	106	– Z –	
tmarus .....	146	zero .....	171
tmax .....	100, 182		
tmin .....	100, 182		
tpart .....	187		
tppt1d .....	116		
tprenc .....	122, 187		
treftth .....	172		
tslagr .....	47		
tstat .....	124, 190		
tstatp .....	193		
ttcabs .....	146		
ttpabs .....	146		
– U –			
uetbor .....	39		
uref .....	176		
– V –			
vagaus .....	48		
varrdt .....	150		
viscl0 .....	173		
viscv0 .....	185		
visls0 .....	175		
visref .....	122, 187		
vitflu .....	46		
vitpar .....	46		
volmol .....	172		
volume .....	33		
– W –			
wmolat .....	100		
wmolg .....	101		
– X –			
xco2 .....	101		
xh2o .....	101		
xkabe .....	101		
xkabel .....	101		
xkappa .....	176		
xlesfd .....	156		
xlesfl .....	84, 156		
xlmt1d .....	117		
xlomlg .....	176		
xmasmr .....	184		
xnp1mx .....	181		
xyzcap .....	141		